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Understanding plasma transport in RMP using XGC0 particle code

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In collaboration with

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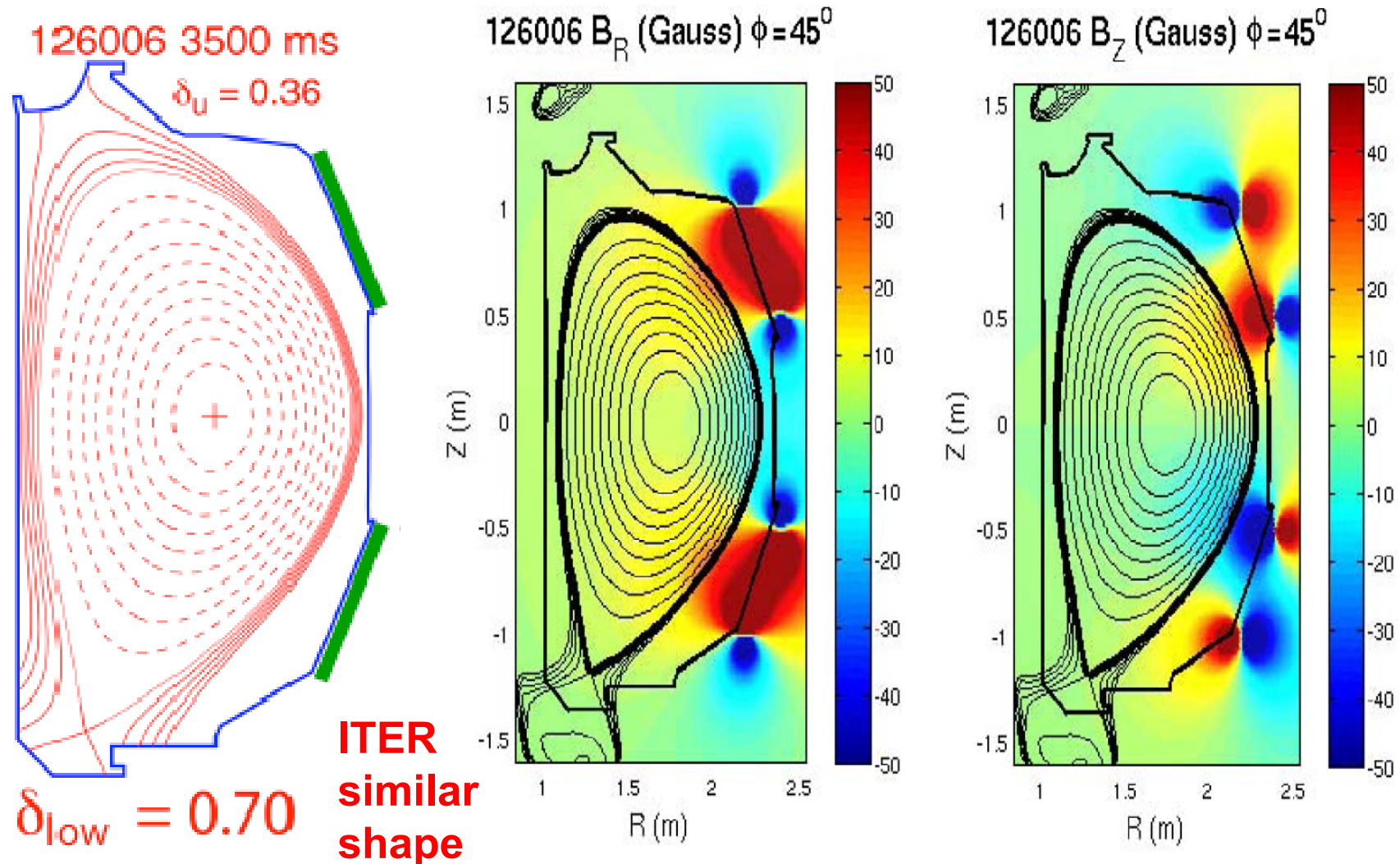
Acknowledgement to DIII-D RMP team

SciDAC FSP Prototype Center for Plasma Edge Simulation

Observed phenomena on DIII-D

- **Burrell, et al, PPCF 47, B37(05); Evans, et al, Nature Physics 2, 419 (06)**
- $q=m/n=11/3, 12/3, 13/3$ within pedestal region and $\delta B/B \sim 3 \times 10^{-4}$
- ITER relevant collisionality: electron collisionality ~ 0.1
- Density collapses (below $2 \times 10^{19} \text{ m}^{-3}$), but not the temperatures
- Toroidal rotation increases promptly by RMP
- T_i increases significantly, T_e steepens just inside the separatrix with little change of ∇P there.
- E_r well persists at narrow radial width just inside the separatrix

B-quilibrium (126006) and vacuum RMP profiles (from I. Joseph)



Simulation condition

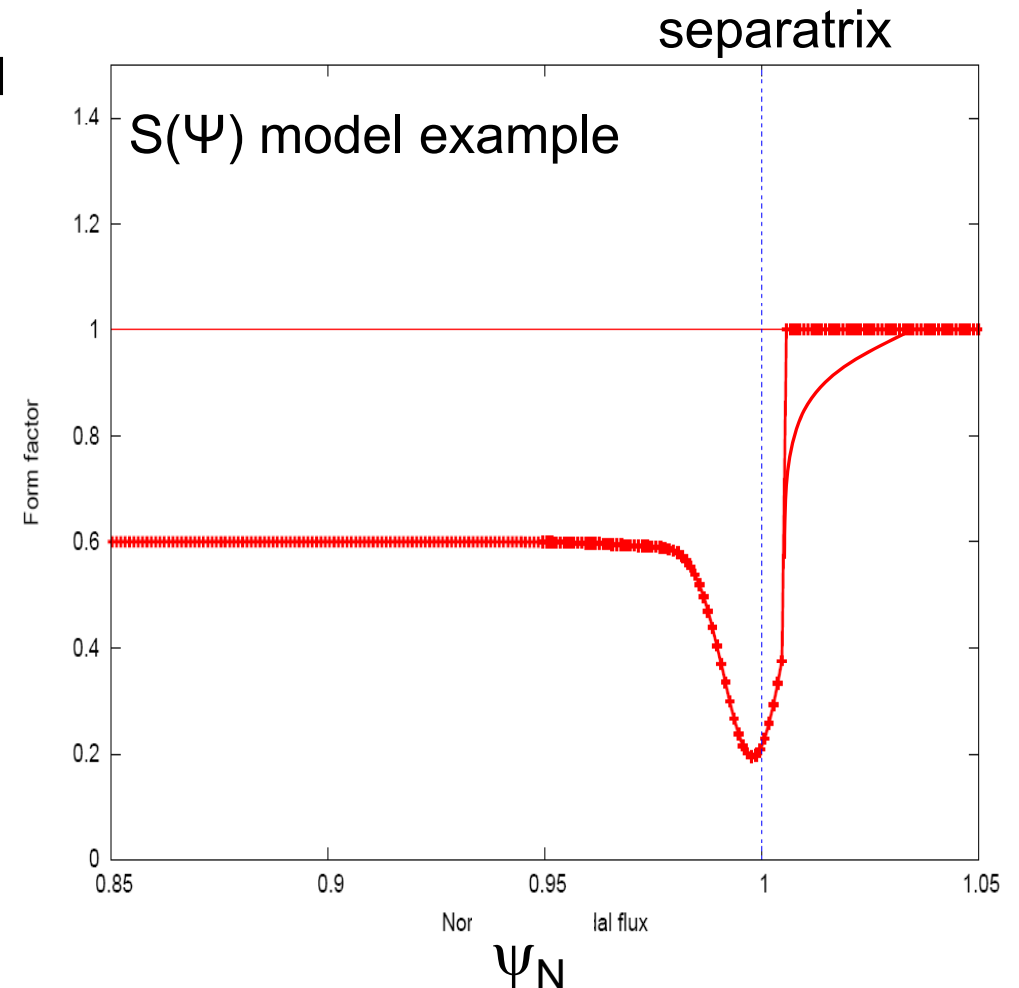
- Use reduced m_i/m_e ($=1000$) for time saving ($\sim 5\%$ error)
- We use both vacuum RMP and a screened RMP model
- RMP is turned on after a neoclassical quasi-steady equilibrium is established (~ 10 ion toroidal transit time).
- Total simulation time ~ 60 ion toroidal transits (~ 5 ms in physical time)
- The goal is to compare the plasma before and after the RMP and to understand the underlying transport physics
- Neutral recycling rate = 0.9 (active pumping)
- Heat flux from core: 3.5 MW to ions and 3.5 MW to electrons
- A random walk anomalous radial diffusion of $D=0.1\text{m}^2/\text{sec}$ is superimposed on the Lagrangian particle guiding center motion. Later on, we found $D=0.25\text{m}^2/\text{s}$ is more appropriate (from experimental comparison).

