Theory, verification and validation of finite- β gyrokinetics

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Structure of this Presentation

Role of electromagnetic effects in given contexts

- 1. Overview of theory and defintions
 - Gyrokinetic equations, transport coefficients, general theory
- 2. Description of common toroidal eigenmodes
 - Parameteric dependence and eigenmode structure
- 3. Verification and validation
 - Examples from DIII-D and NSTX
- 4. Perplexing features of electromagnetic simulation
 - Transport runaway and magnetic stochasticity



Electromagnetic Gyrokinetic Equations

Field dependence through quantity Ψ_a

$$\frac{\partial h_a}{\partial t} + \frac{v_{\parallel}}{\mathcal{J}_{\psi}B} \frac{\partial H_a}{\partial \theta} + \mathbf{v}_d \cdot \nabla H_a + \omega_0 \frac{\partial h_a}{\partial \alpha} + c \left[h_a, \Psi_a\right]_{\psi,\alpha} \\ + c \left(\frac{\partial f_{a0}}{\partial \psi} + \frac{m_a v_{\parallel}}{T_a} \frac{I}{B} \frac{\partial \omega_0}{\partial \psi} f_{a0}\right) \frac{\partial \Psi_a}{\partial \alpha} = C_a^{GL} \left[H_a\right] . \\ H_a(\mathbf{R}) = \frac{e_a f_{a0}}{T_a} \Psi_a(\mathbf{R}) + h_a(\mathbf{R})$$

 $\mathbf{R} \rightarrow$ guiding-center position

 $H_a \rightarrow$ nonadiabatic distribution of species a

 $f_{a0} \rightarrow$ Maxwellian equilibrium (in frame of rotation) of species a



Gyroaverages of Fields

Compressional potential $\delta B_{||}$ requires different average

$$\begin{split} \Psi_{a}(\mathbf{R}) &\doteq \left\langle \delta\phi(\mathbf{R}+\boldsymbol{\rho}) - \frac{1}{c}(\mathbf{V}_{0}+\mathbf{v}) \cdot \delta\mathbf{A}(\mathbf{R}+\boldsymbol{\rho}) \right\rangle_{\mathbf{R}}, \\ &= \mathcal{G}_{0a}\left[\delta\phi(\mathbf{R}) - \frac{v_{\parallel}}{c} \delta A_{\parallel}(\mathbf{R}) \right] + \frac{v_{\perp}^{2}}{\Omega_{ca}c} \mathcal{G}_{1a} \, \delta B_{\parallel}(\mathbf{R}) \, . \end{split}$$

 $ho = \mathbf{b} \times \mathbf{v}' / \Omega_{ca} \rightarrow \text{gyroradius vector}$ $\Omega_{ca} = e_a B / (m_a c) \rightarrow \text{cyclotron frequency}$ $\delta \phi \rightarrow \text{electrostatic potential}$ $\delta A_{||} \rightarrow \text{transverse electromagnetic potential}$ $\delta B_{||} \rightarrow \text{compressional electromagnetic potential}$



Gyroaverages of Fields

Pseudospectral operators valid for all wavelengths

$$\begin{split} \Psi_{a}(\mathbf{R}) &= \mathcal{G}_{0a} \left[\delta \phi(\mathbf{R}) - \frac{v_{\parallel}}{c} \delta A_{\parallel}(\mathbf{R}) \right] + \frac{v_{\perp}^{2}}{\Omega_{ca} c} \mathcal{G}_{1a} \, \delta B_{\parallel}(\mathbf{R}) \ . \\ z(\mathbf{R}) &\doteq \sum_{\mathbf{k}_{\perp}} e^{iS(\mathbf{R})} \, \tilde{z}(\mathbf{k}_{\perp}) \ , \end{split}$$

