Cross-Machine Comparison of Resonant Field Amplification and Resistive Wall Mode Stabilization by Plasma Rotation

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In collaboration with

DIII-D team including

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JET-EFDA contributors including

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NSTX team including

S.A. Sabbagh, J.M. Bialek, J.E. Menard, A.C. Sontag, W. Zhu, D.A. Gates, D. Mueller and R. Raman





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Cross-machine comparison establishes universality of resistive wall mode stabilization by plasma rotation

- Resistive wall mode (RWM) stabilization by plasma rotation has been studied up to high β_N
 - **NSTX**: $\beta_N \leq 7.3$
 - **DIII-D**: $\beta_N \leq 4.3$
 - JET: $\beta_N \leq 3.7$
- Recent DIII-D experiments demonstrate sustained $\beta_N \sim 4$ with active error field correction
 - → A.M. Garofalo invited talk, UI2, Friday 9:30AM
- The devices vary in size and aspect ratio A=R/a







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Plasma rotation can stabilize RWM up to the ideal wall stability limit

• Resistive Wall Mode (RWM):

- External kink mode whose growth is slowed by magnetic field penetration through the conducting wall
- Quasi-static perturbation in a fast toroidal plasma flow: γ_{RWM} , $\omega_{RWM} \sim \tau_W^{-1} << \Omega_{rot}$
- Stabilizing effect of plasma rotation first observed in DIII-D [E.J. Strait et al, PRL 74 (1995)]
 - Stabilization requires a dissipation mechanism [A. Bondeson, D.J. Ward, PRL 72 (1994)]

Resonant field amplification (RFA):

 Externally applied resonant fields can excite the weakly damped RWM [A.H. Boozer, PRL 86 (2001)]





Machine-size comparison between DIII-D and JET and aspect ratio comparison between DIII-D and NSTX

• Machine size comparison: DIII-D and JET vary by a factor of 1.7

- Same resonant field amplification (RFA)
- Same critical plasma rotation for RWM stabilization
 - Importance of q=2 surface for rotational stabilization
- Aspect ratio comparison: DIII-D and NSTX vary by a factor of 2
 - Higher critical rotation at low aspect ratio explained by trapped particles not contributing to RWM stabilization
 - Alternatively, the RWM stabilization is determined by the sound wave velocity rather than the Alfvén velocity
- Target plasmas designed for a large difference between no wall and ideal wall β limits rather than maximum β_{N}





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Match parameters for the RWM drive and for the dissipation mechanism

- Obtain the same external kink mode by matching the ideal MHD properties of the plasma: shape, q-profile, pressure profile
- Express RWM drive by the normalized gain in $\boldsymbol{\beta}$

$$C_{\beta} = \frac{\beta - \beta_{no \ wall}}{\beta_{ideal \ wall} - \beta_{no \ wall}} = \begin{cases} 0 \text{ at no wall limit} \\ 1 \text{ at ideal wall limit} \end{cases}$$

- Stabilization models:
 - Sound wave damping: Fluid approximation, where RWM couples to sound waves, which are then ion Landau damped, described via a parallel viscous force
 - Kinetic description of inertia enhancement and ion Landau damping
- → Normalize plasma rotation frequency with inverse of Alfvén time

$$\tau_A = R_0 \frac{\sqrt{\mu_0 n_e m_i}}{B_0}$$



Matching shape and profiles leads to same ideal MHD no-wall stability limit in DIII-D and JET plasmas





• Ideal MHD no-wall stability limit:



 $-\,\beta_{\text{N,no wall}}\,\text{~2.8}\,\,\boldsymbol{\ell}_{i} \quad \text{ in DIII-D and JET}$



Wall-stabilized regime in DIII-D and JET varies due to different ℓ_i and wall geometry







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DIII-D and NSTX develop common target with a substantial wall-stabilized regime



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DIII-D and JET plasmas probed using non-axisymmetric external control coils with similar geometry

- Apply resonant field pulses B^{ext} with one pair of external control coils (predominantly n=1)
- Detect plasma response B^{plas} with toroidal arrays of B_r sensors









High- β plasmas respond to externally applied n=1 perturbations

- Beta exceeds nowall limit
- Plasma rotation provides RWM stabilization
- Probe plasma with externally applied n=1 field
- RFA leads to plasma response detected at the (toroidal) node of the applied field





RFA in DIII-D and JET increase significantly once β exceeds no wall stability limit

- Increase of RFA for β > β_{no-wall} consistent with previous observations in DIII-D and NSTX [A.C. Sontag et al, Phys. Plasmas (2005)]
 - Low β response in JET differs from DIII-D

 Measured amplification in DIII-D twice as large as in JET



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Comparison of RFA measurements has to account for geometry of magnetic fields and sensor locations

- Similar geometry of applied fields in DIII-D and JET
- Similar geometry of plasma perturbation in DIII-D and JET
- Radial decay of external field and plasma response cause radial dependence of B^{plas}/B^{ext}
 - Cylindrical approximation

$$\frac{B_a^{plas}/B_a^{ext}}{B_s^{plas}/B_s^{ext}} = (r_s/a)^{2m}$$

 Assume effective poloidal mode number m=2 at outboard midplane



	DIII-D	JET
a (m)	0.54	0.95
r _s (m)	0.74	1.61
(r _s /a) ⁴	3.5	8.25

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RFA magnitudes at the plasma boundary in DIII-D and JET are in quantitative agreement

