

# Predictive Gyrokinetic Transport Simulations and Application of Synthetic Diagnostics

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# Preview: The TGYRO Project

Simulate multiple space and time scales using first-principles physics

**TGYRO provides high-level control of GYRO, NEO, TGLF, etc.**

- ▷ Complete hierarchical (ordered) solution of **Fokker-Planck** equations.
- ▷ Proper analytic **equilibrium** solution for **rotating** plasmas.
- ▷ **Neoclassical transport**: direct spectral solution of the drift-kinetic equation.  
**NEO code (Belli)** or **NEO formula module**
- ▷ **Turbulent transport**: direct Eulerian solution of nonlinear gyrokinetic equations.  
**GYRO code** or **TGLF model (Staebler)**
- ▷ Fully-implicit solution of transport equations in **integral form**.
- ▷ **Rapid convergence** for quasilinear transport models (TGLF, IFS-PPPL).
- ▷ Handy tool for transport model **comparison** or **development**.

# Outline of this Talk

1. Review of **Fokker-Planck theory** from start to finish.
2. Summary of **TGYRO**: numerical approach and convergence tests.
3. Overview of **synthetic diagnostic** tools.
4. Sample **application of TGYRO** to DIII-D discharge.

# Fokker-Planck Theory of Plasma Transport

Comprehensive series of papers by Sugama and coworkers

The Fokker-Planck (FP) equation provides the **fundamental theory** for **plasma equilibrium, fluctuations, and transport**:

$$\left[ \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla + \frac{e_a}{m_a} \left( (\mathbf{E} + \hat{\mathbf{E}}) + \frac{\mathbf{v}}{c} \times (\mathbf{B} + \hat{\mathbf{B}}) \right) \cdot \frac{\partial}{\partial \mathbf{v}} \right] (f_a + \hat{f}_a) = C_a(f_a + \hat{f}_a) + S_a$$

$f_a$   $\longrightarrow$  **ensemble-averaged distribution**

$\hat{f}_a$   $\longrightarrow$  **fluctuating distribution**

$S_a$   $\longrightarrow$  **sources (beams, RF, etc)**

$$C_a = \sum_b C_{ab}(f_a + \hat{f}_a, f_b + \hat{f}_b) \longrightarrow \text{nonlinear collision operator}$$

# Fokker-Planck theory

Comprehensive, consistent framework for equilibrium profile evolution

The general approach is to separate the FP equation into **ensemble-averaged**,  $\mathcal{A}$ , and **fluctuating**,  $\mathcal{F}$ , components:

$$\mathcal{A} = \left. \frac{d}{dt} \right|_{\text{ens}} f_a - \langle C_a \rangle_{\text{ens}} - D_a - S_a ,$$

$$\mathcal{F} = \left. \frac{d}{dt} \right|_{\text{ens}} \hat{f}_a + \frac{e_a}{m_a} \left( \hat{\mathbf{E}} + \frac{\mathbf{v}}{c} \times \hat{\mathbf{B}} \right) \cdot \frac{\partial}{\partial \mathbf{v}} (f_a + \hat{f}_a) - C_a + \langle C_a \rangle_{\text{ens}} + D_a ,$$

where

$$\left. \frac{d}{dt} \right|_{\text{ens}} \doteq \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla + \frac{e_a}{m_a} \left( \mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right) \cdot \frac{\partial}{\partial \mathbf{v}} ,$$

$$D_a \doteq - \frac{e_a}{m_a} \left\langle \left( \hat{\mathbf{E}} + \frac{\mathbf{v}}{c} \times \hat{\mathbf{B}} \right) \cdot \frac{\partial \hat{f}_a}{\partial \mathbf{v}} \right\rangle_{\text{ens}} .$$

▷  $D_a$  is the **fluctuation-particle interaction operator**.

# Fokker-Planck theory

Space- and time-scale expansion in powers of  $\rho_* = \rho_s/a$

**Ensemble averages** are expanded in powers of  $\rho_*$  as

$$\begin{aligned}f_a &= f_{a0} + f_{a1} + f_{a2} + \dots , \\S_a &= S_{a2} + \dots \text{ (transport ordering),} \\ \mathbf{E} &= \mathbf{E}_0 + \mathbf{E}_1 + \mathbf{E}_2 + \dots , \\ \mathbf{B} &= \mathbf{B}_0 .\end{aligned}$$

**Fluctuations** are also expanded in powers of  $\rho_*$  as

$$\begin{aligned}\hat{f}_a &= \hat{f}_{a1} + \hat{f}_{a2} + \dots , \\ \hat{\mathbf{E}} &= \hat{\mathbf{E}}_1 + \hat{\mathbf{E}}_2 + \dots , \\ \hat{\mathbf{B}} &= \hat{\mathbf{B}}_1 + \hat{\mathbf{B}}_2 + \dots .\end{aligned}$$

Built-in assumption about scale separation **hard to escape**.

# Fokker-Planck theory

## Lowest-order conditions for flow and gyroangle independence

### Lowest-order Constraints

The lowest-order ensemble-averaged equation gives the **constraints**

$$\mathcal{A}_{-1} = 0 : \quad \mathbf{E}_0 + \frac{1}{c} \mathbf{V}_0 \times \mathbf{B} = 0 \quad \text{and} \quad \frac{\partial f_{a0}}{\partial \xi} = 0$$

where  $\xi$  is the gyroangle.

### Large mean flow

The only **equilibrium flow** that persists on the fluctuation timescale is

$$\mathbf{V}_0 = R \omega_0(\psi) \mathbf{e}_\varphi \quad \text{where} \quad \omega_0 \doteq -c \frac{\partial \Phi_0}{\partial \psi} .$$

[F.L. Hinton and S.K. Wong, Phys. Fluids **28** (1985) 3082].

# Fokker-Planck theory

Equilibrium equation is a formidable nonlinear PDE

## Equilibrium equation

The gyrophase average of the zeroth order ensemble-averaged equation gives the **collisional equilibrium** equation:

$$\int_0^{2\pi} \frac{d\xi}{2\pi} \mathcal{A}_0 = 0 : \quad \left( \mathbf{V}_0 + v'_{\parallel} \mathbf{b} \right) \cdot \nabla f_{a0} = C_a(f_{a0})$$

where  $\mathbf{v}' = \mathbf{v} - \mathbf{V}_0$  is the velocity in the rotating frame.

## Equilibrium distribution function

The exact solution for  $f_{a0}$  is a **Maxwellian in the rotating frame**, such that the centrifugal force causes the density to vary on the flux surface:

$$f_{a0} = n_a(\psi, \theta) \left( \frac{m_a}{2\pi T_a} \right)^{3/2} e^{-m_a(v')^2/2T_a} .$$



# Fokker-Planck theory

## Equations for neoclassical transport and turbulence at $\mathcal{O}(\rho_*)$

### Drift-kinetic equation

Gyroaverage of first-order  $\mathcal{A}_1$  gives expressions for **gyroangle-dependent** ( $\tilde{f}_{a1}$ ) and **gyroangle-independent** ( $\bar{f}_{a1}$ ) distributions:

$$\int_0^{2\pi} \frac{d\xi}{2\pi} \mathcal{A}_1 = 0 : \quad f_{a1} = \tilde{f}_{a1} + \bar{f}_{a1}, \quad \tilde{f}_{a1} = \frac{1}{\Omega_a} \int^\xi d\xi \widetilde{\mathcal{L}f_{a0}}$$

▷ Ensemble-averaged  $\bar{f}_{a1}$  is determined by the **drift kinetic equation (NEO)**.

### Gyrokinetic equation

Gyroaverage of first-order  $\mathcal{F}_1$  gives an expression for **first-order fluctuating distribution** ( $\hat{f}_{a1}$ ) in terms of the distribution of the gyrocenters,  $h_a(\mathbf{R})$ :

$$\int_0^{2\pi} \frac{d\xi}{2\pi} \mathcal{F}_1 = 0 : \quad \hat{f}_{a1}(\mathbf{x}) = -\frac{e_a \hat{\phi}(\mathbf{x})}{T_a} + h_a(\mathbf{x} - \rho)$$

▷ Fluctuating  $\hat{f}_{a1}$  is determined by the **gyrokinetic equation (GYRO)**.

# Drift-Kinetic Equation for Neoclassical Transport

NEO gives complete solution with full kinetic e-i-impurity coupling

$$v'_{\parallel} \mathbf{b} \cdot \nabla \bar{g}_a - C_a^L(\bar{g}_a) = \frac{f_{a0}}{T_a} \left[ -\frac{1}{N_a} \frac{\partial N_a T_a}{\partial \psi} W_{a1} - \frac{\partial T_a}{\partial \psi} W_{a2} + c \frac{\partial^2 \Phi_0}{\partial \psi^2} W_{aV} + \frac{\langle B E_{\parallel}^A \rangle}{\langle B^2 \rangle^{1/2}} W_{aE} \right]$$

$$\bar{g}_a \doteq \bar{f}_{a1} - f_{a0} \frac{e_a}{T_a} \int^{\ell} \frac{dl}{B} \left( B E_{\parallel} - \frac{B^2}{\langle B^2 \rangle} \langle B E_{\parallel} \rangle \right),$$

$$W_{a1} \doteq \frac{m_a c}{e_a} v'_{\parallel} \mathbf{b} \cdot \nabla \left( \omega_0 R + \frac{I}{B} v'_{\parallel} \right),$$

$$W_{a2} \doteq W_{a1} \left( \frac{\varepsilon}{T_a} - \frac{5}{2} \right),$$

$$W_{aV} \doteq \frac{m_a c}{2e_a} v'_{\parallel} \mathbf{b} \cdot \nabla \left[ m_a \left( \omega_0 R + \frac{I}{B} v'_{\parallel} \right)^2 + \mu \frac{R^2 B_p^2}{B} \right],$$

$$W_{aE} \doteq \frac{e_a v'_{\parallel} B}{\langle B \rangle^{1/2}}.$$

# Gyro-Kinetic Equation for Turbulent Transport

GYRO gives complete solution with full  $(\phi, A_{\parallel}, B_{\parallel})$  electromagnetic physics.

$$\begin{aligned} \frac{\partial h_a(\mathbf{R})}{\partial t} + \left( \mathbf{V}_0 + v'_{\parallel} \mathbf{b} + \mathbf{v}_{da} - \frac{c}{B} \nabla \hat{\Psi}_a \times \mathbf{b} \right) \cdot \nabla h_a(\mathbf{R}) - C_a^{GL} \left( \hat{f}_{a1} \right) \\ = f_{a0} \left[ -\frac{\partial \ln(N_a T_a)}{\partial \psi} \hat{W}_{a1} - \frac{\partial \ln T_a}{\partial \psi} \hat{W}_{a2} + \frac{c}{T_a} \frac{\partial^2 \Phi_0}{\partial \psi^2} \hat{W}_{aV} + \frac{1}{T_a} \hat{W}_{aT} \right] \end{aligned}$$

$$\hat{W}_{a1}(\mathbf{R}) \doteq -\frac{c}{B} \nabla \hat{\Psi}_a \times \mathbf{b} \cdot \nabla \psi ,$$

$$\hat{W}_{a2}(\mathbf{R}) \doteq \hat{W}_{a1} \left( \frac{\varepsilon}{T_a} - \frac{5}{2} \right) ,$$

$$\hat{W}_{aV}(\mathbf{R}) \doteq -\frac{m_a R c}{B} \left\langle (\mathbf{V}_0 + \mathbf{v}') \cdot \mathbf{e}_{\varphi} \nabla \left( \hat{\phi} - \frac{1}{c} (\mathbf{V}_0 + \mathbf{v}') \cdot \hat{\mathbf{A}} \right) \times \mathbf{b} \cdot \nabla \psi \right\rangle_{\xi} ,$$

$$\hat{W}_{aT}(\mathbf{R}) \doteq e_a \left\langle \left( \frac{\partial}{\partial t} + \mathbf{V}_0 \cdot \nabla \right) \left( \hat{\phi} - \frac{1}{c} (\mathbf{V}_0 + \mathbf{v}') \cdot \hat{\mathbf{A}} \right) \right\rangle_{\xi} .$$

$$\begin{aligned} \hat{\Psi}_a(\mathbf{R}) \doteq \left\langle \hat{\phi}(\mathbf{R} + \boldsymbol{\rho}) - \frac{1}{c} (\mathbf{V}_0 + \mathbf{v}') \cdot \hat{\mathbf{A}}(\mathbf{R} + \boldsymbol{\rho}) \right\rangle_{\xi} \\ \rightarrow J_0 \left( \frac{k_{\perp} v'_{\perp}}{\Omega_a} \right) \left( \hat{\phi}(\mathbf{k}_{\perp}) - \frac{\mathbf{V}_0}{c} \cdot \hat{\mathbf{A}}(\mathbf{k}_{\perp}) - \frac{v'_{\parallel}}{c} \hat{A}_{\parallel}(\mathbf{k}_{\perp}) \right) + J_1 \left( \frac{k_{\perp} v'_{\perp}}{\Omega_a} \right) \frac{v'_{\perp}}{c} \frac{\hat{B}_{\parallel}(\mathbf{k}_{\perp})}{k_{\perp}} . \end{aligned}$$

# Gyro-Kinetic Equation for Turbulent Transport

GYRO gives complete solution with full  $(\phi, A_{\parallel}, B_{\parallel})$  electromagnetic physics.

Must also solve the electromagnetic field equations on the **fluctuation scale**:

$$\begin{aligned}\frac{1}{\lambda_D^2} \left( \hat{\phi}(\mathbf{x}) - \frac{\mathbf{V}_0}{c} \cdot \hat{\mathbf{A}} \right) &= 4\pi \sum_a e_a \int d^3v \hat{h}_a(\mathbf{x} - \boldsymbol{\rho}) , \\ -\nabla_{\perp}^2 \hat{A}_{\parallel}(\mathbf{x}) &= \frac{4\pi}{c} \sum_a e_a \int d^3v \hat{h}_a(\mathbf{x} - \boldsymbol{\rho}) v'_{\parallel} , \\ \nabla \hat{B}_{\parallel}(\mathbf{x}) \times \mathbf{b} &= \frac{4\pi}{c} \sum_a e_a \int d^3v \hat{h}_a(\mathbf{x} - \boldsymbol{\rho}) \mathbf{v}'_{\perp} .\end{aligned}$$

- ▷ Can one compute equilibrium-scale potential  $\Phi_0$  from the Poisson equation?
  - ▷ Practically, no; need higher-order theory and extreme numerical precision.
  - ▷ All codes must take care to avoid **nonphysical potential** at long wavelength
- ... F. Parra, next talk.
- ▷ TGYRO gets  $\omega_0(\psi) = -c\partial_{\psi}\Phi_0$  from the **momentum transport equation**.