Plasma shielding effect can be examined by multiplying a model screening form factor $S(\Psi)$

$$\delta\mathbf{B} = \nabla \times (S(\psi) \delta\mathbf{A}_{\text{vac}})$$

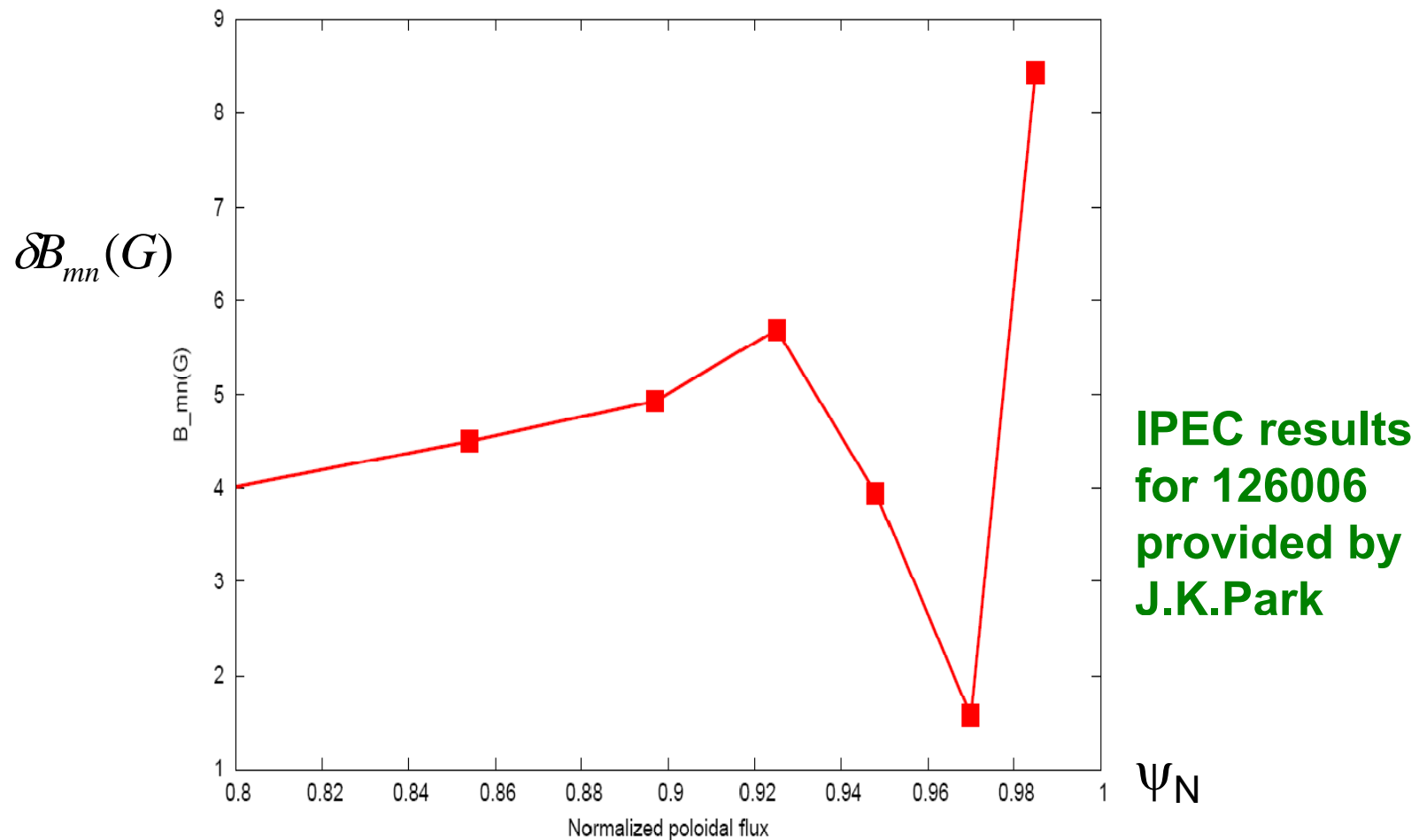
[Fitzpatrick (98), J. Park, et al., and Strauss, et al, IAEA2008]

- Strong plasma rotation can screen RMP penetration
- Plasma can also amplify RMP by resonant interaction



Example: IPEC ideal MHD simulation shows RMP screening with a local minimum inside separatrix

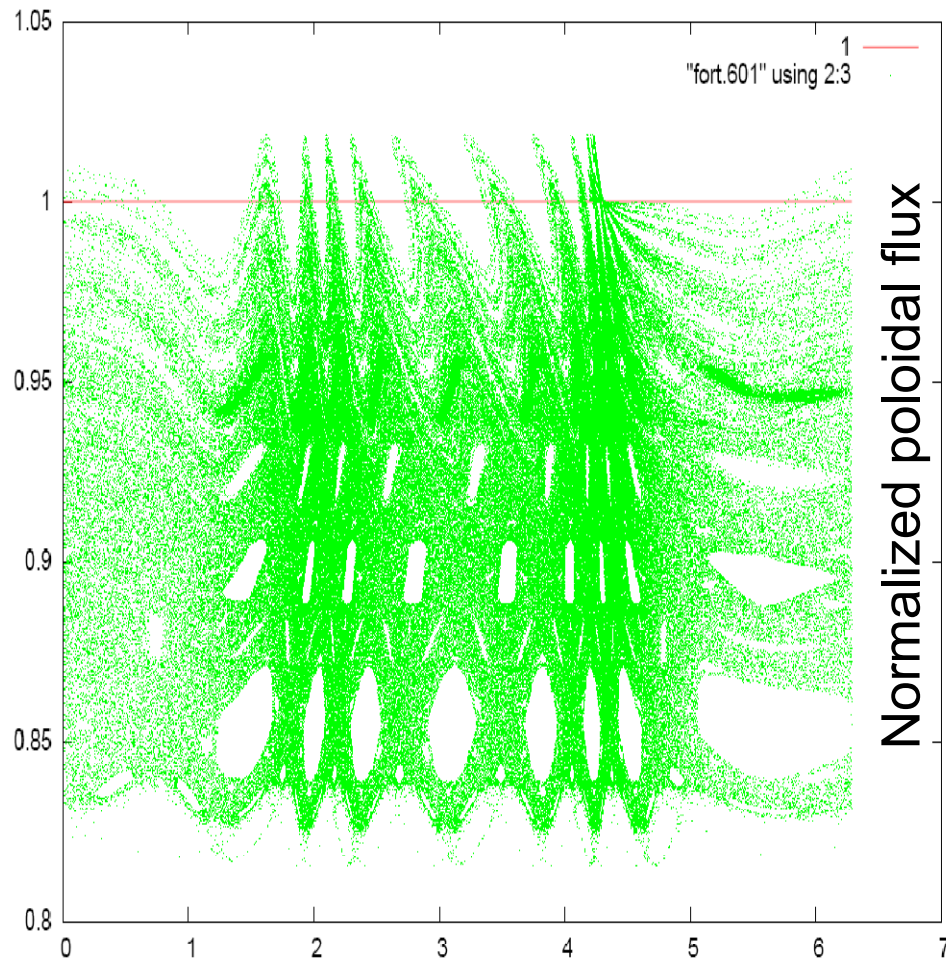
IPEC: ideally perturbed equilibrium code, **J.K.Park, PoP 14, 052110 (2007)**