 \mathcal{G}_{0a} and \mathcal{G}_{1a} are **pseudospectral** operators in real space, with Bessel function representations in wavenumber space:

$$abla_{\perp}^2
ightarrow - k_{\perp}^2 ,$$
 $\mathcal{G}_{0a}
ightarrow J_0(k_{\perp}\rho_a) ,$
 $\mathcal{G}_{1a}
ightarrow rac{1}{2} \left[J_0(k_{\perp}\rho_a) + J_2(k_{\perp}\rho_a)
ight] .$



Electromagnetic Maxwell Equations $N_FIELD = 1, 2, 3$

Poisson equation

$$-\nabla_{\perp}^2 \delta \phi(\mathbf{x}) = 4\pi \sum_a e z_a \, \delta n_a = 4\pi \sum_a e_a \int d^3 v \, \hat{f}_{a1}(\mathbf{x}) \; .$$

Parallel Ampère's Law

$$-\nabla_{\perp}^2 \delta A_{\parallel}(\mathbf{x}) = \frac{4\pi}{c} \sum_a \delta j_{\parallel,a} = \frac{4\pi}{c} \sum_a e_a \int d^3 v \, v_{\parallel} \, \hat{f}_{a1}(\mathbf{x})$$

Perpendicular Ampère's Law

$$\nabla_{\perp} \delta B_{\parallel}(\mathbf{x}) \times \mathbf{b} = \frac{4\pi}{c} \sum_{a} \delta \mathbf{j}_{\perp,a} = \frac{4\pi}{c} \sum_{a} e_{a} \int d^{3}v \, \mathbf{v}_{\perp} \hat{f}_{a1}(\mathbf{x})$$



Overview and General Considerations

Connecting particle distribution to gyrocenter distribution

Right-hand sides can be written in terms of H_a

$$\int d^3 v \, \hat{f}_{a1}(\mathbf{x}) = -\frac{n_a e_a}{T_a} \, \delta \phi(\mathbf{x}) + \int d^3 v \, H_a(\mathbf{x} - \boldsymbol{\rho}) \,,$$
$$\int d^3 v \, v_{\parallel} \, \hat{f}_{a1}(\mathbf{x}) = \int d^3 v \, v_{\parallel} \, H_a(\mathbf{x} - \boldsymbol{\rho}) \,,$$
$$\int d^3 v \, \mathbf{v}_{\perp} \, \hat{f}_{a1}(\mathbf{x}) = \int d^3 v \, \mathbf{v}_{\perp} \, H_a(\mathbf{x} - \boldsymbol{\rho}) \,,$$

 $\hat{f}_{a1}(\mathbf{x}) o$ fluctuating part of perturbed 6-D distribution $\mathbf{x} = \mathbf{R} + oldsymbol{
ho}$ oparticle position



Flux surfaces labeled by effective minor radius, r.

Generalises the Waltz-Miller midplane minor radius

r is the **half-width** of the flux-surface at the elevation of the centroid.





The effective field B_{unit} and other flux functions

Meaning is perpetual source of confusion for users

• B_{unit} is the effective magnetic field.

$$B_{\text{unit}}(r) \doteq \frac{1}{r} \frac{d\chi_t}{dr} = \frac{q}{r} \frac{d\psi}{dr}$$

- Arguably the most elegant choice for local simulations
- Effective gyroradius

$$\rho_{s,\text{unit}} = \frac{c_s}{eB_{\text{unit}}/(m_i c)}$$

Effective electron beta

$$\beta_{e,\text{unit}} = \frac{8\pi n_e T_e}{B_{\text{unit}}^2}$$



Transport Coefficients

Suitable ensemble averages, $\langle\!\langle \cdot
angle\!
angle$, must be taken