# Transport Equations

## Flux-surface-averaged moments of Fokker-Planck equation

$$\left\langle \int d^3v \mathcal{A} \right\rangle_{\theta} \quad \text{density}$$
$$\left\langle \int d^3v \varepsilon \mathcal{A} \right\rangle_{\theta} \quad \text{energy}$$
$$\sum_a \left\langle \int d^3v m_a v'_{\varphi} \mathcal{A} \right\rangle_{\theta} \quad \text{toroidal momentum}$$

Only terms of order  $\rho_*^2$  survive these averages

$$\rho_*^{-1} = 10^3 \quad \rho_*^0 = 1 \quad \rho_*^1 = 10^{-3} \quad \rho_*^2 = 10^{-6}$$

# Transport Equations

Flux-surface-averaged moments of Fokker-Planck equation to  $\mathcal{O}(\rho_*^2)$

$$n_a(r) : \quad \frac{\partial \langle n_a \rangle}{\partial t} + \frac{1}{V'} \frac{\partial}{\partial r} (V' \Gamma_a) = S_{n,a}$$

$$T_a(r) : \quad \frac{3}{2} \frac{\partial \langle n_a T_a \rangle}{\partial t} + \frac{1}{V'} \frac{\partial}{\partial r} (V' Q_a) + \Pi_a \frac{\partial \omega_0}{\partial \psi} = S_{W,a}$$

$$\omega_0(r) : \quad \frac{\partial}{\partial t} (\omega_0 \langle R^2 \rangle \sum_a m_a n_a) + \frac{1}{V'} \frac{\partial}{\partial r} (V' \sum_a \Pi_a) = \sum_a S_{\omega,a}$$

$$S_{n,a} = S_{n,a}^{\text{beam}} + S_{n,a}^{\text{wall}} \quad \text{and} \quad \Gamma_a = \Gamma_a^{\text{GV}} + \Gamma_a^{\text{neo}} + \Gamma_a^{\text{tur}}$$

$$S_{W,a} = S_{W,a}^{\text{aux}} + S_{W,a}^{\text{rad}} + S_{W,a}^{\alpha} + S_{W,a}^{\text{tur}} + S_{W,a}^{\text{col}} \quad \text{and} \quad Q_a = Q_a^{\text{GV}} + Q_a^{\text{neo}} + Q_a^{\text{tur}}$$

$$\Pi_a = \Pi_a^{\text{GV}} + \Pi_a^{\text{neo}} + \Pi_a^{\text{tur}}$$

**RED: TGYRO**    **GREEN: NEO**    **BLUE: GYRO**

# A Comment on Full-F Gyrokinetic Models

## Focus on collisionless Hamiltonian substructure of model

Consider the collisionless **full-F gyrokinetic equation** (Hahm 1988):

$$\frac{dF_a}{dt} = \frac{\partial F_a}{\partial t} + [F_a, \langle H \rangle] = 0$$

- ▷ Symplectic averaging technique for **collisionless**  $d/dt$ .
- ▷ **Elegant formalism** for expansion  $\langle H \rangle = \sum_n H_n \rho_*^n$ .
- ▷  $C_a(F_a)$  and  $S_a$  are usually **ignored**, or  $C_a$  **linearized**.
- ▷ Unclear how to implement nonlinear  $C_a(F)$ .
- ▷ Unclear how to apply to transport timescale simulation  
... **M. Barnes**, previous talk.

# TGYRO Design Goals and Philosophy

## Integrated solution of Fokker-Planck hierarchy

### Goal

Solve transport equations for profiles  $n_a(r)$ ,  $T_a(r)$ ,  $\omega_0(r)$  that drive fluxes that balance specified input power (from experiment or model scenario).

### Essence of numerical approach

- ▷ Choose initial **profiles** and radial **pivot point**.
- ▷ Calculate fluxes using GYRO+NEO at radii  $\{r_j\}$ , where  $0 < r_j \leq r_{\max} < a$ .
- ▷ **Adjust gradients** until GYRO+NEO fluxes balance heating sources.
- ▷ Effectively a nonlinear **root-finding** problem.



# Brief Description of the TGYRO Solver

## Illustration: Steady-state problem for energy transport

Balance between **transport flux**,  $Q_a = Q_a^{\text{tur}} + Q_a^{\text{neo}}$ , and **target flux**,  $Q_a^T$ :

**divergence form (standard)**  $\frac{1}{V'} \frac{\partial}{\partial r} (V' Q_a) = S_a(r)$

**integral form (TGYRO)**  $Q_a(\{z_a, T_a\}, r) = Q_a^T(r)$

$$Q_a^T(r) \doteq \frac{1}{V'(r)} \int_0^r dx V'(x) S_a(x) \quad \text{and} \quad z_a \doteq -\frac{1}{T_a} \frac{\partial T_a}{\partial r} .$$

Specify the temperature at a matching radius  $r_*$  and discretize  $r \rightarrow \{r_j\}$ :

$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

$$T_a(r_{j-1}) = T_a(r_j) \exp \left\{ \frac{1}{2} [z_a(r_j) + z_a(r_{j-1})] [r_j - r_{j-1}] \right\} .$$

# Brief Description of the TGYRO Solver

## Use integrated form of the transport equations

Solve the discrete equations using gyroBohm normalization ( $\widehat{Q} \doteq Q/Q_{\text{GB}}$ ):

$$\widehat{Q}_{a,j}(\{z_{a,j}\}) = \widehat{Q}_{a,j}^T(\{z_{a,j}\})$$

The problem is thus reduced to  **$n$ -dimensional root-finding**:

$$f_{a,j}(\{z_{a,j}\}) = 0 \quad \rightarrow \quad \left( \frac{\partial f}{\partial z} \right)_{aa',jj'} \delta z_{a',j'} = -\eta_{a,j} f_{a,j}$$

▷  $\widehat{Q}_{a,j} = \widehat{Q}_{a,j}^{\text{tur}} + \widehat{Q}_{a,j}^{\text{neo}}$  are the **transport fluxes** (GYRO, NEO, TGYRO)

→ for economy, **use block-diagonal Jacobian**:  $\frac{\partial \widehat{Q}_{a,j}}{\partial z_{a',j}} \delta_{jj'}$ .

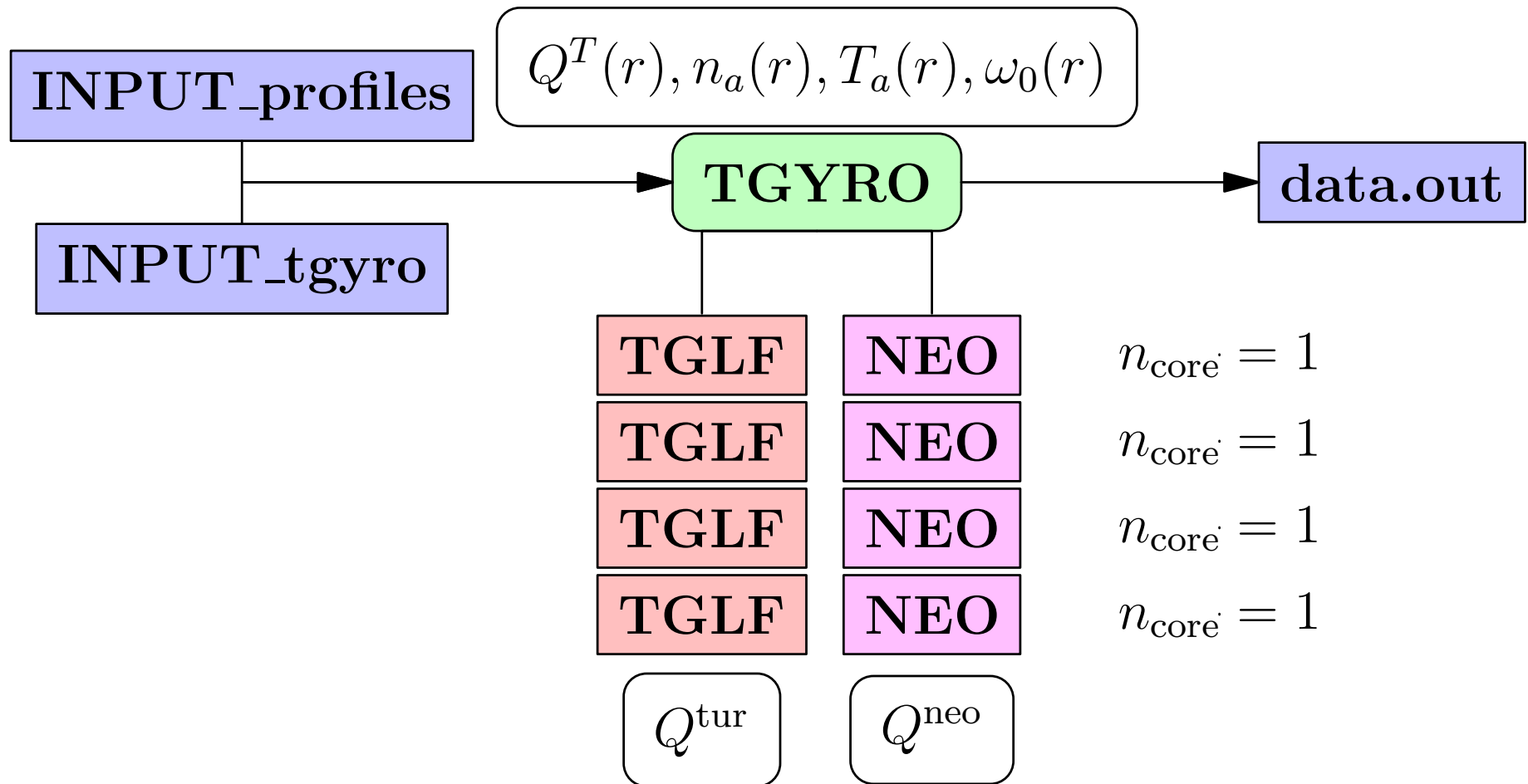
→ statistical steady state and large  $\Delta z_{a,j}$  required for Jacobian calculation.

▷  $\widehat{Q}_{a,j}^T$  are the **target fluxes** (INEXPENSIVE)

→ for accuracy, **use dense Jacobian**:  $\frac{\partial \widehat{Q}_{a,j}^T}{\partial z_{a',j'}}$ .