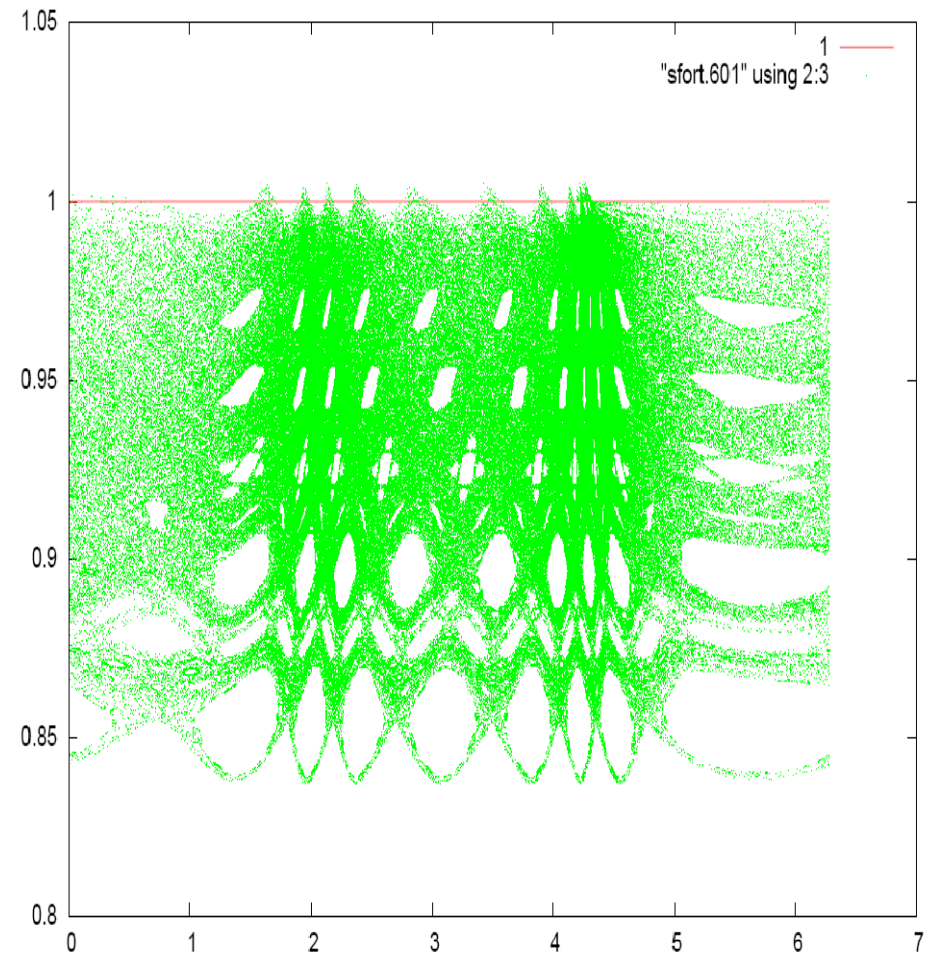
Radial magnetic field components here is in Hamada coordinates

Poincare plots for vacuum and screened RMP fields

Vacuum RMP



Screened RMP (analytic model)



Poloidal angle(radian)

Theoretical quasilinear electron heat flux in a collisionless limit

Rechester-Rosenbluth PRL (1978)

- Heat transport along the stochastic magnetic field assuming Maxwellian distribution function for electron ensemble

$$\Gamma_r = -\frac{1}{\sqrt{\pi}} D_m V_{the} \left(\frac{1}{n} \frac{\partial n}{\partial r} + \frac{1}{2T_e} \frac{\partial T_e}{\partial r} + \frac{eE_r}{T_e} \right) n,$$

$$Q_r = -\frac{2}{\sqrt{\pi}} D_m V_{the} \left(\frac{1}{n} \frac{\partial n}{\partial r} + \frac{3}{2T_e} \frac{\partial T_e}{\partial r} + \frac{eE_r}{T_e} \right) n T_e,$$

where quasilinear expression for field line diffusion coefficient D_m is given by

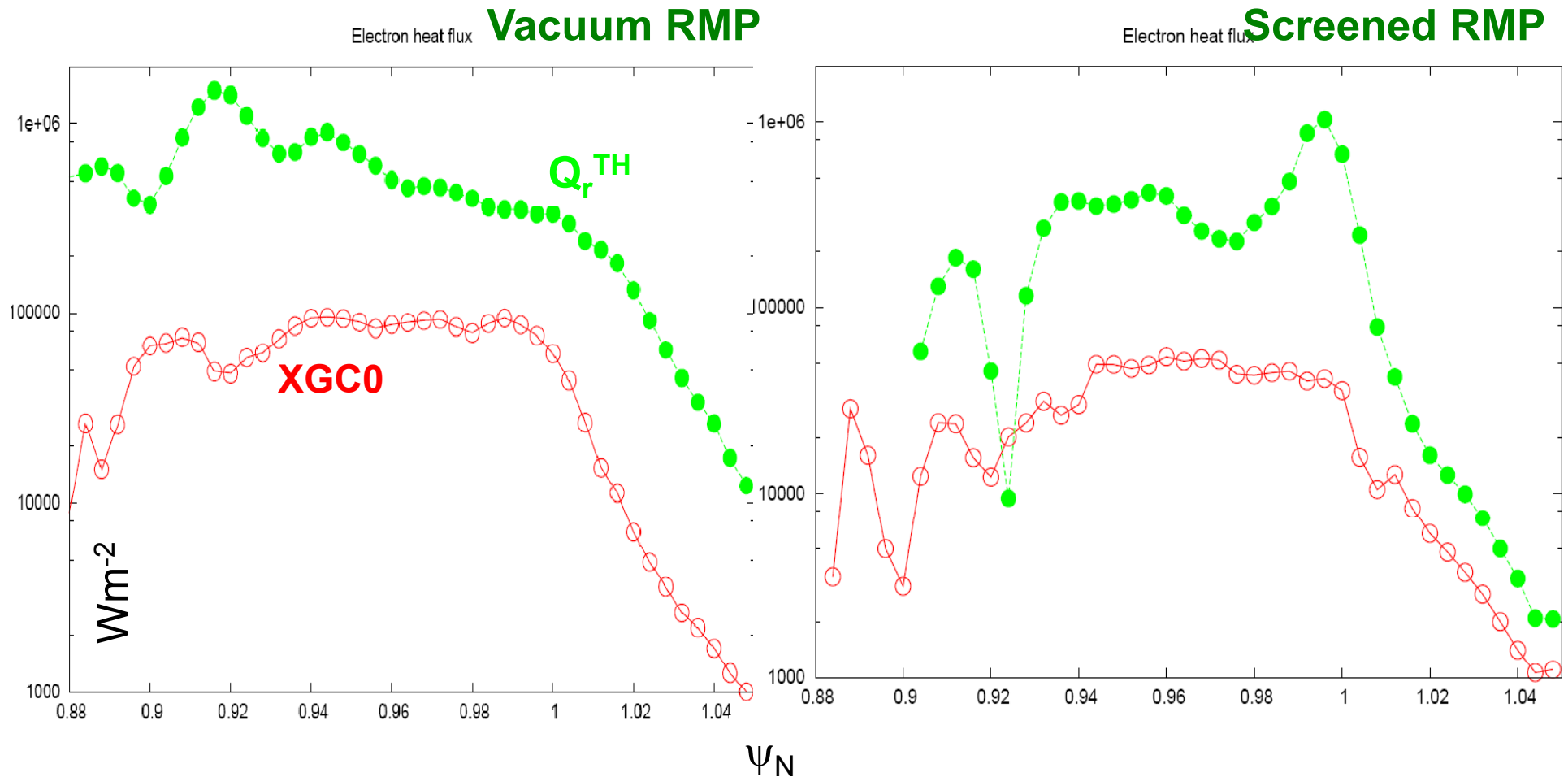
$$D_m = \pi q R_0 |\delta B / B|^2 \quad \text{which is evaluated at each resonance surface.}$$