$$\begin{split} \Gamma_{a} &= \frac{c}{\psi'} \left\langle \!\! \left\langle \int d^{3} v \, H_{a}^{*}(\mathbf{R}) \, \frac{\partial \Psi_{a}}{\partial \alpha} \right\rangle \!\! \right\rangle, \\ Q_{a} &= \frac{c}{\psi'} \left\langle \!\! \left\langle \int d^{3} v \, H_{a}^{*}(\mathbf{R}) \, \frac{1}{2} m_{a} v^{2} \, \frac{\partial \Psi_{a}}{\partial \alpha} \right\rangle \!\! \right\rangle, \\ \Pi_{a} &= \frac{c}{\psi'} \left\langle \!\! \left\langle \int d^{3} v \, H_{a}^{*}(\mathbf{R}) m_{a} R \left[\left(V_{0} + v_{\parallel} \frac{B_{t}}{B} \right) \frac{\partial \Psi_{a}}{\partial \alpha} + v_{\perp} \frac{B_{p}}{B} \frac{\partial \mathcal{X}_{a}}{\partial \alpha} \right] \right\rangle \!\! \right\rangle, \\ S_{a} &= \frac{c}{\psi'} \left\langle \!\! \left\langle \int d^{3} v \, H_{a}^{*}(\mathbf{R}) \, e_{a} \left(\frac{\partial}{\partial t} + \omega_{0} \frac{\partial}{\partial \alpha} \right) \Psi_{a} \right\rangle \!\! \right\rangle. \\ &\left\langle \!\! \left\langle \cdot \right\rangle \!\! \right\rangle \doteq \lim_{t_{*} \to \infty} \frac{1}{2\pi L \tau} \int_{o}^{L} dr \int_{0}^{2\pi} d\alpha \int_{0}^{\tau} dt \, \mathcal{F} \cdot \,, \end{split}$$



Transport Coefficients

GyroBohm normalizations

$$\Gamma_a \to \Gamma_{\rm GB} \doteq n_e c_s (\rho_{s,\rm unit}/a)^2$$
$$Q_a \to Q_{\rm GB} \doteq n_e c_s T_e (\rho_{s,\rm unit}/a)^2$$
$$\Pi_a \to \Pi_{\rm GB} \doteq n_e a T_e (\rho_{s,\rm unit}/a)^2$$
$$S_a \to S_{\rm GB} \doteq n_e (c_s/a) T_e (\rho_{s,\rm unit}/a)^2$$



Form of Maxwell Equations used in Practice

Poisson equation

$$-\frac{1}{4\pi}\nabla_{\perp}^{2}\delta\phi + \sum_{a}n_{a}\frac{e_{a}^{2}}{T_{a}}\int d^{3}v F_{Ma}\left(1-\mathcal{G}_{0a}^{2}\right)\delta\phi$$
$$-\sum_{a}n_{a}\frac{e_{a}^{2}}{T_{a}}\int d^{3}v F_{Ma}\mathcal{G}_{0a}\mathcal{G}_{1a}\frac{v_{\perp}^{2}}{\Omega_{ca}c}\,\delta B_{\parallel} = \sum_{a}e_{a}\int d^{3}v \,\mathcal{G}_{0a}h_{a}$$

Parallel Ampère's Law

$$-\frac{1}{4\pi}\nabla_{\perp}^{2}\delta A_{\parallel} + \sum_{a} n_{a}\frac{e_{a}^{2}}{T_{a}}\int d^{3}v \,\frac{v_{\parallel}^{2}}{c^{2}}F_{Ma}\mathcal{G}_{0a}^{2}\delta A_{\parallel} = \sum_{a} e_{a}\int d^{3}v \,\frac{v_{\parallel}}{c}\mathcal{G}_{0a}h_{a}$$

Perpendicular Ampère's Law

$$\begin{split} \frac{1}{4\pi} \delta B_{\parallel} + \sum_{a} n_{a} \frac{e_{a}^{2}}{T_{a}} \int d^{3}v F_{Ma} \left(\frac{v_{\perp}^{2}}{\Omega_{ca}c} \mathcal{G}_{1a} \right)^{2} \delta B_{\parallel} \\ + \sum_{a} n_{a} \frac{e_{a}^{2}}{T_{a}} \int d^{3}v F_{Ma} \frac{v_{\perp}^{2}}{\Omega_{ca}c} \mathcal{G}_{1a} \mathcal{G}_{0a} \delta \phi &= -\sum_{a} e_{a} \int d^{3}v \mathcal{G}_{1a} \frac{v_{\perp}^{2}}{\Omega_{ca}c} h_{a} \end{split}$$