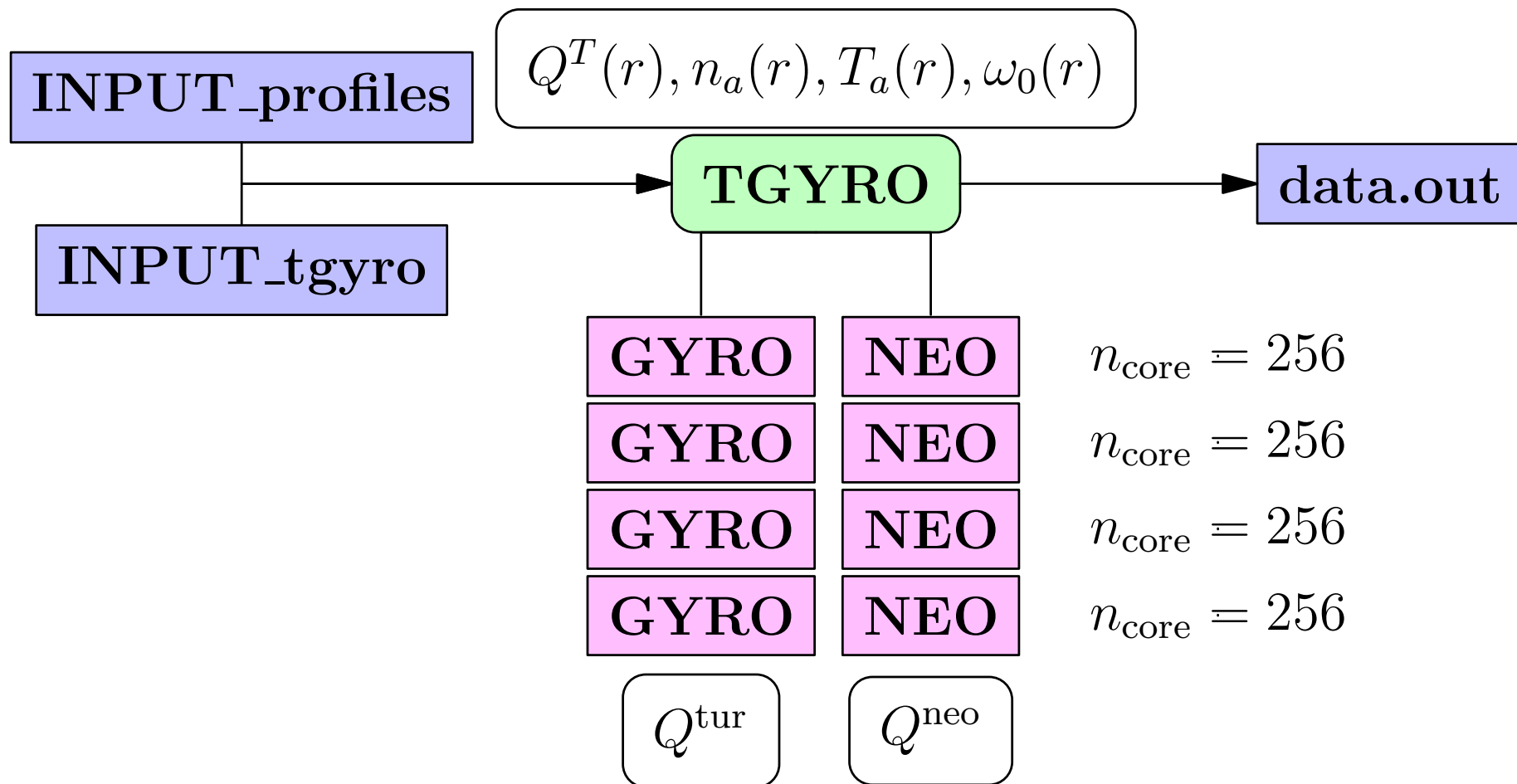
# TGYRO Structure

TGYRO can use TGLF for turbulent fluxes,  $Q^{\text{tur}}$



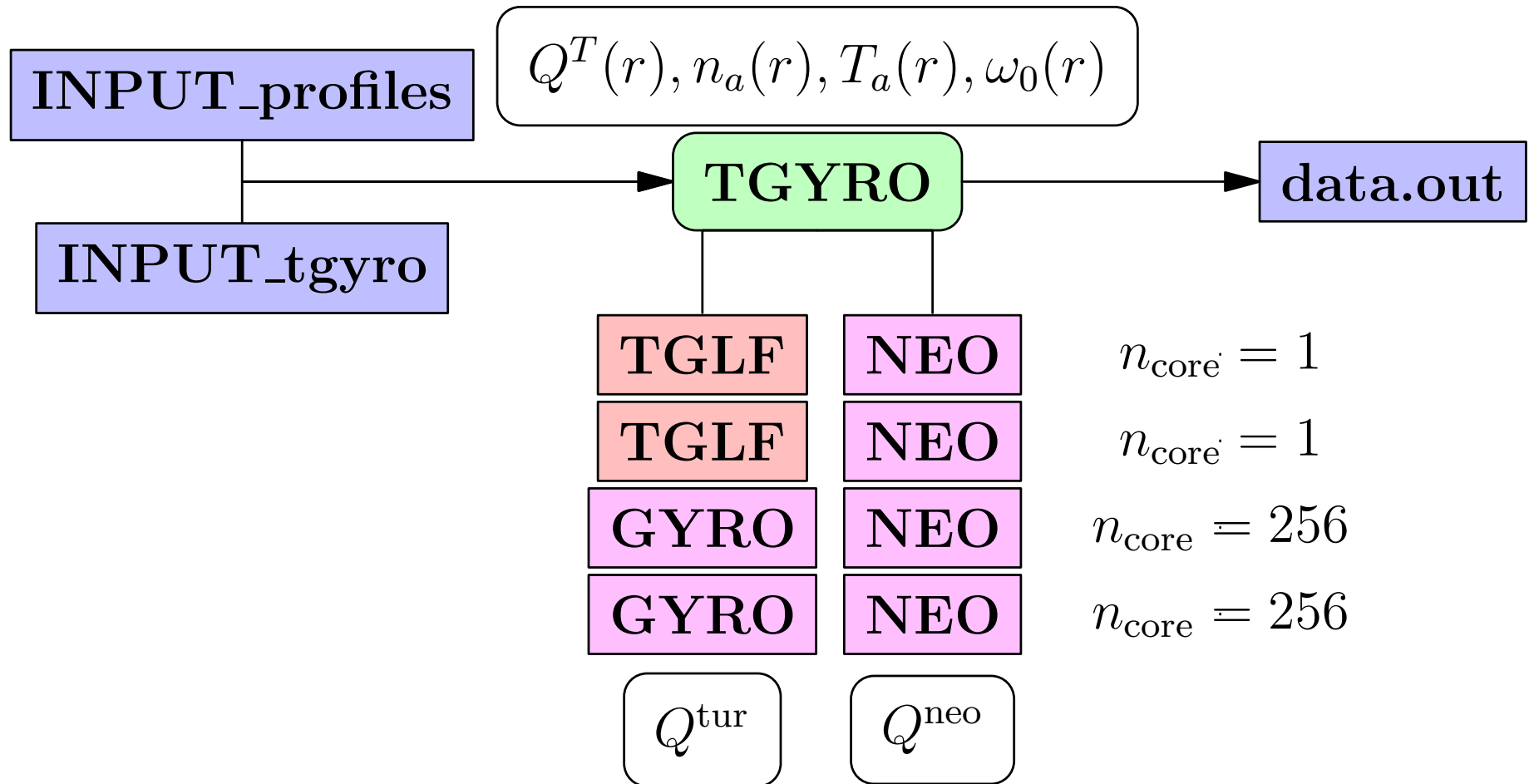
# TGYRO Structure

TGYRO can use full nonlinear GYRO runs for turbulent fluxes,  $Q^{\text{tur}}$



# TGYRO Structure

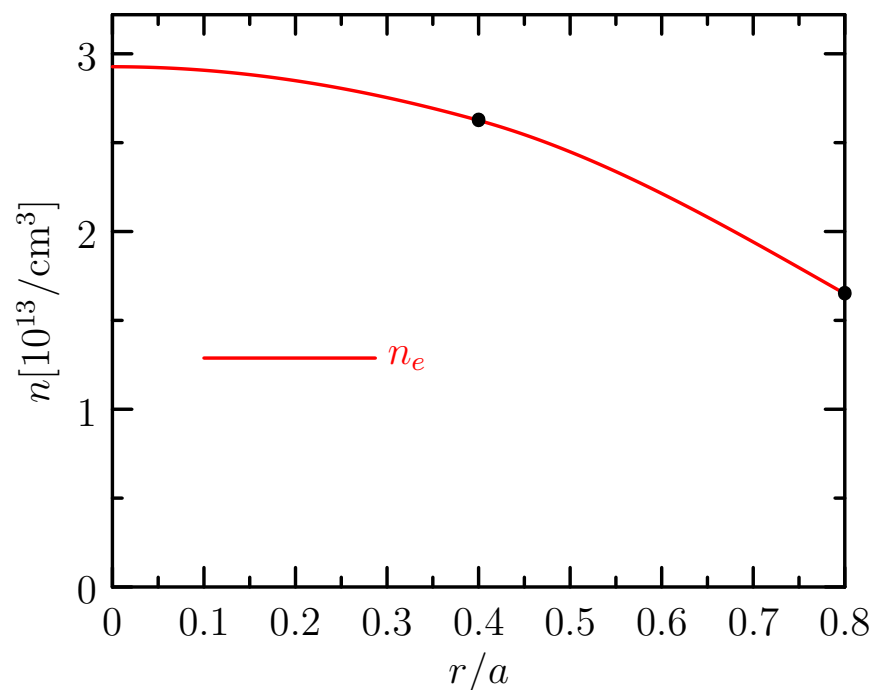
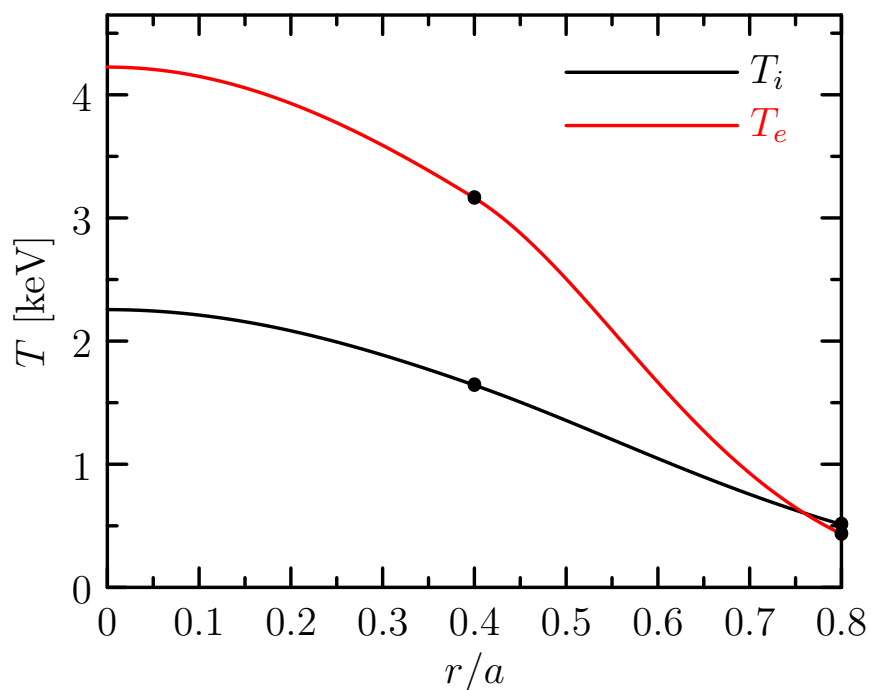
We can also mix turbulence models



# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

2 radial nodes



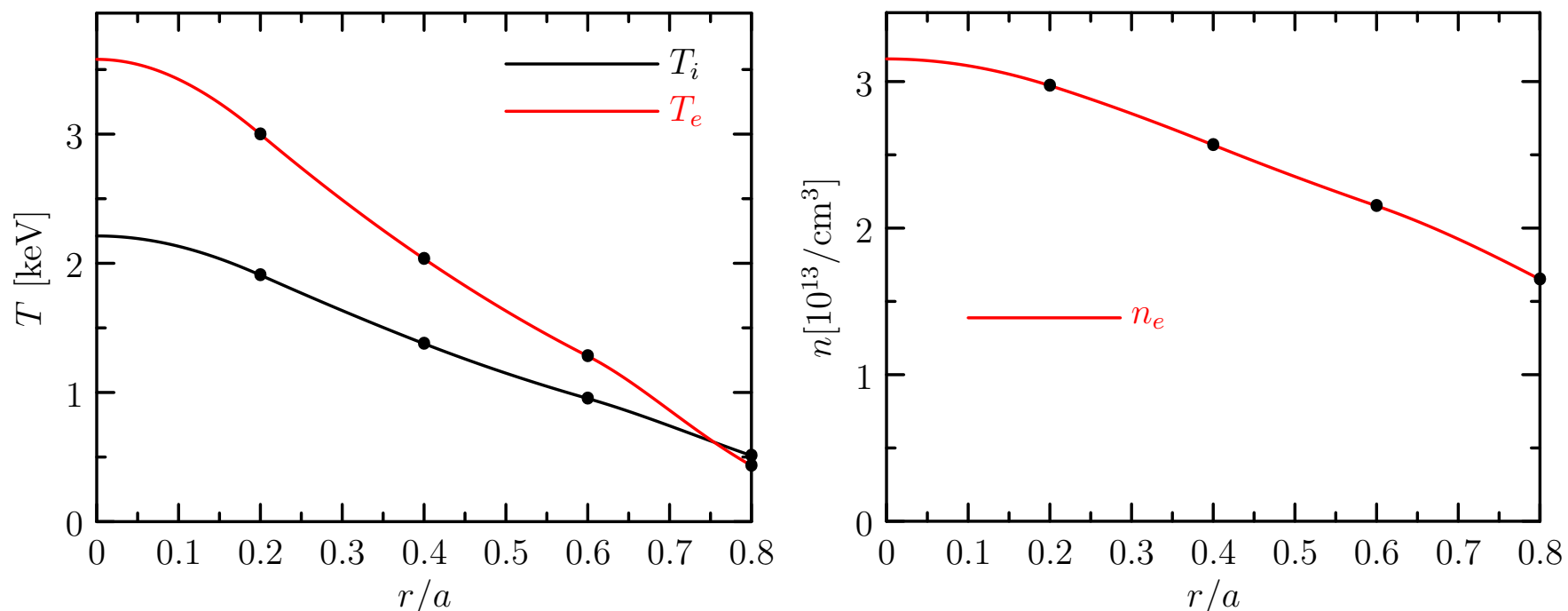
$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

$z_a(x)$  taken to be **piecewise-linear**

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

4 radial nodes



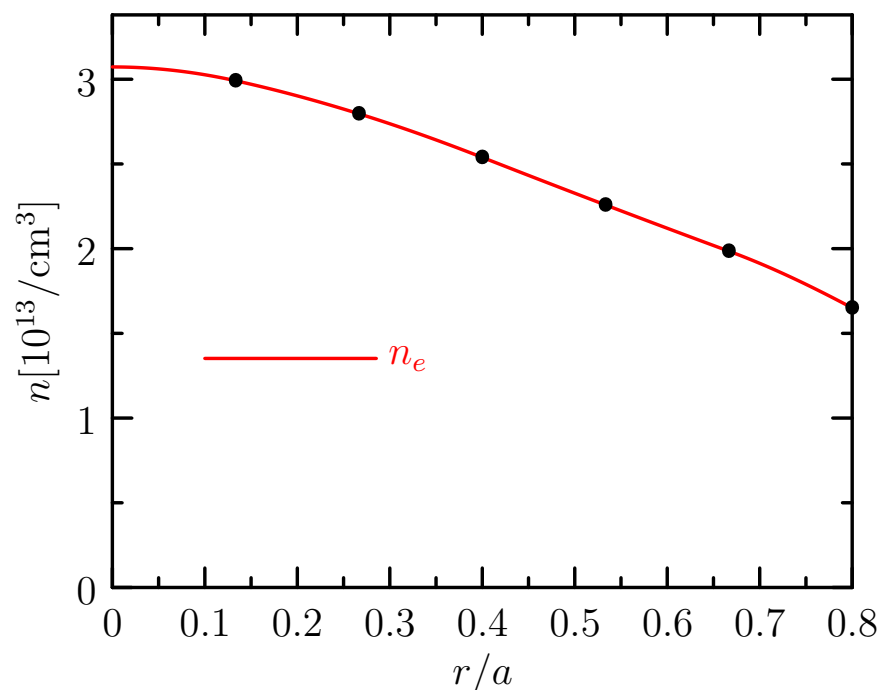
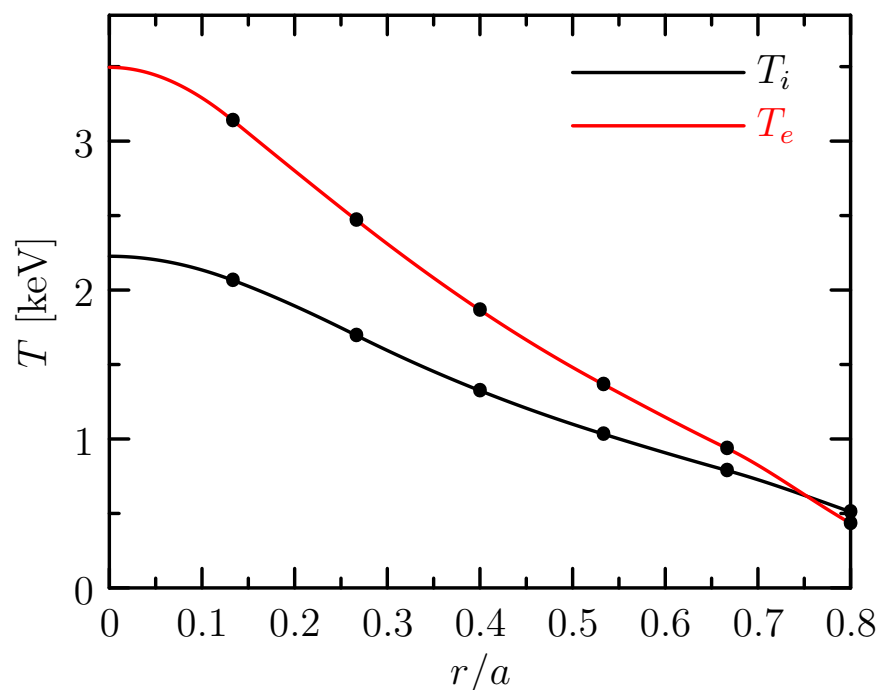
$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