- Setting Γ_r to zero gives theoretical ambipolar radial electric field and the corresponding heat flux expression

$$E_r^{TH} = -\frac{T_e}{e} \left[\frac{1}{n} \frac{\partial n}{\partial r} + \frac{1}{2T} \frac{\partial T}{\partial r} \right],$$

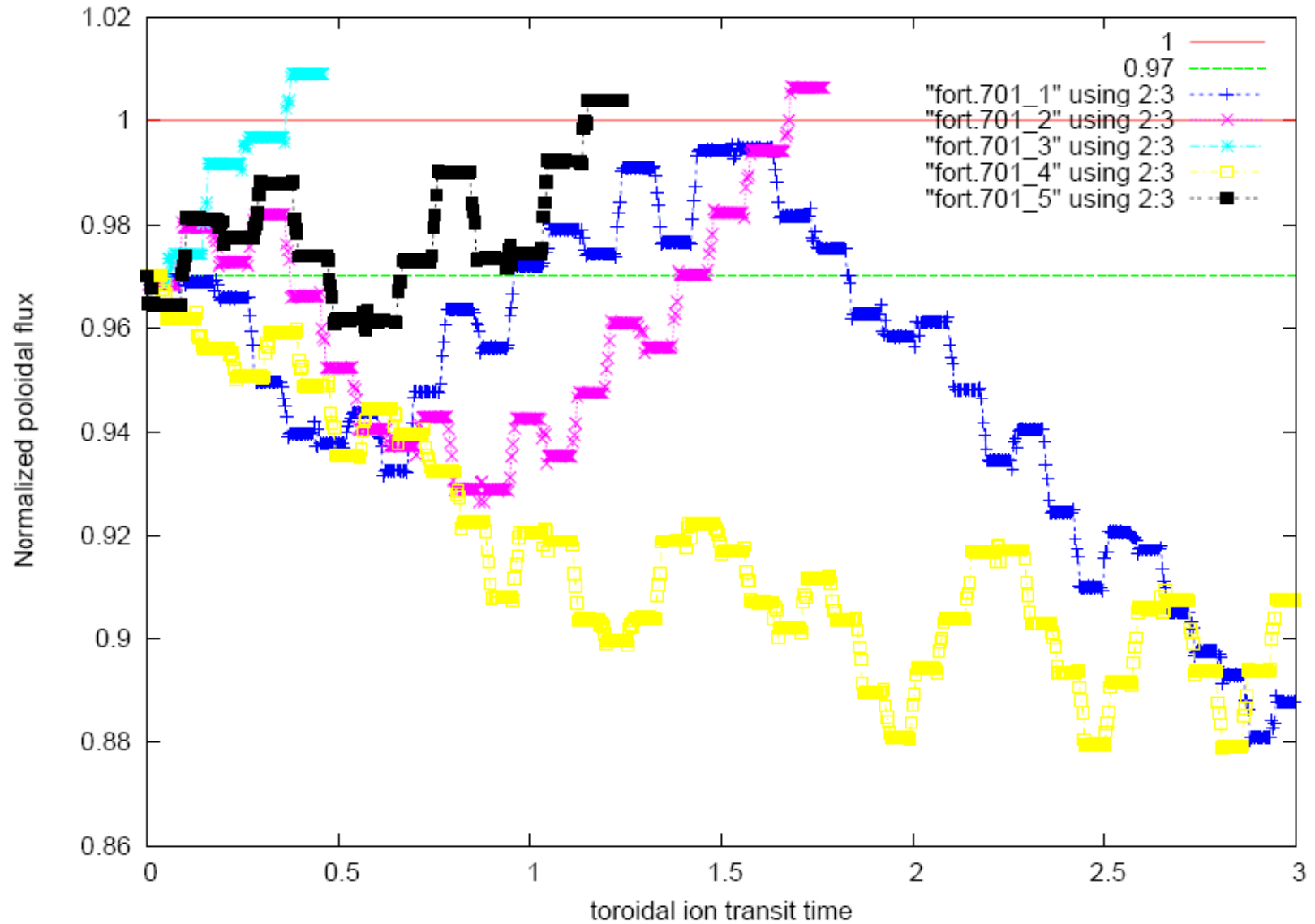
$$Q_r^{TH} = -\frac{2}{\sqrt{\pi}} D_m V_{the} n \frac{\partial T}{\partial r}$$

XGC0 electron heat flux is much smaller than Collisionless Rechester-Rosenbluth



Actual heat flux is much different from Rechester-Rosenbluth formula due to several nonideal effects such as incomplete chaos, E_r , mirror trapping, etc.