The Ampère Cancellation Problem

• Let's assume a pure plasma with $T_i = T_e$ and $k_{\perp} \rho_{s, {\rm unit}} \ll 1$:

$$-\frac{2k_{\perp}^2\rho_{s,\text{unit}}^2}{\beta_{e,\text{unit}}}\delta\hat{A}_{\parallel} + \left(1 + \frac{m_i}{m_e}\right)\delta\hat{A}_{\parallel} = \sum_a e_a \int d^3v \,\hat{v}_{\parallel a}\hat{h}_a$$

- The factor m_i/m_e is artificial.
- It is cancelled by a corresponding term in \hat{h}_a .
- Attempting to perform field integral analytically will lead to pain.
- Must devise a numerical scheme for which artificial pieces cancel.



Finite- β effects appear in two different ways/places:

1. Magnetic fluctuations: $\beta_{e,\text{unit}} \rightarrow \text{AMPERE}_\text{SCALE} \times \beta_{e,\text{unit}}$

$$-\frac{2\rho_{s,\text{unit}}^2}{\beta_{e,\text{unit}}}\nabla_{\perp}^2\delta\hat{A}_{\parallel} + \sum_a \alpha_a z_a^2 V[\hat{v}_{\parallel a}^2 \mathcal{G}_{0a}^2 \delta\hat{A}_{\parallel}] = \sum_a z_a V[\hat{v}_{\parallel a} \mathcal{G}_{0a} \hat{h}_a]$$

2. Geometry/drift motion: $\nabla p \to \texttt{GEO_BETAPRIME_SCALE} \times \nabla p$

$$\mathbf{v}_{d} = \frac{v_{\parallel}^{2} + \mu B}{\Omega_{ca} B} \,\mathbf{b} \times \nabla B + \frac{2v_{\parallel}\omega_{0}}{\Omega_{ca}} \,\mathbf{b} \times \mathbf{s} + \frac{4\pi v_{\parallel}^{2}}{\Omega_{ca} B^{2}} \,\mathbf{b} \times \nabla p$$



Workflow enabled with profiles_gen command-line tool

\$ profiles_gen -i iterdb

Supported interfaces for reading profile data from:

- 1. ITERDB ASCII
- 2. ITERDB NetCDF
- 3. Plasma State NetCDF
- 4. CORSICA ASCII
- 5. ASTRA ASCII
- 6. PEQDSK/ELITE ASCII
- 7. UFILE (ITPA database)

Options for profiles_gen usage: EFIT gfile

- GATO mapper automatically extracts high-resolution flux surfaces
- Fitter simultaneously generates model (Miller-type) fit

$$R(r,\theta) = R_0(r) + r\cos(\theta + \arcsin\delta\sin\theta) ,$$

$$Z(r,\theta) = Z_0(r) + \kappa r\sin(\theta + \zeta\sin 2\theta) ,$$

• And general (up-down asymmetric) fit

$$\begin{aligned} R(r,\theta) &= \frac{1}{2} a_0^R(r) + \sum_{n=1}^N \left[a_n^R(r) \cos(n\theta) + b_n^R(r) \sin(n\theta) \right] , \\ Z(r,\theta) &= \frac{1}{2} a_0^Z(r) + \sum_{n=1}^N \left[a_n^Z(r) \cos(n\theta) + b_n^Z(r) \sin(n\theta) \right] . \end{aligned}$$



General fit only required close to separatrix (r/a = 0.99 shown below)



Options for $profiles_gen$ usage: NEO calculation of E_r

• Use **NEO** to compute rotation profile

$$\omega_0(r) = \frac{cE_r}{RB_p} = -c\frac{\partial\Phi}{\partial\psi}$$

- Calculation unique given measured Carbon $v_{\phi,c}(r)$ profile.
- Logic for handling multiple ions, species lumping, sonic rotation, etc.
- **Diagnostic** calculation of all ion velocities: $v_{\phi,a}$, $v_{\theta,a}$.