$z_a(x)$  taken to be **piecewise-linear**

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

6 radial nodes



$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

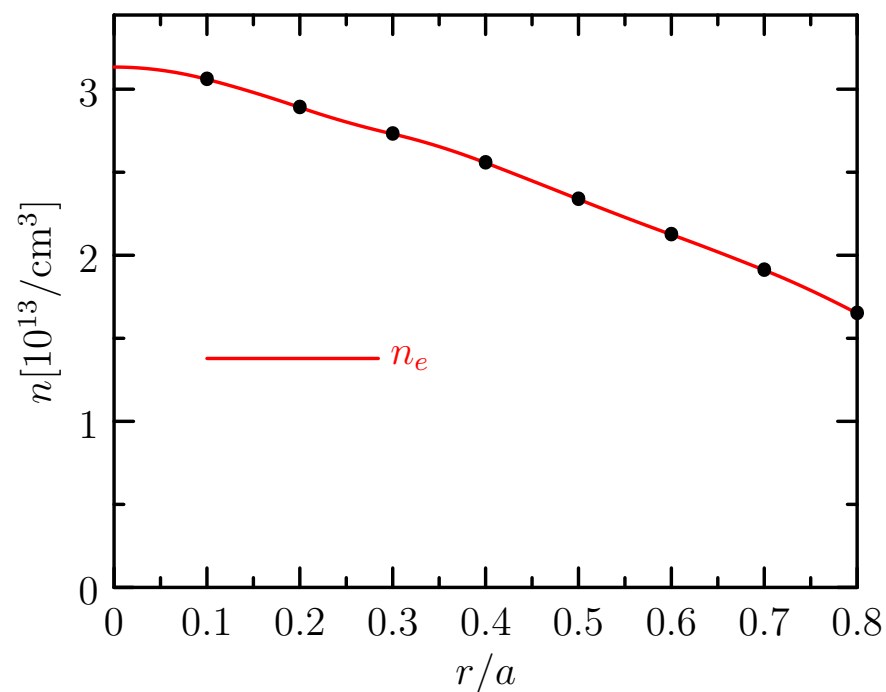
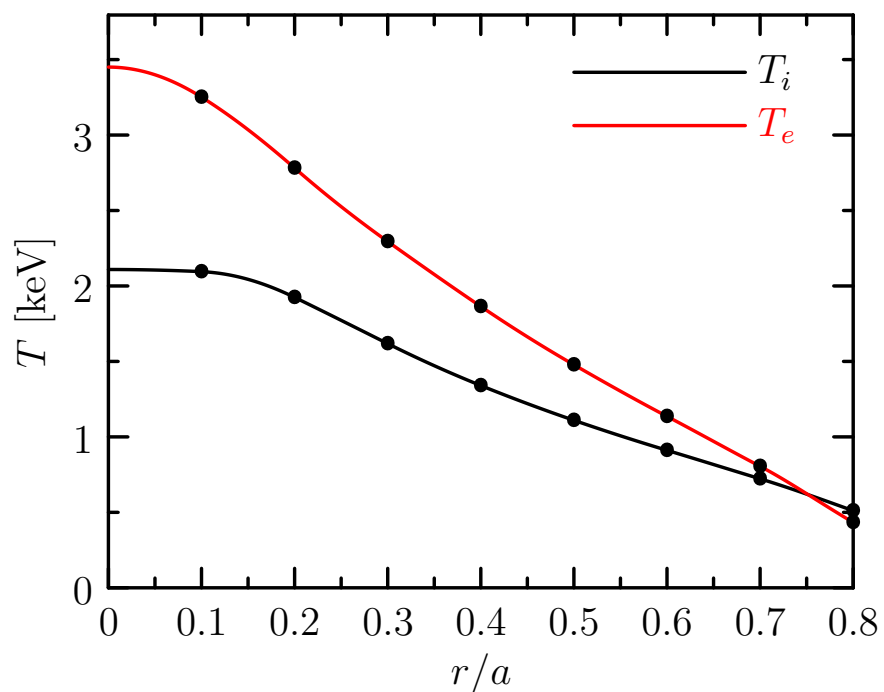
$z_a(x)$  taken to be **piecewise-linear**



# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

8 radial nodes



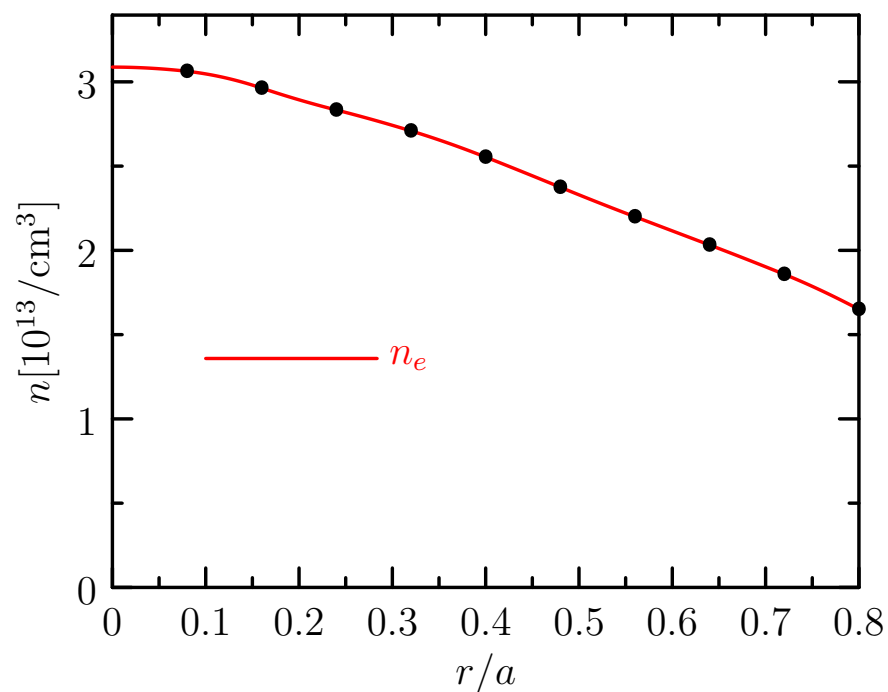
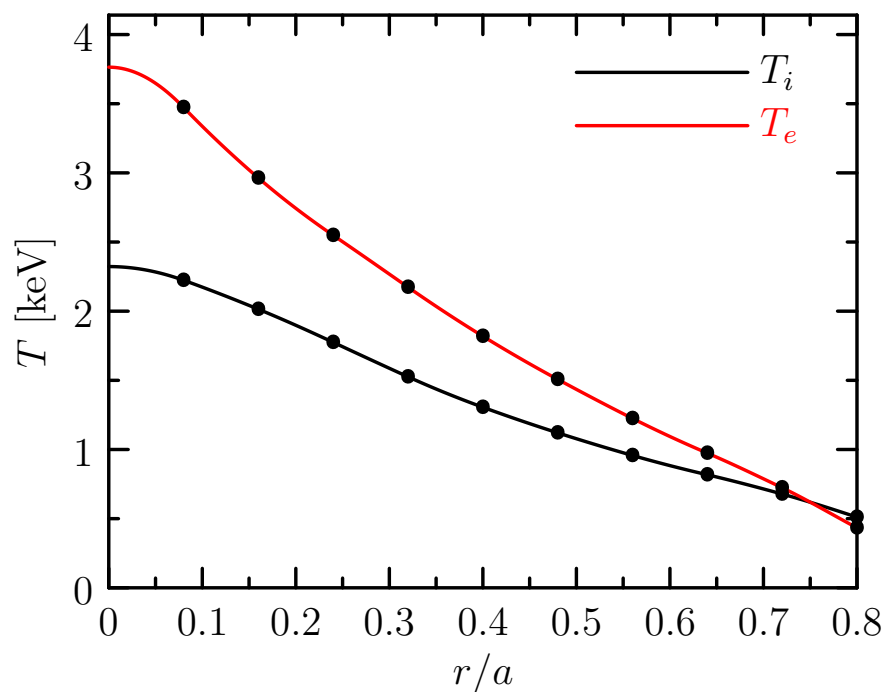
$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

$z_a(x)$  taken to be **piecewise-linear**

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

10 radial nodes



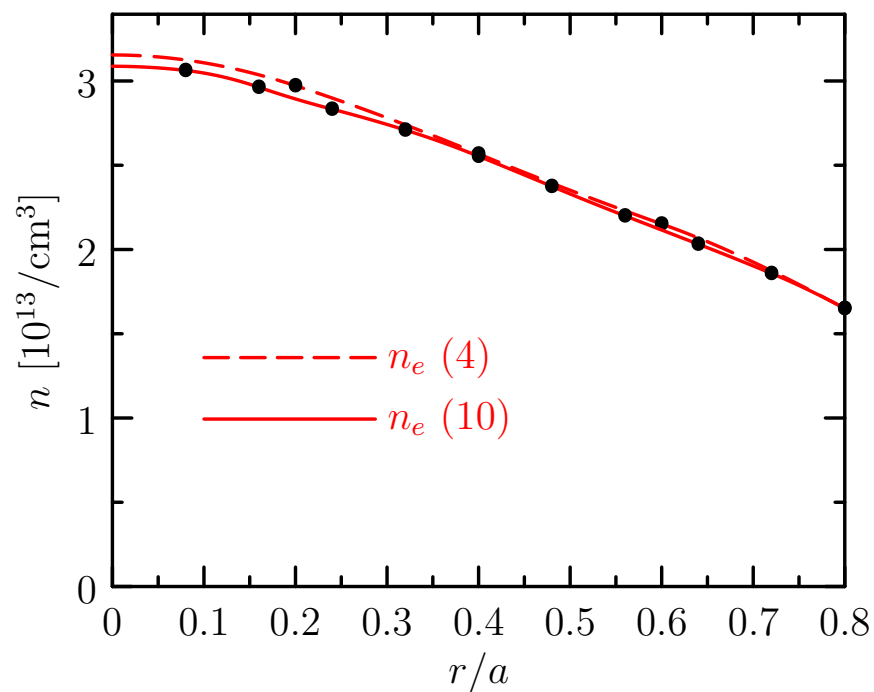
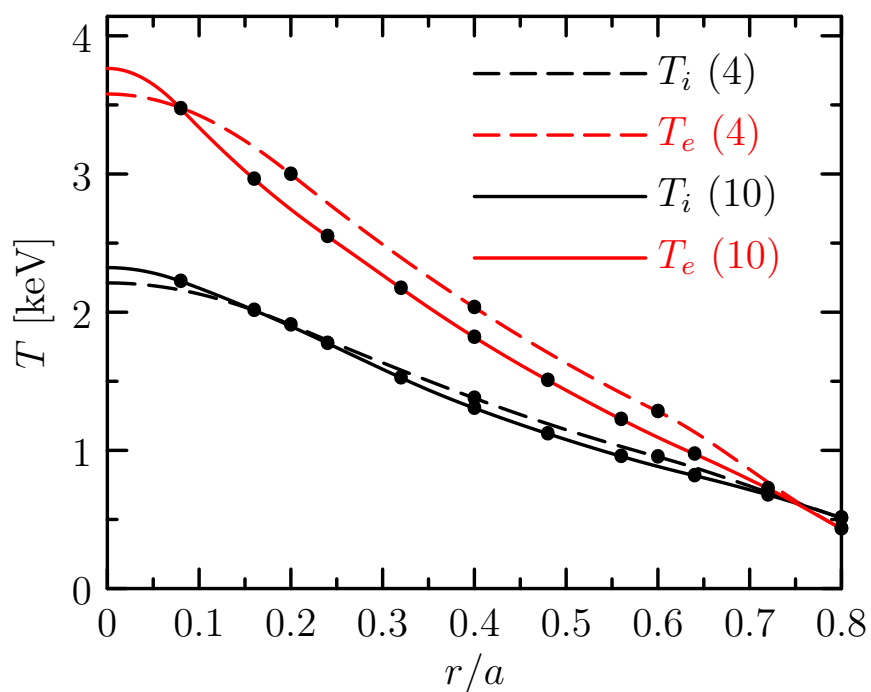
$$T_a(r) = T_{a*} \exp \left( \int_r^{r_*} dx z_a(x) \right)$$