Single electron motions in Vacuum RMP: Far from Rechester-Rosenbluth type

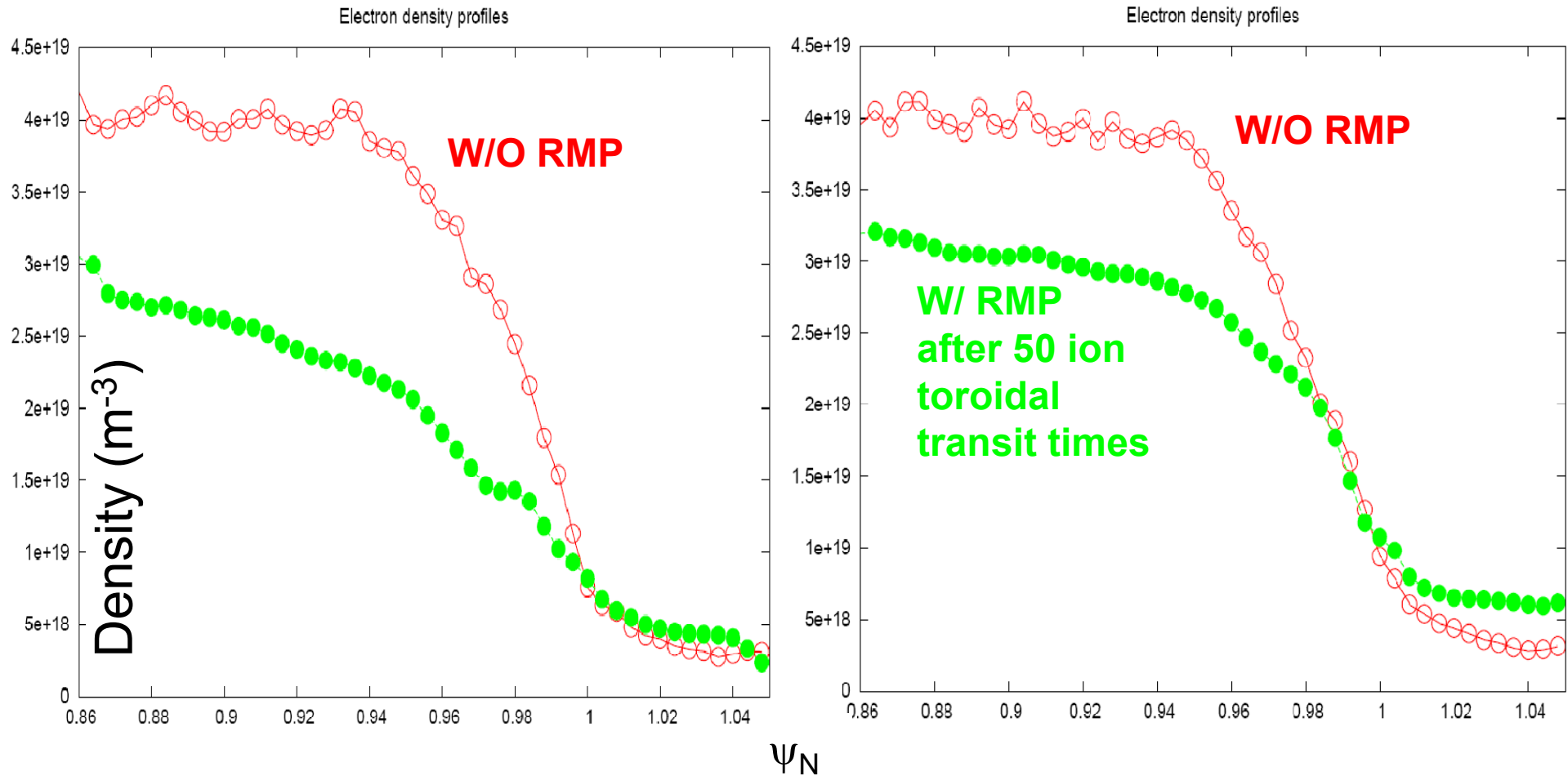


Density is reduced by RMPs

Recycling coefficient=0.9

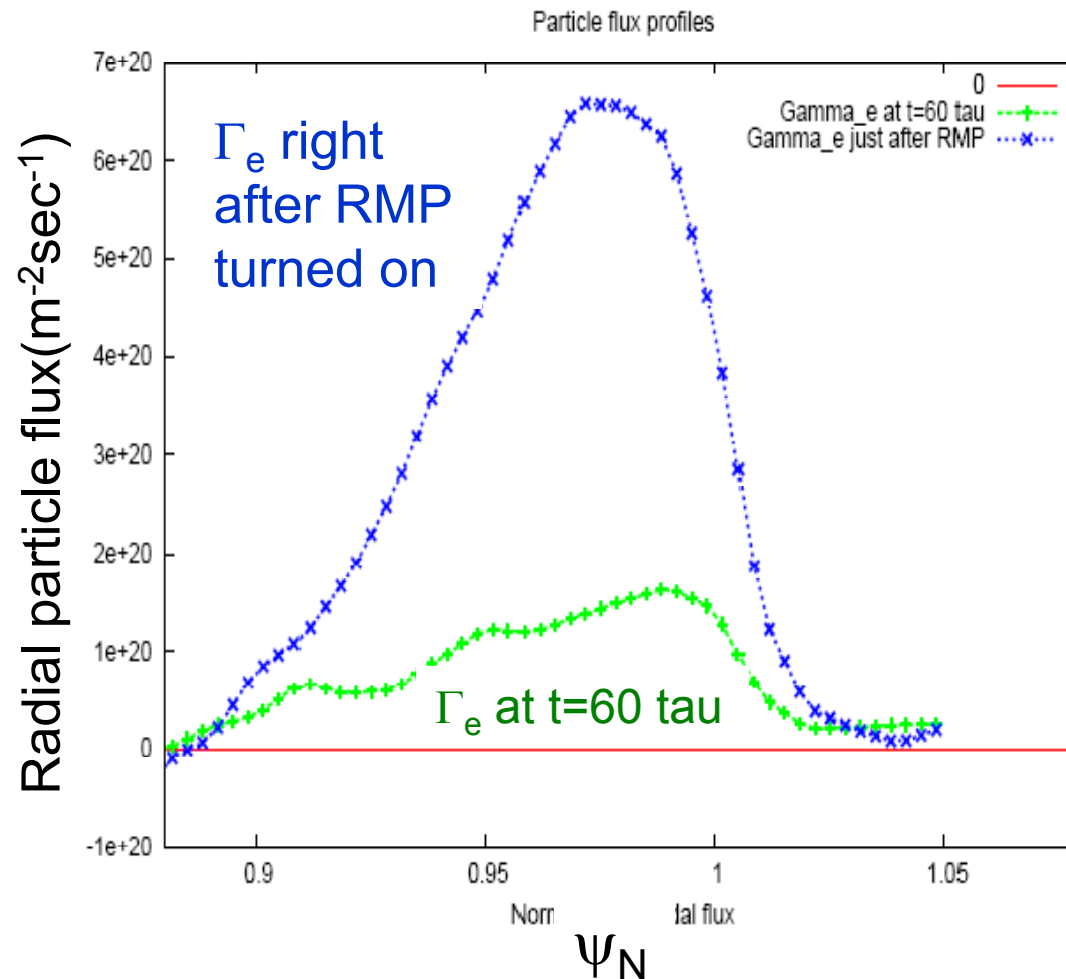
Vacuum RMP

Screened RMP



- As soon as RMP is turned on, density drops promptly

Radial particle flux after RMP

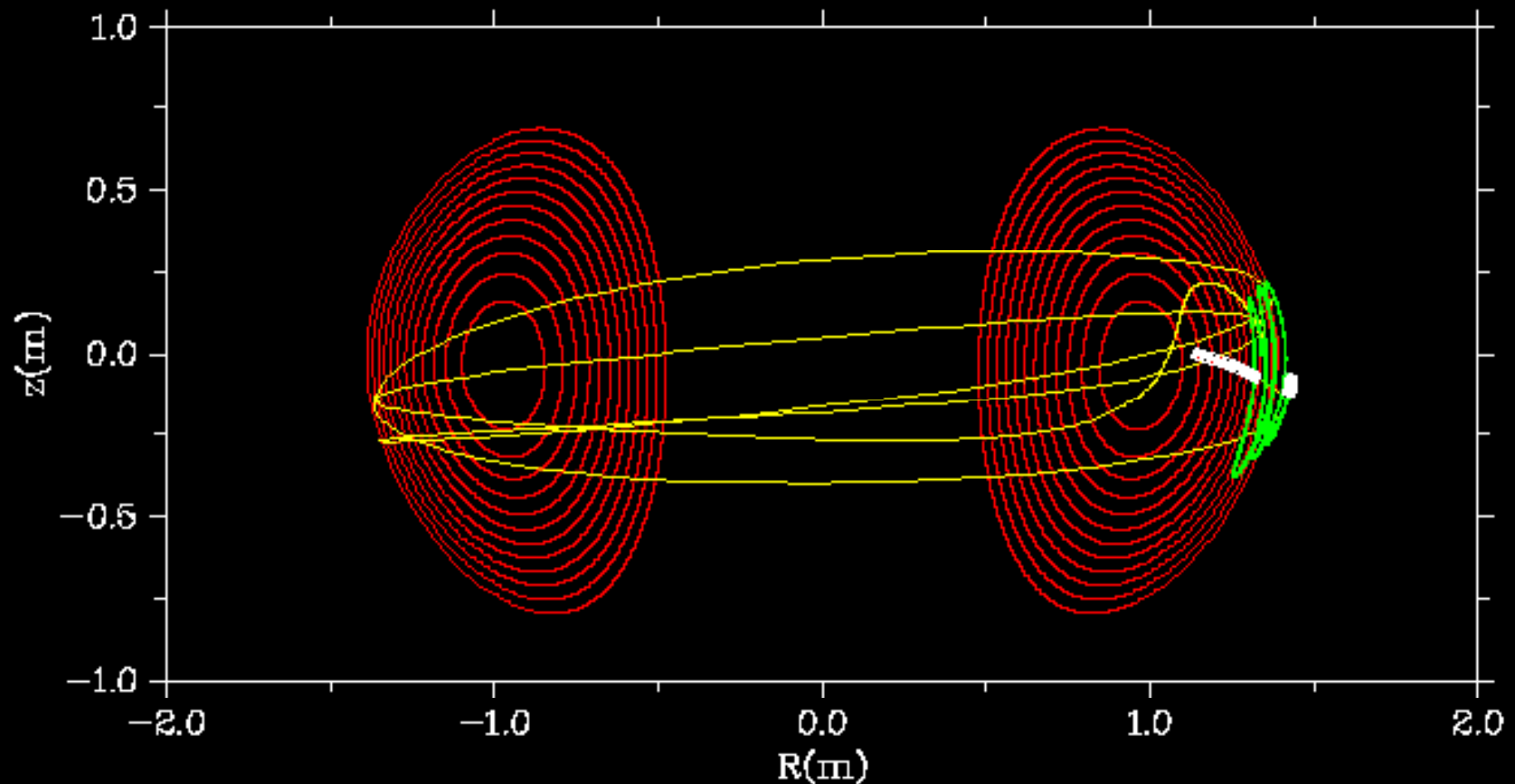