GA Standard Case (Miller circle) β **scan** $\alpha_{\rm MHD} = 0$

Ion Temperature Gradient (ITG) Mode $\beta_e = 0.2\%, \alpha_{MHD} = 0$

Trapped Electron Mode (TEM) $\beta_e = 0.2\%, \alpha_{MHD} = 0$

Kinetic Ballooning Mode (KBM) $\beta_e = 2.2\%, \alpha_{MHD} = 0$

Tearing-parity mode (TPM) $\beta_e = 2.2\%, \alpha_{MHD} = 0$

GA Standard Case β scan

Self-consistent $lpha_{ m MHD}$

Ion temperature gradient scan

Electron temperature gradient scan

Density gradient scan

Collision frequency scan

SENERAL ATOMICS

Finite- β version of the Cyclone Case Belli POP 2010

MHD critical beta occurs at about $\beta_{e,\text{unit}} \simeq 1.2\%$

Excellent GYRO-GS2 agreement on Holland validation case Bravenec PoP 2011: Revisit DIII-D 128913 at $\rho = 0.5$

Excellent GYRO-GS2 agreement on Holland validation case

Bravenec PoP 2011: Good agreement in all channels

Effect of α_{MHD} retained but δA_{\parallel} ignored.

DIII-D High- β plasmas

Holland PoP 2012

- Most validation studies have focused on low-power L-mode discharges
- Key difference: $\rho_{s,\text{unit}}/a$ larger in H-mode than L-mode
- Profile shearing effects can contribute some stabilization
- Focus on discharges created for study of transport scaling with eta
- Transport in 128385 quenched at full β .

DIII-D High- β plasmas Holland PoP 2012

Experimental results bracketed by $0.5 < \beta_e/\beta_e^{\rm expt} < 1$

💠 GENERAL ATOMICS

Electromagnetic Cyclone Eigenmodes including δB_{\parallel} Belli POP 2010

- Sea of modes in NSTX made initial-value linear simulations problematic
- Near mode crossings, eigenmode fails to emerge clearly
- Impossible to generate smooth curves of frequency versus parameter.
- Creation of GYRO field **eigenvalue solver** was motivated
- This is in contrast to Bass' more comprehensive distribution eigenvalue solver

Electromagnetic Cyclone Eigenmodes including δB_{\parallel} Belli POP 2010

💠 GENERAL ATOMICS

NSTX Eigenmodes at $k_{ heta} ho_s = 0.25$ Belli POP 2010

(a-c): KBM, (d-i): Hybrid ITG/KBM

NSTX Eigenmodes at $k_{\theta} \rho_s = 0.6$ Belli POP 2010

(a-c): KBM, (d-f): ITG-like, (g-i): Hybrid ITG/KBM

NSTX Eigenmodes at $k_{\theta} \rho_s = 15$ Belli POP 2010

Alfvénic drift eigenfunctions

NSTX Eigenmodes at $k_{\theta} \rho_s = 15$ Belli POP 2010

Compressional electron drift waves

Significant advances and innovations by Guttenfelder

- Wide range of parameters
- H-mode Q_i often near neoclassical levels
- Treat core region, $0.4 \leq r/a \leq 0.8$ with GYRO
- Electrostatic ITG/TEM found at lower β
- ETG found above $a/L_{T_e, crit}$.
- Microtearing at high β_e
 - $\chi_{e,\rm EM} \simeq 6m^2/s$
 - $\Delta x \leq 0.2 \rho_{s,\text{unit}}$
 - Transport increases with ν_{ei} .