$z_a(x)$  taken to be **piecewise-linear**

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

Rapid profile convergence with number of radial nodes

## 4 vs 10 radial nodes

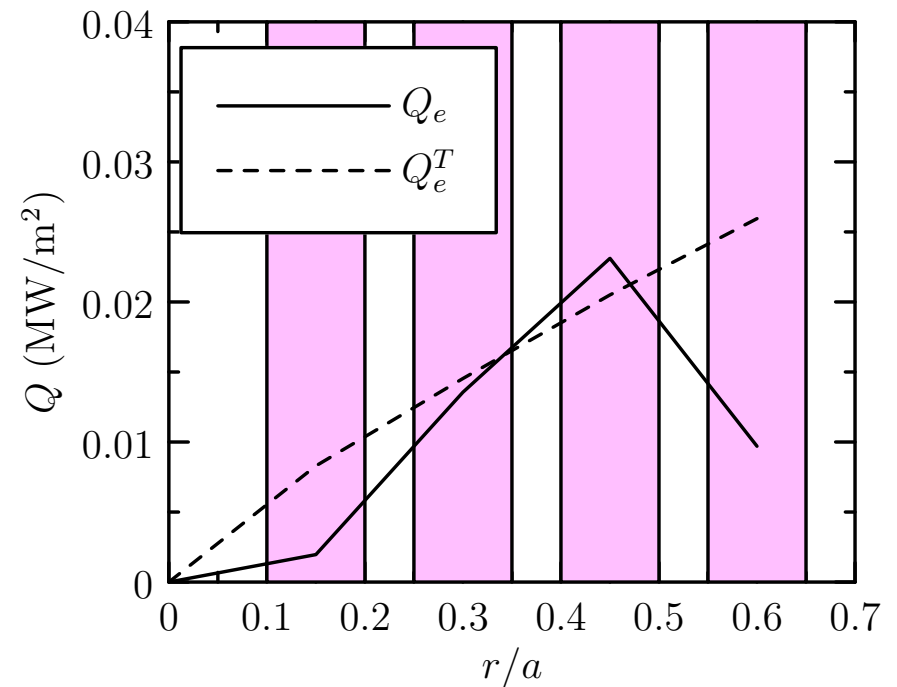
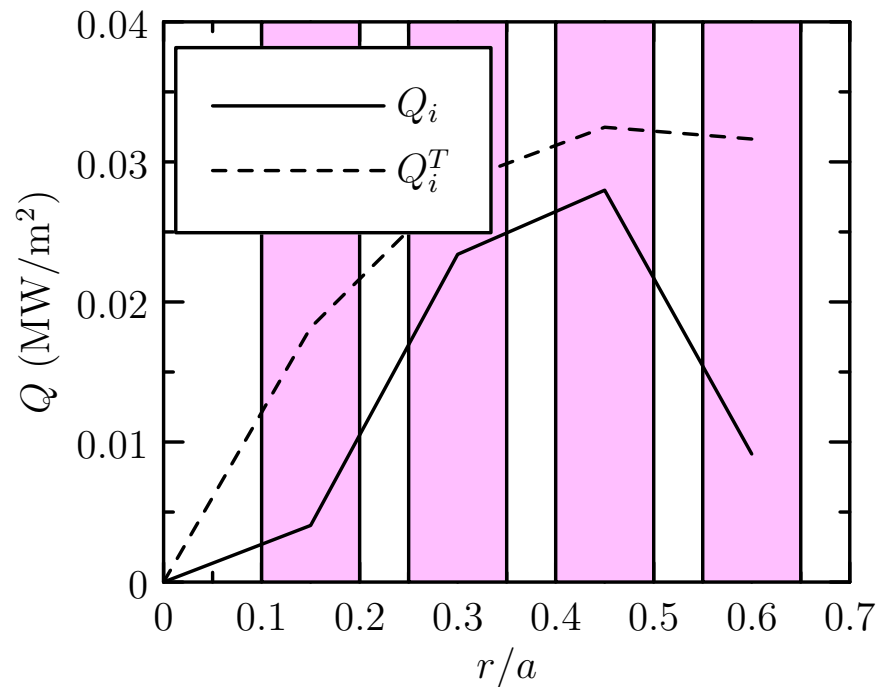


- ▷ Prediction for typical DIII-D L-mode plasma converged with 10 radial nodes.
- ▷ Very good result with **only 4 nodes!**

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 0



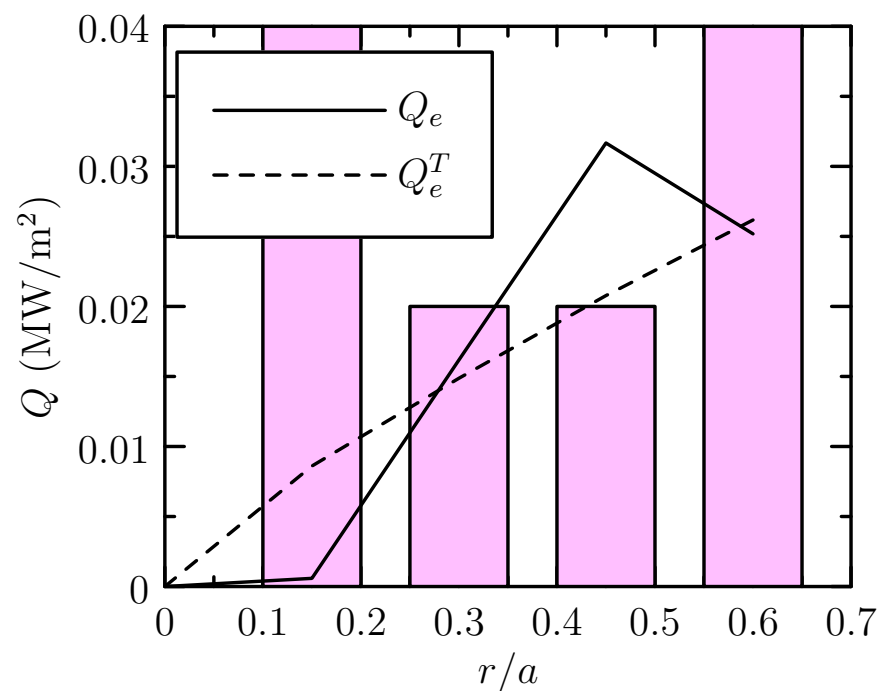
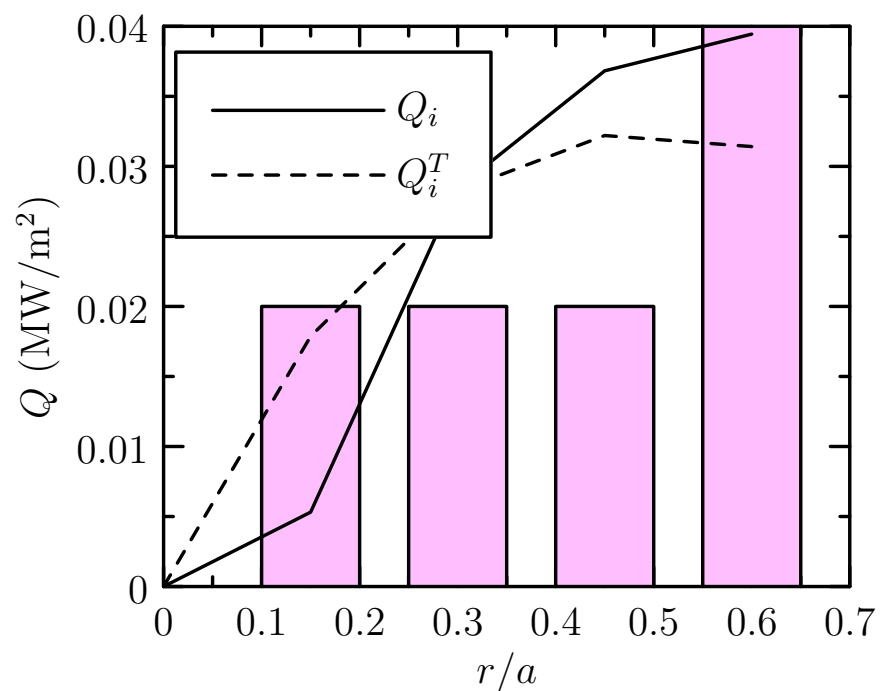
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 1



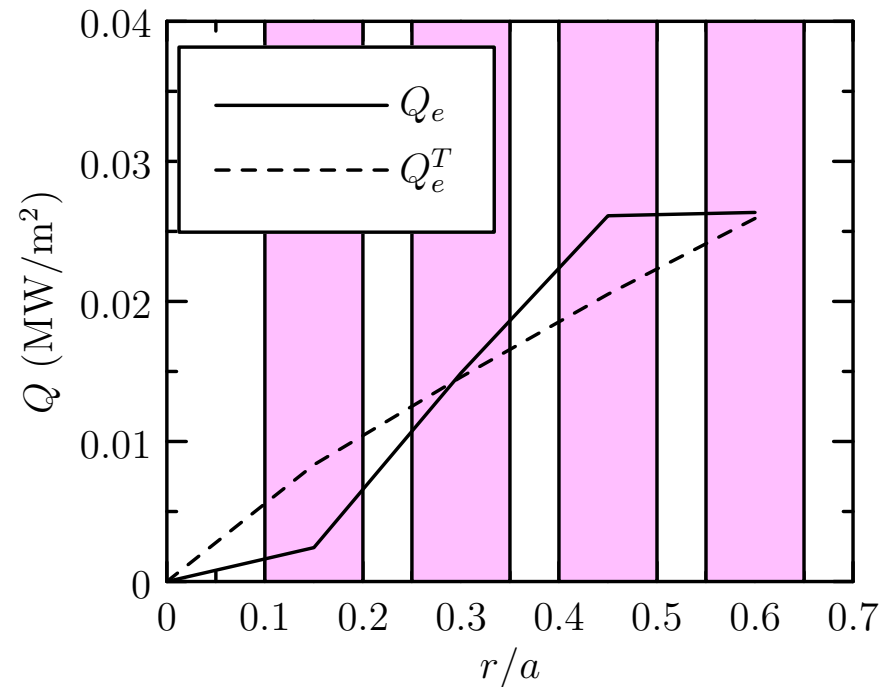
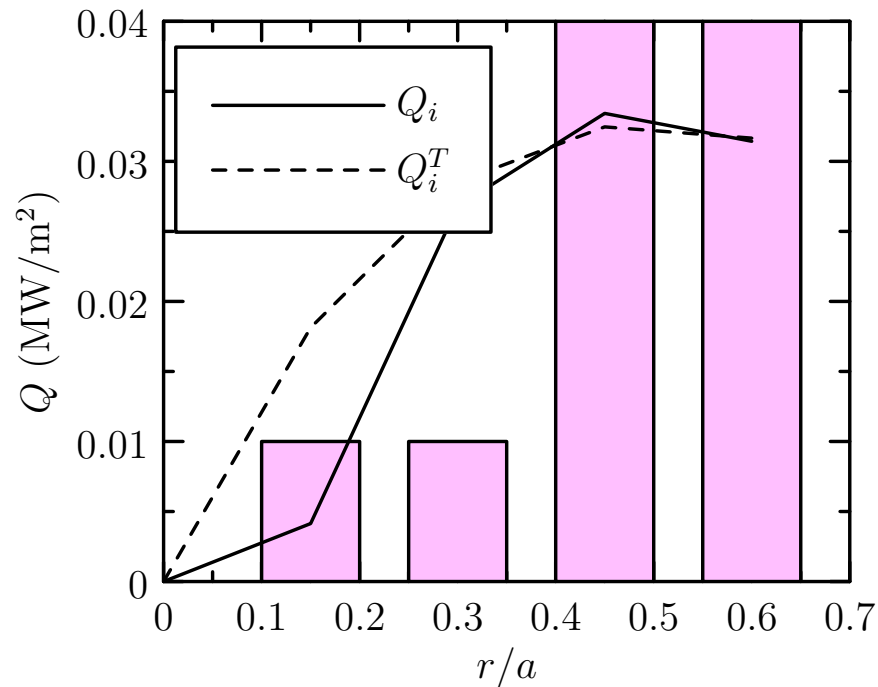
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 2



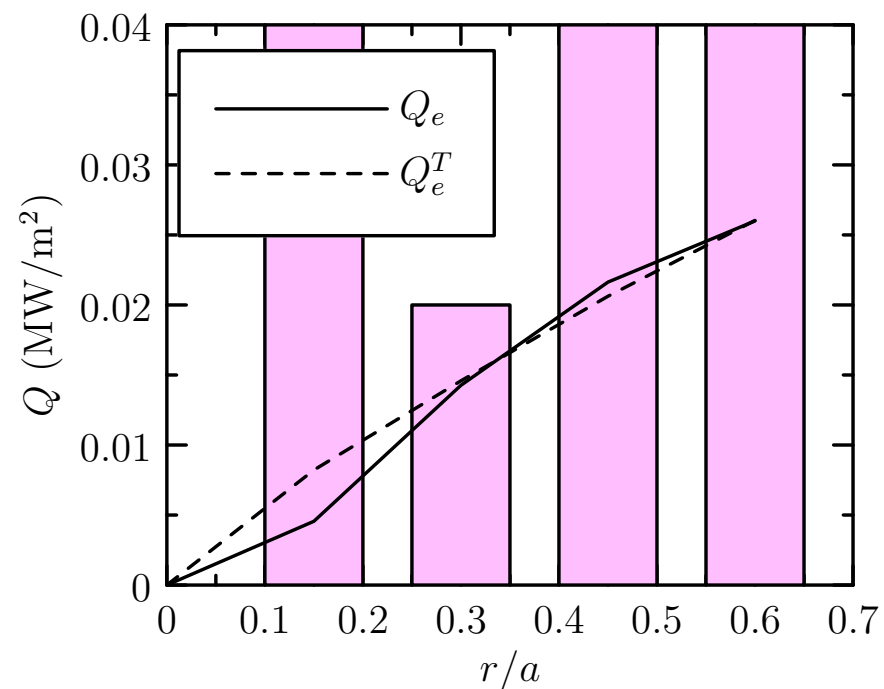
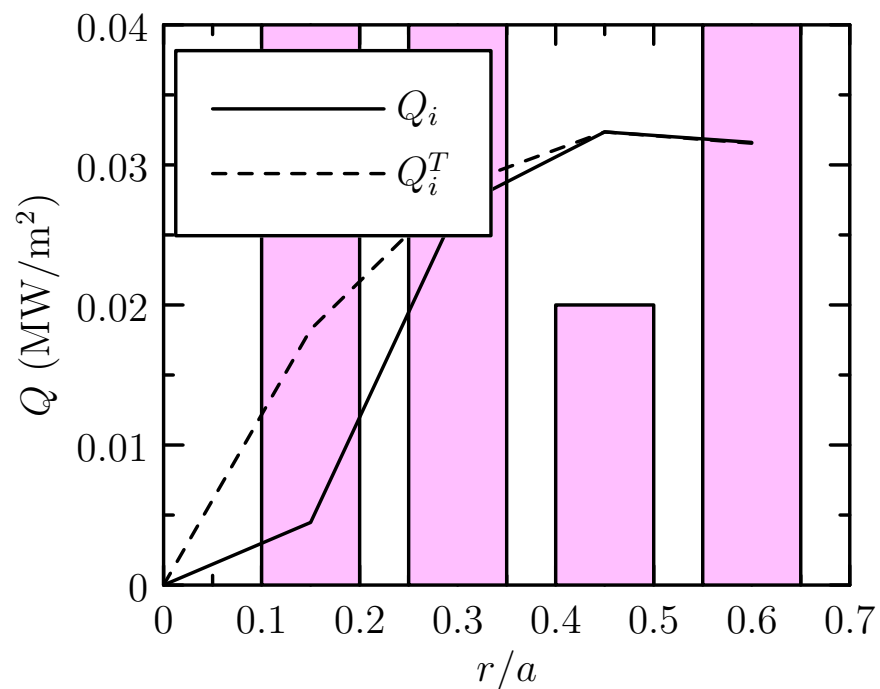
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 3