- Map RFA to the same location, e.g. plasma boundary
- RWM drive described by normalized gain over no wall limit C_{β}





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Braking of plasma rotation needed for RWM onset

- NBI torque in NSTX, DIII-D and JET is usually provides sufficient rotation for RWM stabilization
- Increase drag by applying non-axisymmetric fields
 - Neoclassical toroidal viscosity (NTV) [Shaing, Phys. Plasmas 10, 1443 (2003)]:

$$T_{NTV} \propto \left(\delta B / B \right)^2$$

- Non-linear RWM onset
 - Magnetic braking:
 - RWM dispersion relation:
 - Evolution of perturbed field (RFA and RWM):

$$\frac{d}{dt}\Omega_{rot} \propto \delta B^2$$

$$\gamma_{RWM} \tau_W = f(\Omega_{rot})$$

$$\tau_W \left(\frac{d}{dt} \delta B - \gamma_{RWM} \delta B \right) = \delta B^{ext}$$



Magnetic braking leads to non-linear RWM onset

• Apply n=1 field

Plasma rotation decreases

 Onset of fast growth is preceded by increasing RFA

- Critical rotation Ω_{crit} measured at onset of fast growth







Time of marginal stability determined from start of fast mode growth





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Critical rotation in the center of DIII-D and JET plasmas decreases with increasing q₉₅







Evaluating Ω_{crit} at q=2 removes q_{95} -dependence and leads to quantitative agreement between DIII-D and JET



- q₉₅-dependence caused by q-surfaces moving inwards towards higher rotation
 - Stabilization mechanism depends on local q, e.g. kinetic damping [A. Bondeson and M.S. Chu, Phys. Plasmas (1996)],

 $\Omega_{\it crit} \propto 1 / q^2$

- Quantitative agreement in DIII-D and JET indicates prominent role of q=2 surface
 - Consistent with predictions for sound wave damping
 [D. Gregoratto, et al, Plasma Phys. Control. Fusion (2001)]
- Variations of JET data partially caused by β-dependence





Ω_{crit} has weak β -dependence

- Beta-dependence of $\Omega_{\rm crit}$ can account for some of the scatter of the measurement
- Consistent with increase of RFA with β







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- Beta-dependence of Ω_{crit} can account for some of the scatter of the measurement
- Consistent with increase of RFA with β
- Value of Ω_{crit} consistent with predictions (using MARS-F) for
 - Sound wave damping







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- Value of Ω_{crit} consistent with predictions (using MARS-F) for
 - Sound wave damping
 - Kinetic damping







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Critical rotation in NSTX exceeds critical rotation in DIII-D

 q₉₅-dependence in NSTX similar to DIII-D and JET

- Evaluate Ω_{crit} at the same value of q

- NSTX critical rotation $\Omega_{crit} \tau_A$ at same q always equal or higher
 - Rotation for $q \ge 3$ close to zero
 - Single resonant surface can be sufficient for RWM stabilization





Aspect ratio dependence suggests that trapped particles do not contribute to RWM stabilization

- Ion Landau damping significantly reduced for trapped particles
 - Assume that only passing particles contribute to RWM stabilization

$$\Omega_{crit}\tau_A \propto \frac{1}{1-\sqrt{\epsilon}}$$

- Observed doubling of $\Omega_{crit} \tau_A$ consistent with stabilization by passing particles only
 - Effect included in the kinetic but not in the sound wave damping model







Alternatively - RWM damping could be determined by sound wave velocity

 Coupling to sound waves depends on sound time

$$\tau_{S} = R_0 \sqrt{\frac{m_i}{k_B T_e + k_B T_i}}$$

- Alfvén time and sound time linked via β
 - $\frac{\tau_{A}}{\tau_{S}} \propto \beta^{1/2} \propto \left(\epsilon \beta_{N} / q_{95}\right)^{1/2}$
 - Link is broken by aspect ratio
- Normalization on sound time removes aspect ratio dependence







Comparison of NSTX, DIII-D and JET establishes universality of RWM stabilization by plasma rotation

- q_{95} -dependence of $\Omega_{crit}\tau_A$ explained by re-location of q-surfaces
 - RWM stabilization depends on the local q
- Quantitative agreement of $\Omega_{crit}\tau_A$ evaluated at q=2 in DIII-D and JET
 - Physics determined by ideal MHD drive and normalized rotation
 - q=2 surface plays prominent role in stabilization mechanism
- Quantitative agreement of RFA in DIII-D and JET
 - Increase of RFA above $\beta_{\text{no wall}}$ in qualitative agreement with NSTX
 - RFA is manifestation of a weakly damped RWM
- Aspect ratio dependence of $\Omega_{crit}\tau_A$ in DIII-D and NSTX explained by trapped particles not contributing to RWM stabilization
 - Alternatively, the stabilization is determined by the sound wave velocity rather than the Alfvén wave velocity







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