Radial particle flux is maximal right after RMP application, and then decays to quasi-steady state. E_r adjusts itself to achieve $\Gamma_i(E_{\text{RMP}}) = \Gamma_e(E_{\text{RMP}})$ for $E_{\text{RMP}} \neq E_{\text{neo}}$

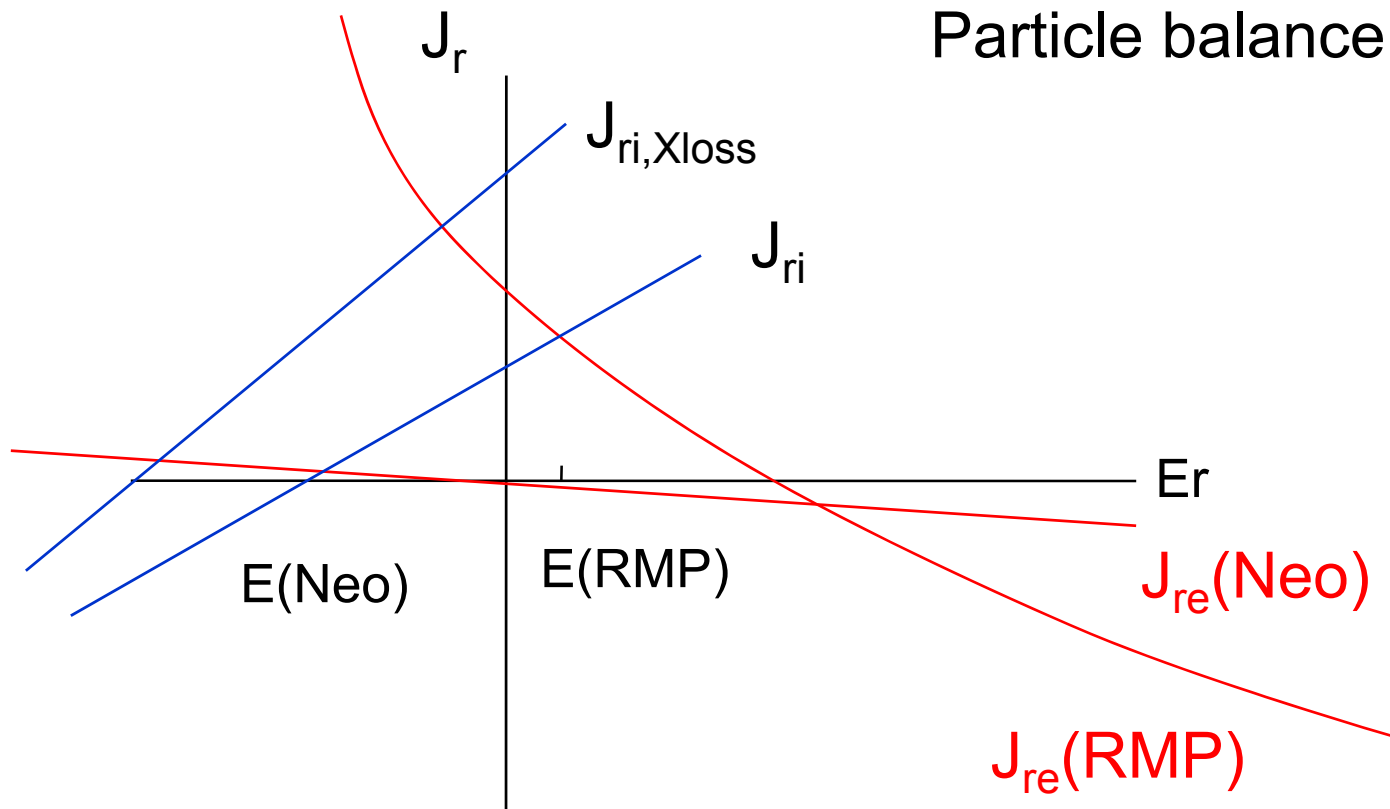
Initial ion flux is from strong polarization drift

[Baek, Ku, Chang, PoP, 13, 012503 (2006)]