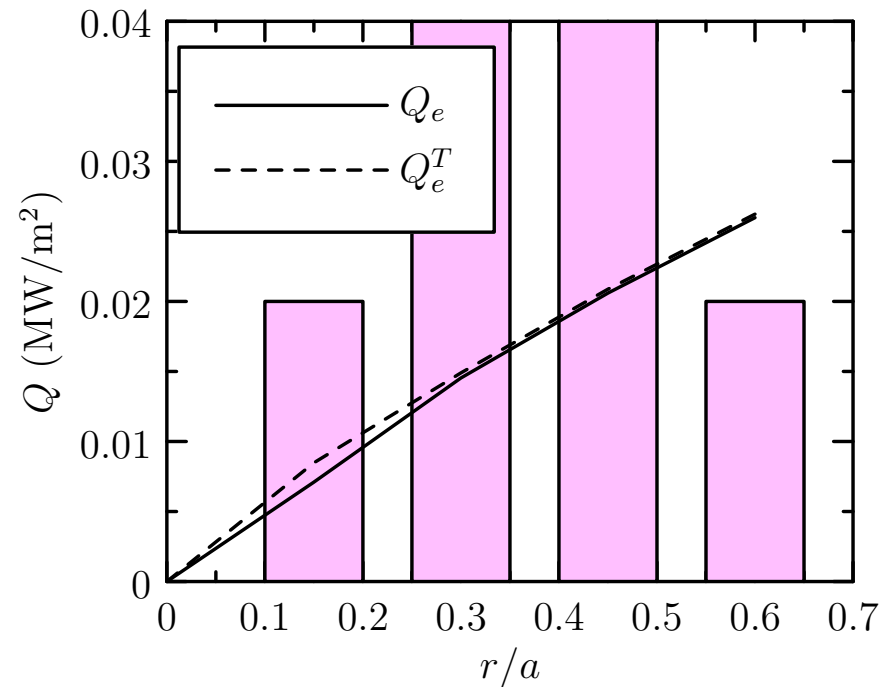
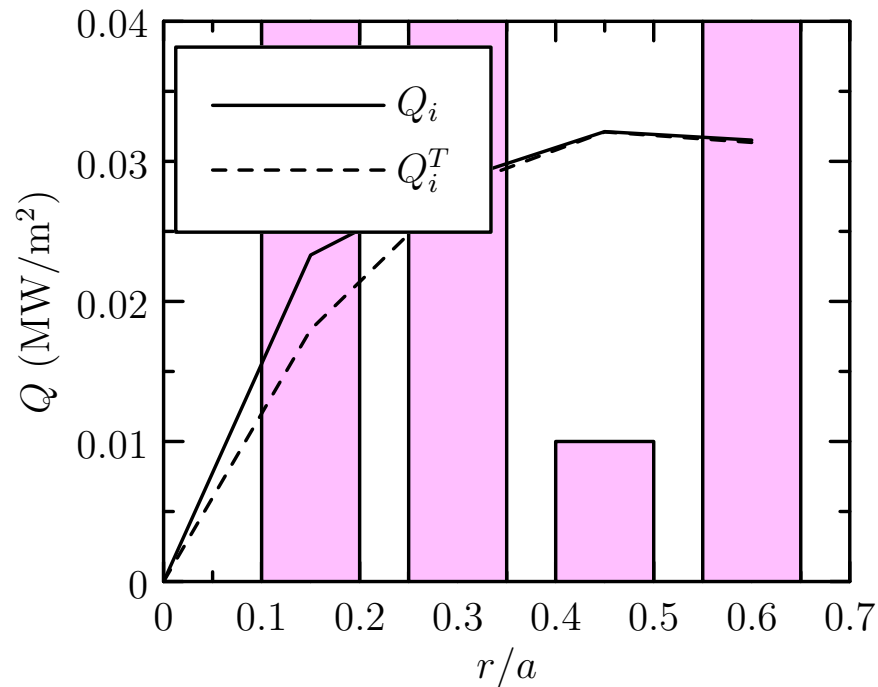
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 4



Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

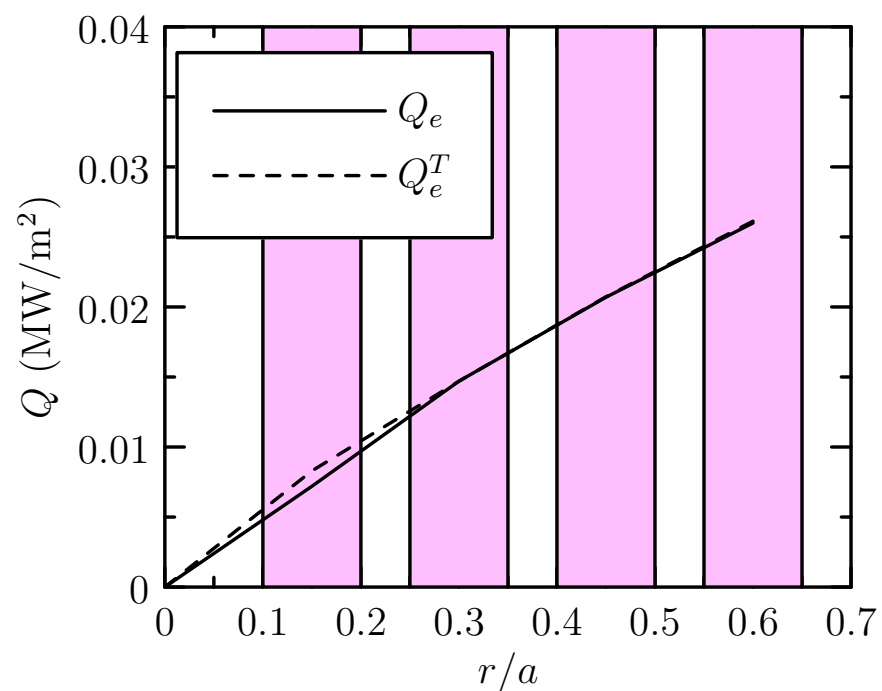
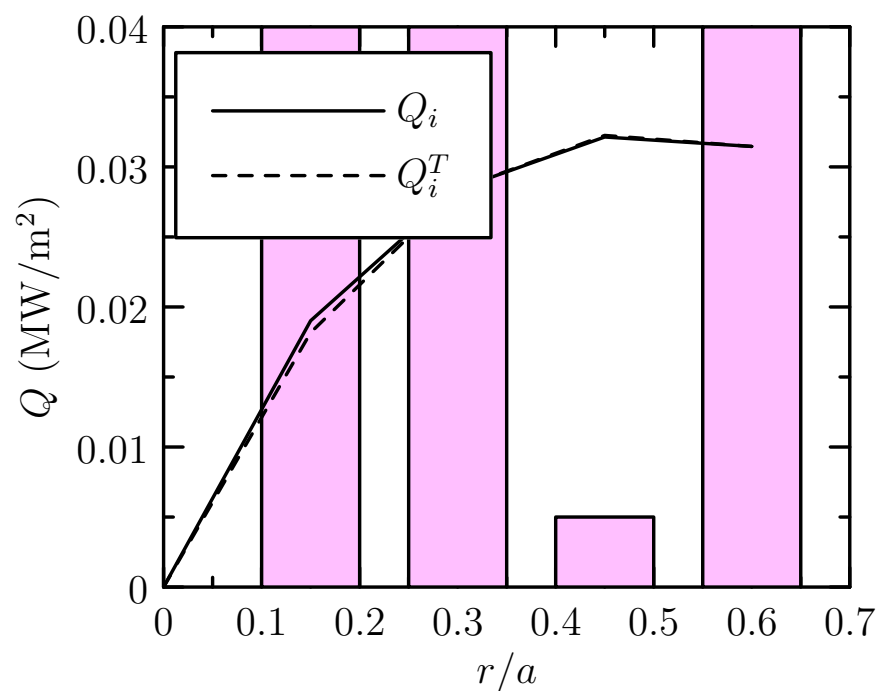
$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$



# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 5



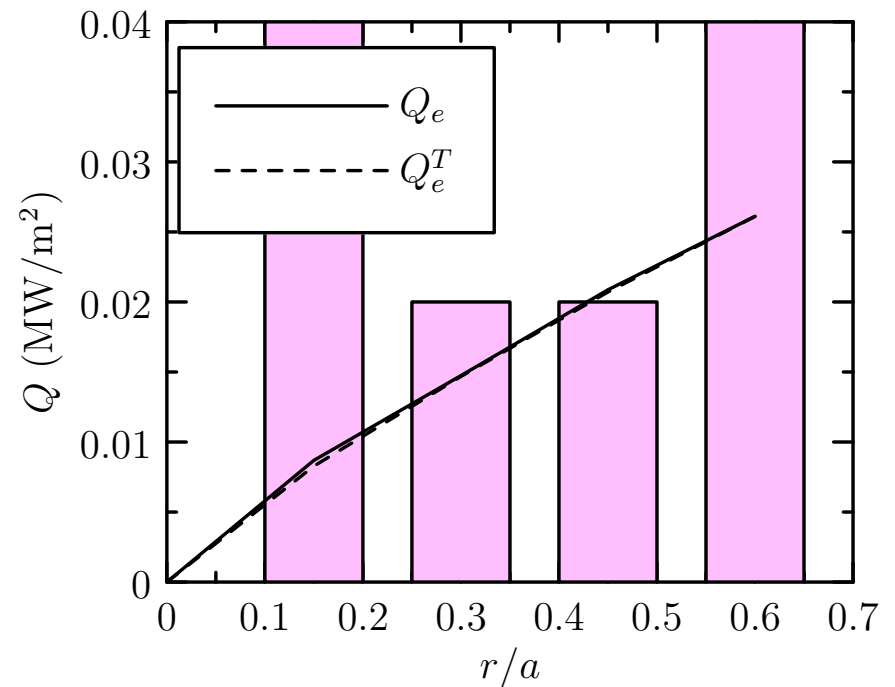
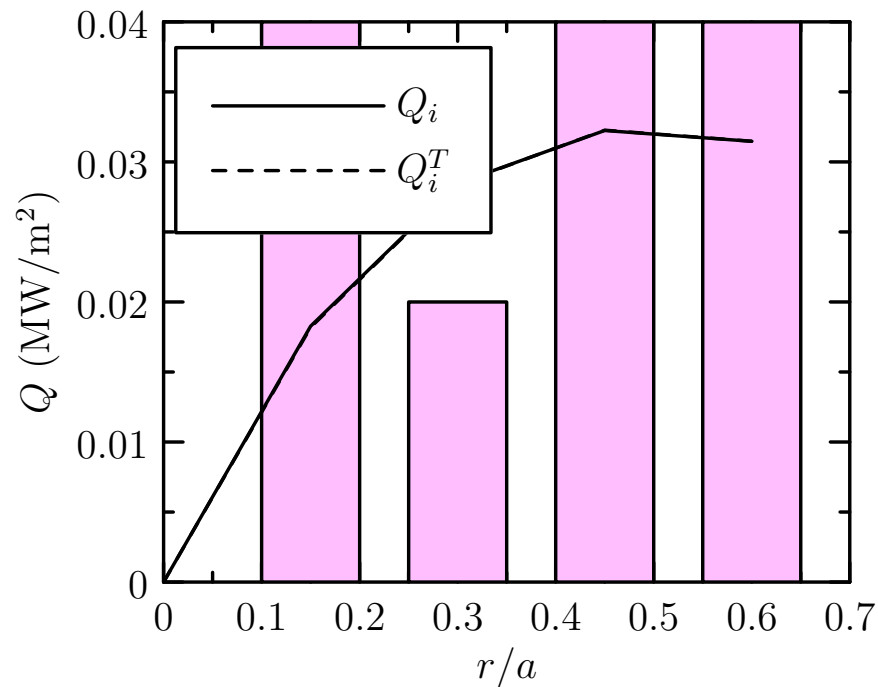
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 6