129041: KBM unstable at high $\alpha_{\rm MHD} \sim \beta'$.

Phenomenology of the TEM/KBM branch

TEM: Destabilized by a/L_{T_e} , a/L_n , weakly dependent on a/L_{T_i} , stabilized by ν_{ei} KBM: Growth rate scaling unified by $\alpha_{\rm MHD} = -q^2 R \beta'$

Large contribution from compressional transport channel: $Q_e^{\delta B_{\parallel}}$

Nearly half of Q_e from compressional motion:

 $\frac{\delta B_{\parallel}}{B_{\rm unit}}$

 $\simeq 0.08\%$.

Local linear AE modes

Bass PoP 2010: Simultaneous EPM, TAE, ITG, TEM

Local nonlinear AE simulations (half-torus)

Bass PoP 2010: Saturated nonlinear states at lower EP fraction

Global linear AE modes (eigenvalue solver)

Bass In Press 2012: Three simultaneous modes (DIII-D 142111)

AE Simulation Challenges

Kinetic Energetic Particles

- EP orbits:
 - resolving orbit motion requires smaller timestep (factor of 10)
 - large orbits require wider gyroaverage stencil
- Near-marginality of Alfvénic modes requires long simulation times
- Multi-scale coupling requires simultaneous resolution of
 - low-k Alfvén (large domain) dynamics
 - intermediate-k ITG/TEM (fine-scale) turbulence
- Global linear analysis
 - Gyrokinetic eigensolver solves $1.15M \times 1.15M$ eigensystem!

Original β -scaling paper and the runaway Candy POP 2005

- Original β scans showed something strange happening
- Simulations ran away at about $\beta = (2/3)\beta_{crit}$.
- Did NOT appear to be a numerical instability.

Original β -scaling paper and the runaway Candy POP 2005

- Runaway motivated the development of the IMEX-RK semi-implicit method.
 - Wasn't a miracle cure
 - SIDE-BENEFIT: linear simulations with real m_i/m_e much faster
- Cancellation issue eventually ruled out as culprit

Original β -scaling paper and corrugations Candy POP 2005

• Significant radial structure about lowest-order rational surfaces

- Related to full (non-fluid) kinetic electron response
- Physical pole-like structure in electron propagator

Electromagnetic Fluctuations and Magnetic Stochasticity

- Connection between runaway and stochasticity was suggested ca. 2006
- Quantification required magnetic-field-line mapping capability
- Tedious because of ballooning representation
- Poincaré mapper part of GYRO: gacode/gyro/tools/fieldline
- There were various diversions related to runaway "cures"
 - better numerical methods
 - more physical realism (collisions)
 - higher resolution (electron-scale grid)
- Runaway is **correct solution** of model equations

Electromagnetic Fluctuations and Magnetic Stochasticity Wang POP 2011, PRL 2011

• GYRO, GENE and GKW were ultimately in agreement about the runaway

Magnetic Stochasticity Wang POP 2011, PRL 2011

- Remarkable discovery 1: Stochasticity observed at smallest values of β
- Chaos in this case is not "simple"; appear to be bounding tori.

Magnetic Stochasticity Wang POP 2011, PRL 2011

- Remarkable discovery 2: EM electron transport is almost purely chaotic
- Correlation in time suprisingly high

Magnetic Stochasticity Wang POP 2011, PRL 2011

• Stochastic energy flux, Q_{st} , is stochastic particle flux, d_m , times tentative conversion factor:

$$Q_{\rm st} = \sqrt{\frac{8}{\pi}} d_m \frac{v_{\rm th}}{L_T} n_{\rm pass} T$$

• Bursts in NSTX (Guttenfelder) not understood

Subcritical MHD β -limit

Waltz POP 2010

- Total pressure profile **corrugated** at finite transport levels
- Some evidence that regions of larger p' lower the effective β limit