NSTX:Neoclassical Polarization Drift(Outward), 10keV



S.H.Hahn, Aug 2003

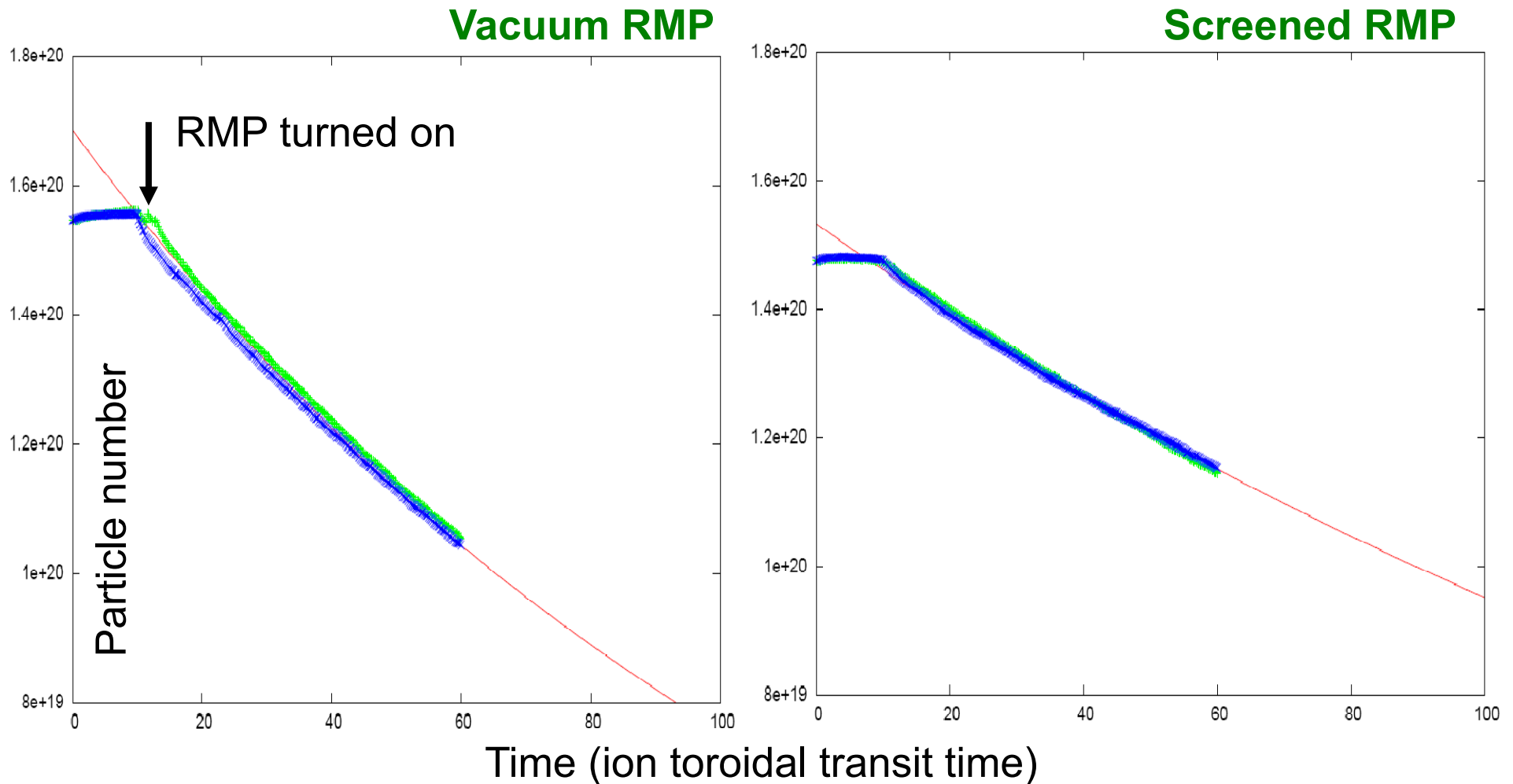


Neoclassical transport is automatically ambipolar only in neoclassical equilibrium \Rightarrow strong radial conductivity.

$E(Neo)$

$$u_{i\parallel} = (cT_i/eB_p) [k d \log T_i / dr - d \log p_i / dr - (e/T_i) d\phi / dr]$$

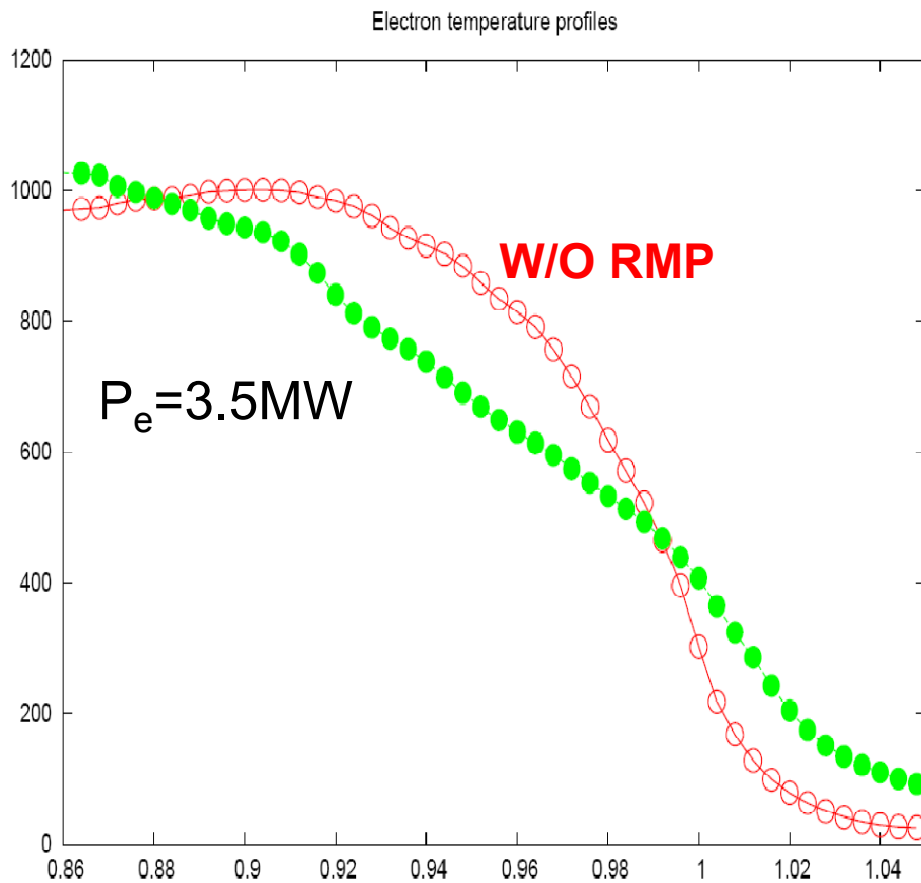
Temporal changes of particle content inside the separatrix gives density e-folding time