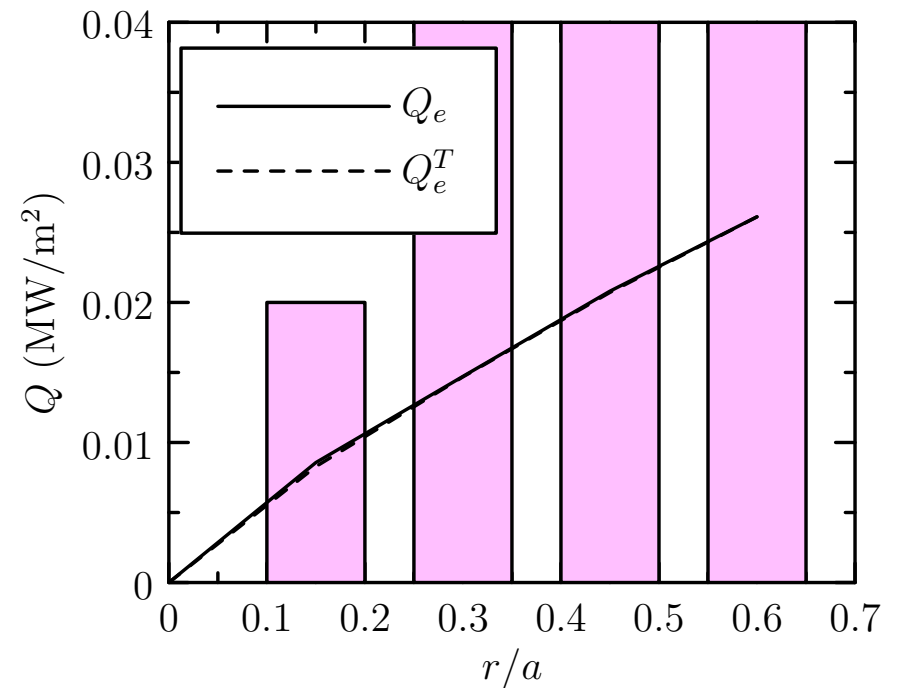
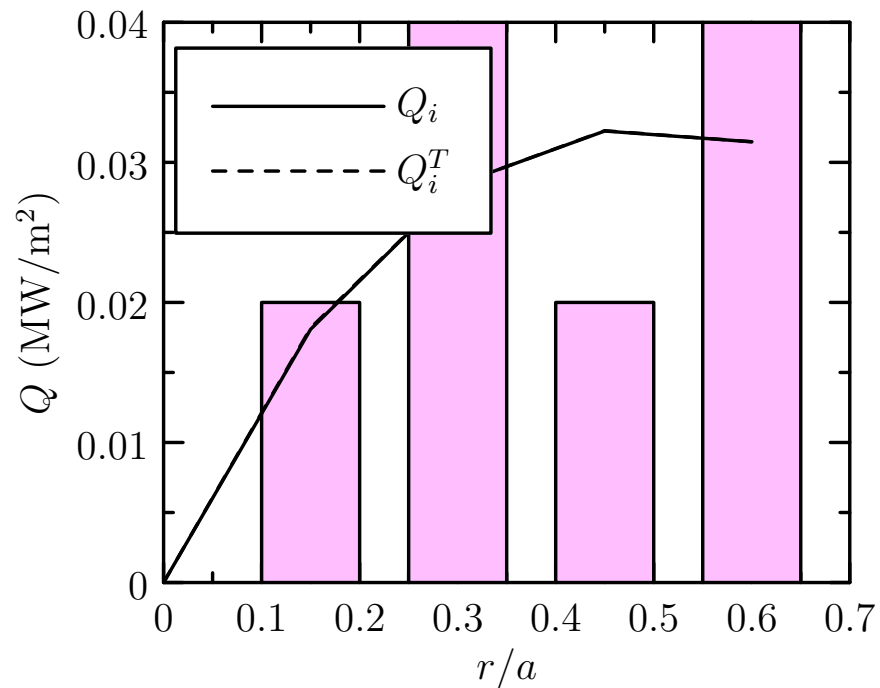
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 7



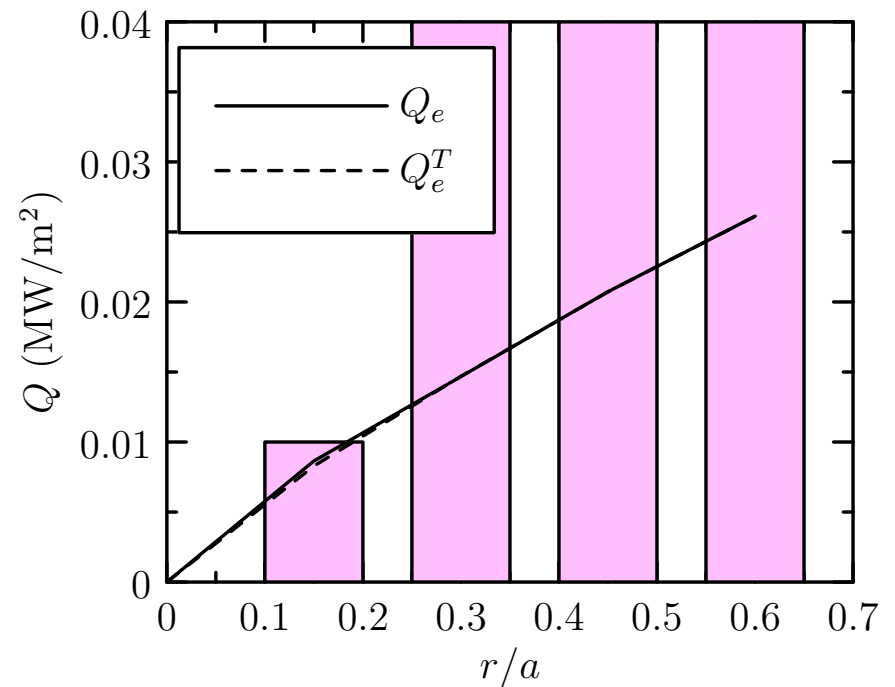
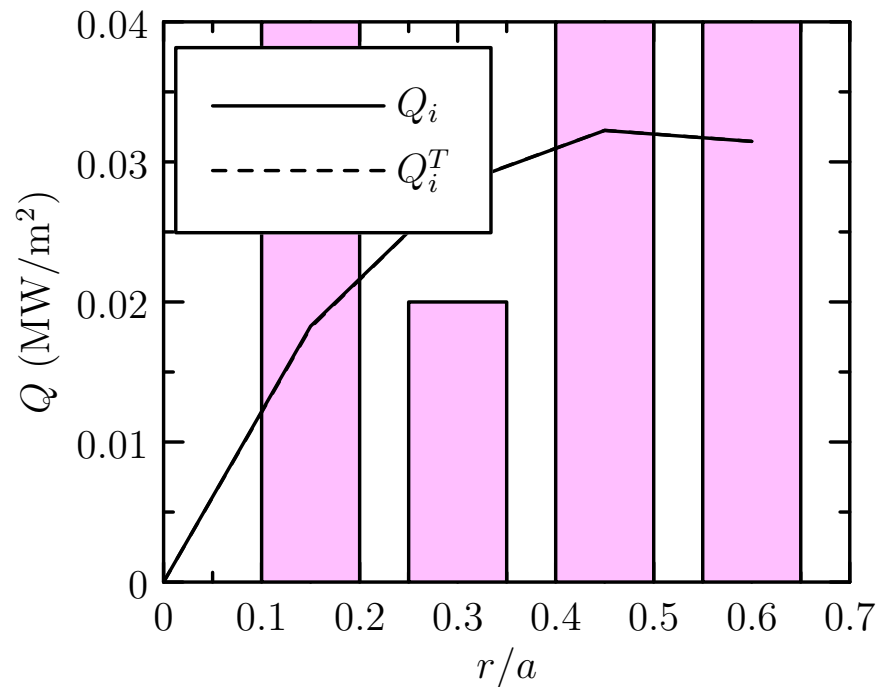
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 8



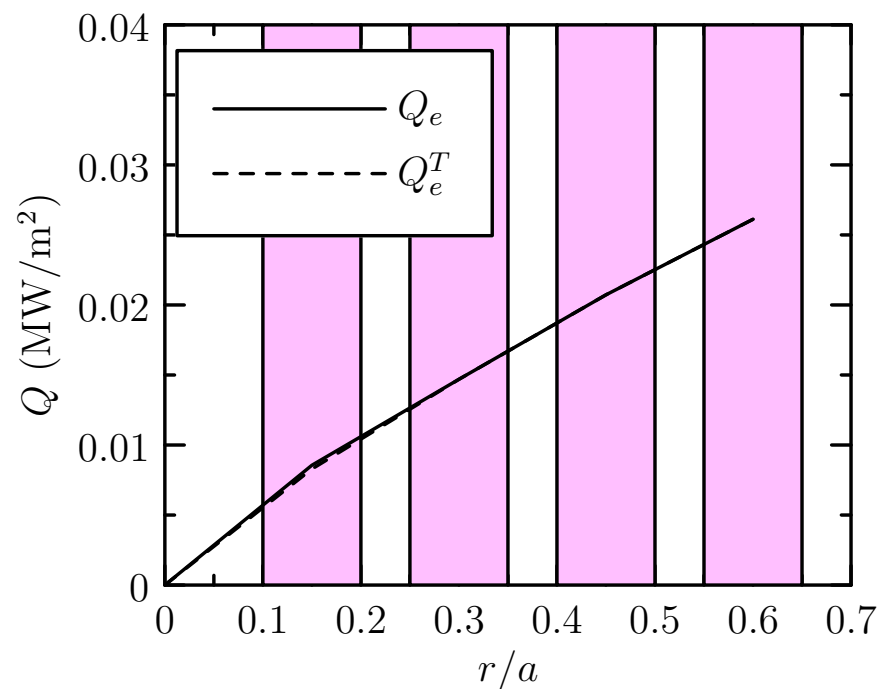
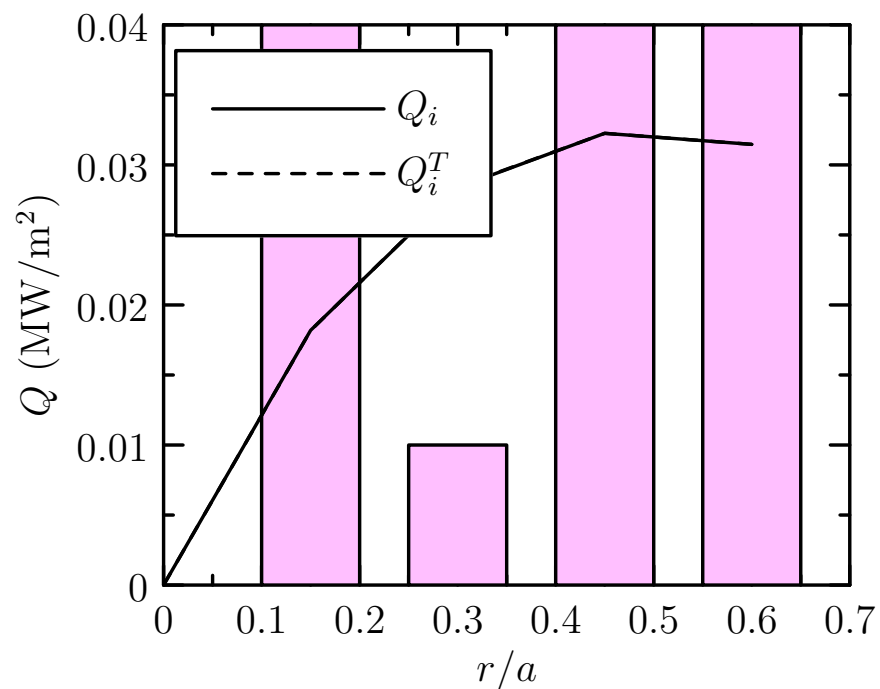
Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# TGYRO-TGLF Test Case (DIII-D L-mode discharge 128913)

4-node case, no density evolution

Iteration 9



Adjust relaxation parameter  $\eta_{a,j}$  based on **local residual**

$$R_{a,j} = \sum_{a,j} \left| \hat{Q}_{a,j} - \hat{Q}_{a,j}^T \right|$$

# Application of Synthetic Diagnostics

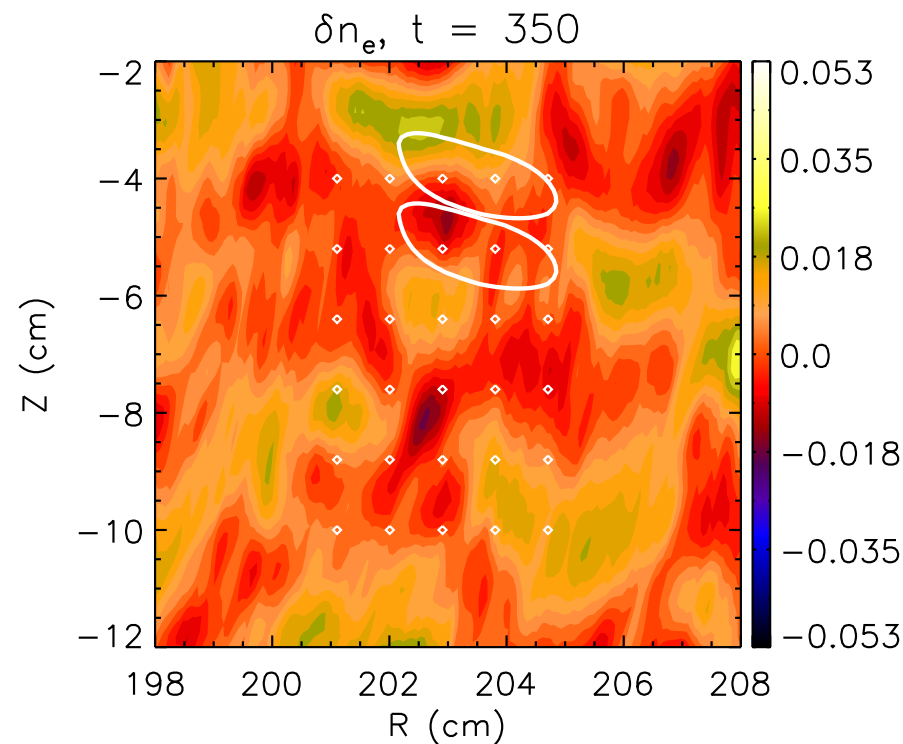
## Essential component of code-experiment comparisons

- ▷ Experimental diagnostic output is often not directly comparable with simulation data (for example, frequency/wavenumber sensitivity).
- ▷ One solution is to mimic the effect of the diagnostic via a **transfer function** which acts on the raw simulation data to produce **filtered** simulation data.
- ▷ Work still underway to define statistical metrics to quantify the **goodness of agreement** (poster by C. Holland, Thurs. 9:30am, DIII-D section, ID TP8.00012).
- ▷ For complete details:
  - C. Holland, A.E. White, G.R. McKee, M.W. Shafer, J. Candy, R.E. Waltz, et al., Phys. Plasmas **16** (2009) 052301.
- ▷ In order of complexity, we can perform synthetic diagnostic analysis on
  1. GYRO simulation using **pure experimental input data**.
  2. GYRO simulation based on converged **TGYRO[TGLF+NEO] prediction**.
  3. GYRO simulation based on converged **TGYRO[GYRO+NEO] prediction**.

# Synthetic Beam Emission Spectroscopy (BES)

From C. Holland *et al.*, Phys. Plasmas 16 (2009) 052301.

BES measures **local density fluctuations** via 2D array ( $5 \times 6$ );  
Poloidal channels cross-correlated to improve S/N ratio;  
50% **point-spread function** contours shown.

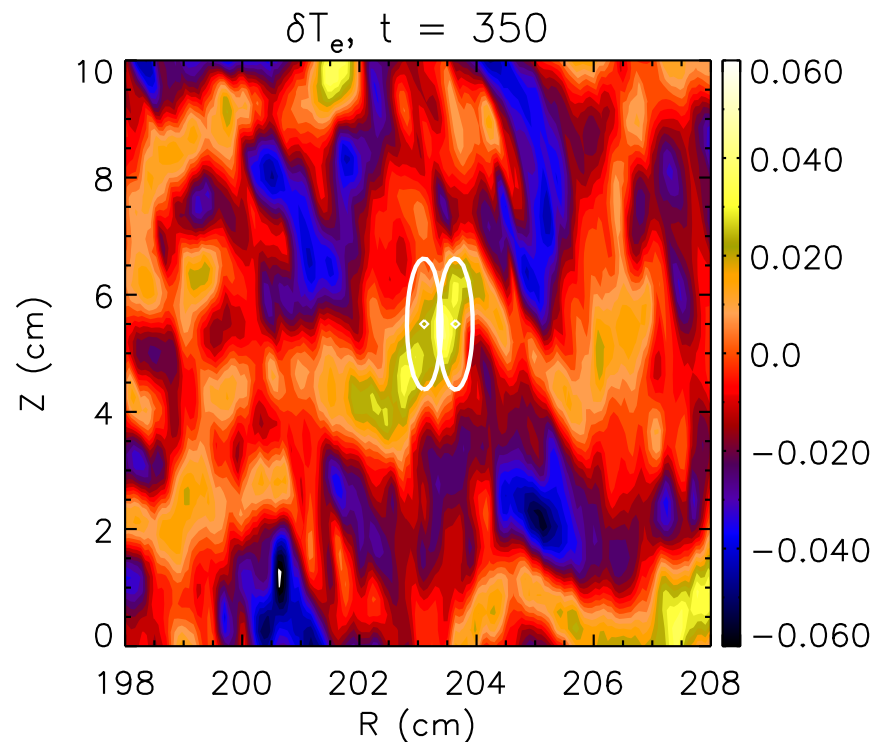


G.R. McKee, R.J. Fonck, D.K. Gupta, D.J. Schlossberg, *et al.*, Plasma Fusion Res. **2**, S1025 2007.

# Synthetic Correlation Electron Cyclotron Emission (CECE)

From C. Holland *et al.*, Phys. Plasmas 16 (2009) 052301.

Measures local  $T_e$  fluctuations via radial cross-correlation;  
50% point-spread function contours shown.



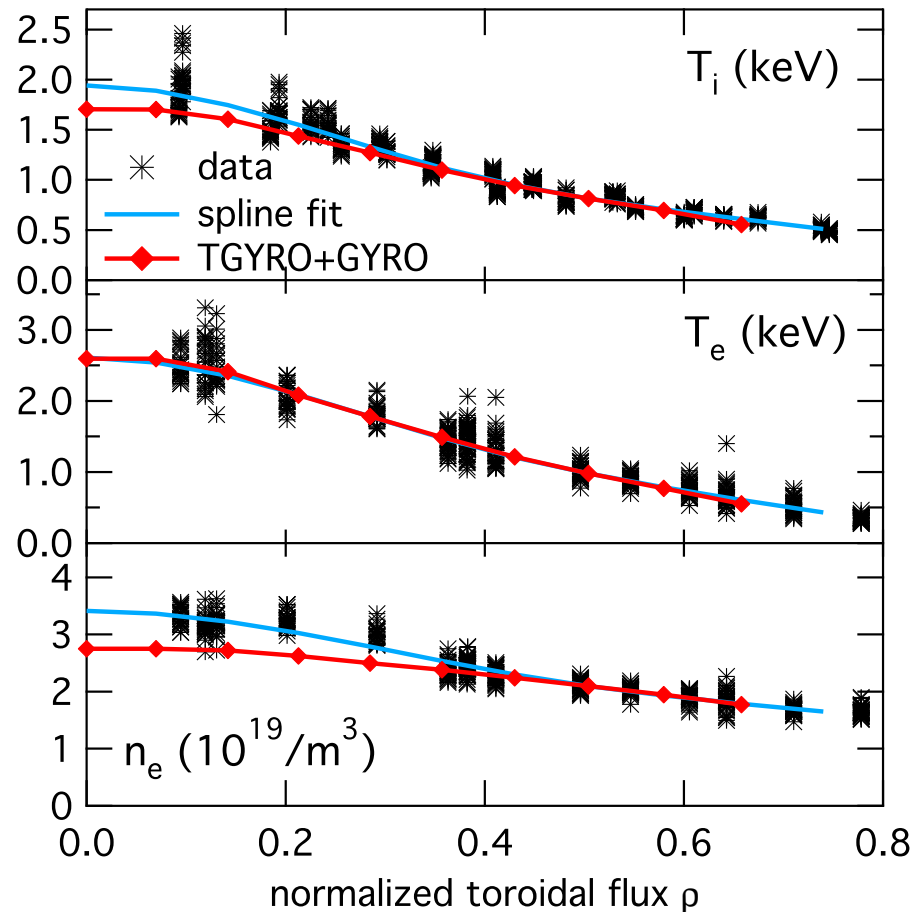
A.E. White, L. Schmitz, G.R. McKee, C. Holland, *et al.*, Phys. Plasmas **15**, 056116 2008.



# Measured Versus Simulated Profiles (DIII-D 128913 1500ms)

Plasma core is non-stationary.

GYRO simulation near top of pedestal **does not converge**; **strong ETG regime**

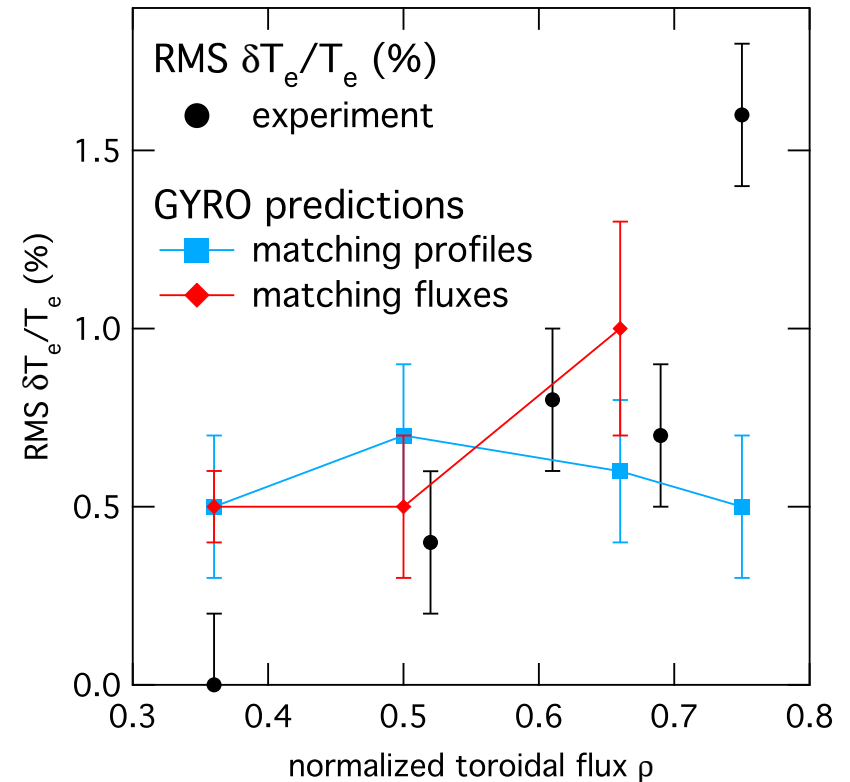
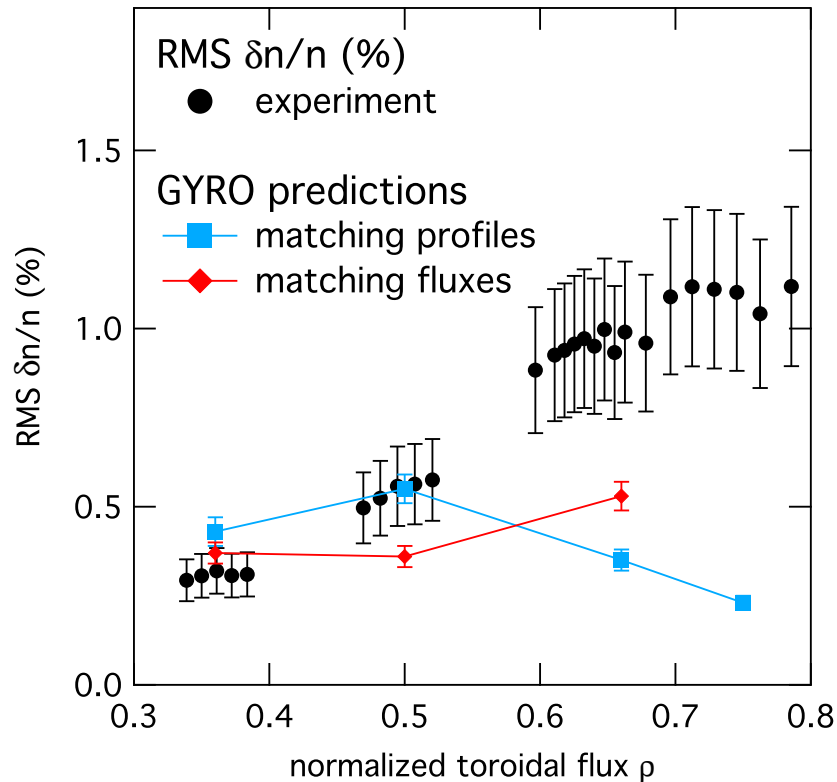


DIII-D 128913 1300-1700 ms

# Measured Versus Simulated Fluctuations (DIII-D 128913 1500n)

Underprediction previously reported by White, Holland APS07

## Fluctuation level sensitive to instability threshold



DIII-D 128913-23 1300-1700 ms

**Red diamonds: MATCH POWER**    **Blue squares: MATCH PROFILES**

# Summary

1. TGYRO solves **Fokker-Planck hierarchy** through **4 orders in  $\rho_*$** .
2. We have implemented a **novel approach** to solve the transport problem:
  - Alternative to fixed-gradient gyrokinetic/neoclassical simulations.
  - New **integral method** for solving transport equations.
3. Results can be routinely analyzed with **synthetic diagnostic** tools:
  - Critical for in-depth code **validation**.
4. Demonstrated **application of TGYRO** to DIII-D discharge:
  - Observe improved agreement with experimental trends.
  - Perhaps **multi-scale ITG-ETG simulations** are needed in pedestal region.
5. Coming soon: **ITER** core performance projections:
  - Computational cost roughly the same as for DIII-D.

# Extra Slides

# Fokker-Planck Equation with COLLISIONS

Complete theory derived by Sugama and coworkers

1. H. Sugama and W. Horton, **Neoclassical and anomalous transport in axisymmetric toroidal plasmas with electrostatic turbulence**, Phys. Plasmas **2** (1995) 2989.
2. H. Sugama, M. Okamoto, W. Horton and M. Wakatani, **Transport processes and entropy production in toroidal plasmas with gyrokinetic electromagnetic turbulence**, Phys. Plasmas **3** (1996) 2379.
3. H. Sugama and W. Horton, **Transport processes and entropy production in toroidally rotating plasmas with electrostatic turbulence**, Phys. Plasmas **4** (1997) 405.
4. H. Sugama and W. Horton, **Neoclassical electron and ion transport in toroidally rotating plasmas**, Phys. Plasmas **4** (1997) 2215.
5. H. Sugama and W. Horton, **Nonlinear electromagnetic gyrokinetic equation for plasmas with large mean flows**, Phys. Plasmas **5** (1998) 2560.

# Fokker-Planck Transport Theory

## Transport of Density on $\mathcal{O}(\rho_*^2)$ timescale

$$\begin{aligned} & \Gamma_a^{\text{GV}} && \text{Gyroviscous particle flux} \\ \Gamma_a^{\text{neo}} &= \left\langle \int d^3v \bar{g}_a W_{a1} \right\rangle && \text{Neoclassical particle flux} \\ \Gamma_a^{\text{tur}} &= \left\langle \left\langle \int d^3v \hat{h}_a \hat{W}_{a1} \right\rangle \right\rangle && \text{Turbulent particle flux density} \\ & S_{n,a}^{\text{beam}} && \text{Beam density source rate} \\ & S_{n,a}^{\text{wall}} && \text{Wall density source rate} \end{aligned}$$

# Fokker-Planck Transport Theory

## Transport of Energy on $\mathcal{O}(\rho_*^2)$ Timescale

	$Q_a^{\text{GV}}$	Gyroviscous energy flux
$Q_a^{\text{neo}} = \left\langle \int d^3v \bar{g}_a \frac{\varepsilon}{T_a} W_{a1} \right\rangle$		Neoclassical energy flux
$Q_a^{\text{tur}} = \left\langle \left\langle \int d^3v \hat{h}_a \frac{\varepsilon}{T_a} \hat{W}_{a1} \right\rangle \right\rangle$		Turbulent energy flux
	$S_{W,a}^{\text{aux}}$	Auxiliary heating power density
	$S_{W,a}^{\text{rad}}$	Radiation heating power density
	$S_{W,a}^{\alpha}$	Alpha heating power density
$S_{W,a}^{\text{tur}} = \left\langle \left\langle \int d^3v \hat{h}_a W_{aT} \right\rangle \right\rangle$		Turbulent exchange power density
	$S_{W,a}^{\text{col}}$	Collisional exchange power density

# Fokker-Planck Transport Theory

## Transport of Toroidal Angular Momentum on $\mathcal{O}(\rho_*^2)$ Timescale

$$\begin{aligned} \Pi_a^{\text{GV}} & \text{ Gyroviscous angular momentum flux} \\ \Pi_a^{\text{neo}} &= \left\langle \int d^3v \bar{g}_a W_{aV} \right\rangle \text{ Neoclassical angular momentum flux} \\ \Pi_a^{\text{tur}} &= \left\langle \left\langle \int d^3v \hat{h}_a \hat{W}_{aV} \right\rangle \right\rangle \text{ Turbulent angular momentum flux} \\ S_{\omega,a} & \text{ Angular momentum density source rate} \end{aligned}$$