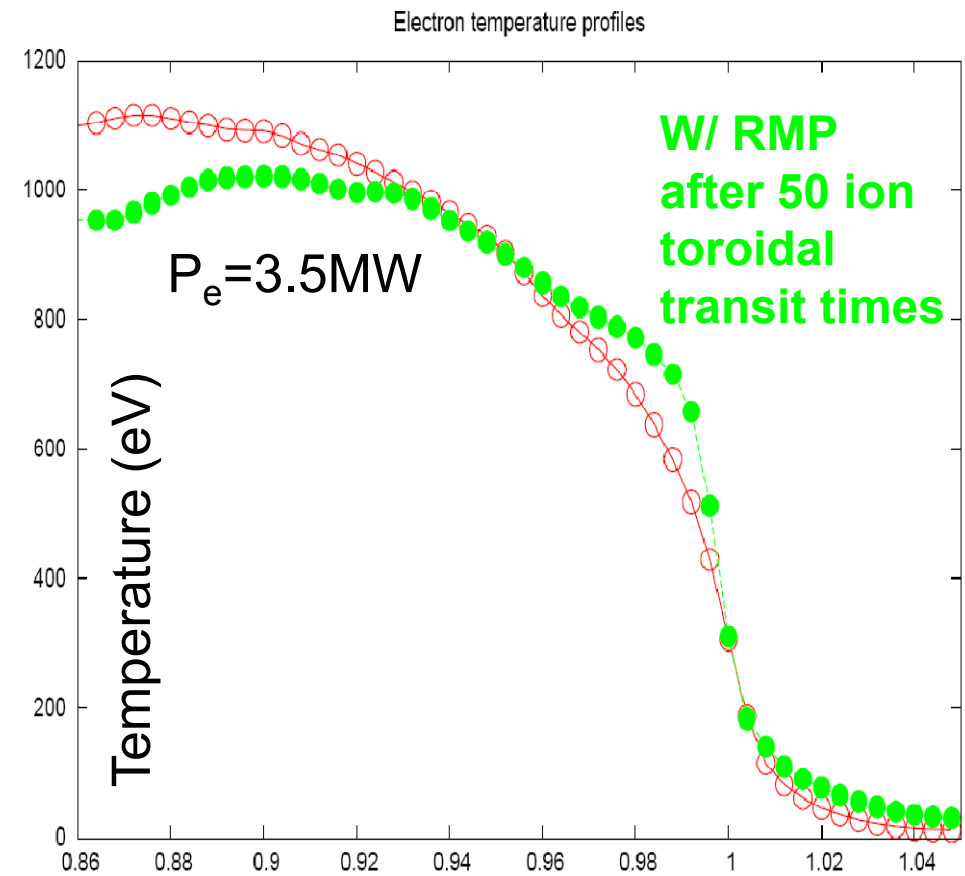
- Density e-folding time is about 10 ms for vacuum RMP and 20 ms for screened RMP.

Electron temperature profile steepens near the separatrix under screened RMP

Vacuum RMP



Screened RMP

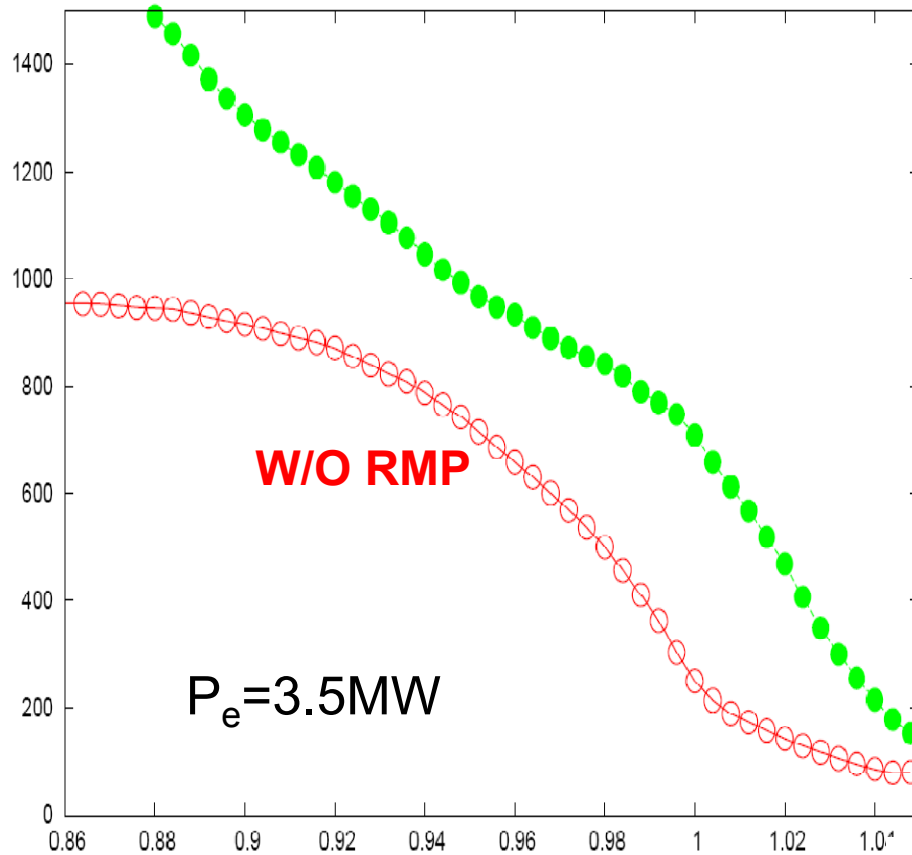


Ψ_N

Ion temperature continues to rise, fueled by strong heat flow from core (ion thermal transport is still low under RMP, and the density is lower)

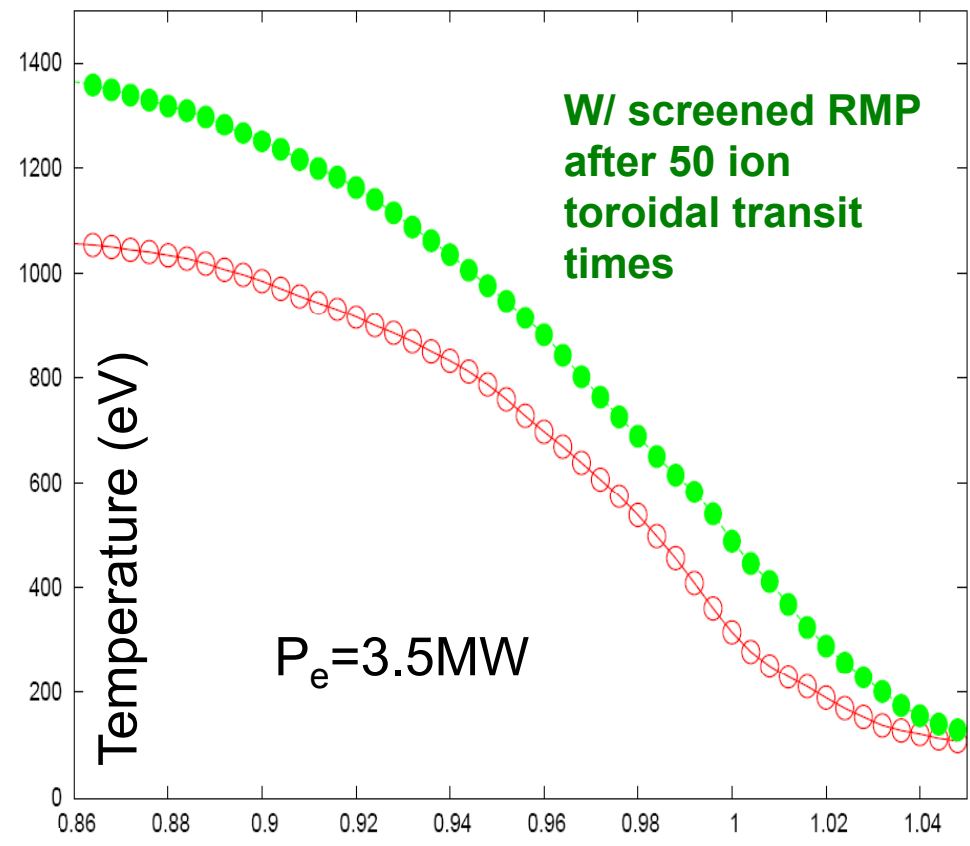
Vacuum RMP

Ion temperature profiles



Screened RMP

Ion temperature profiles

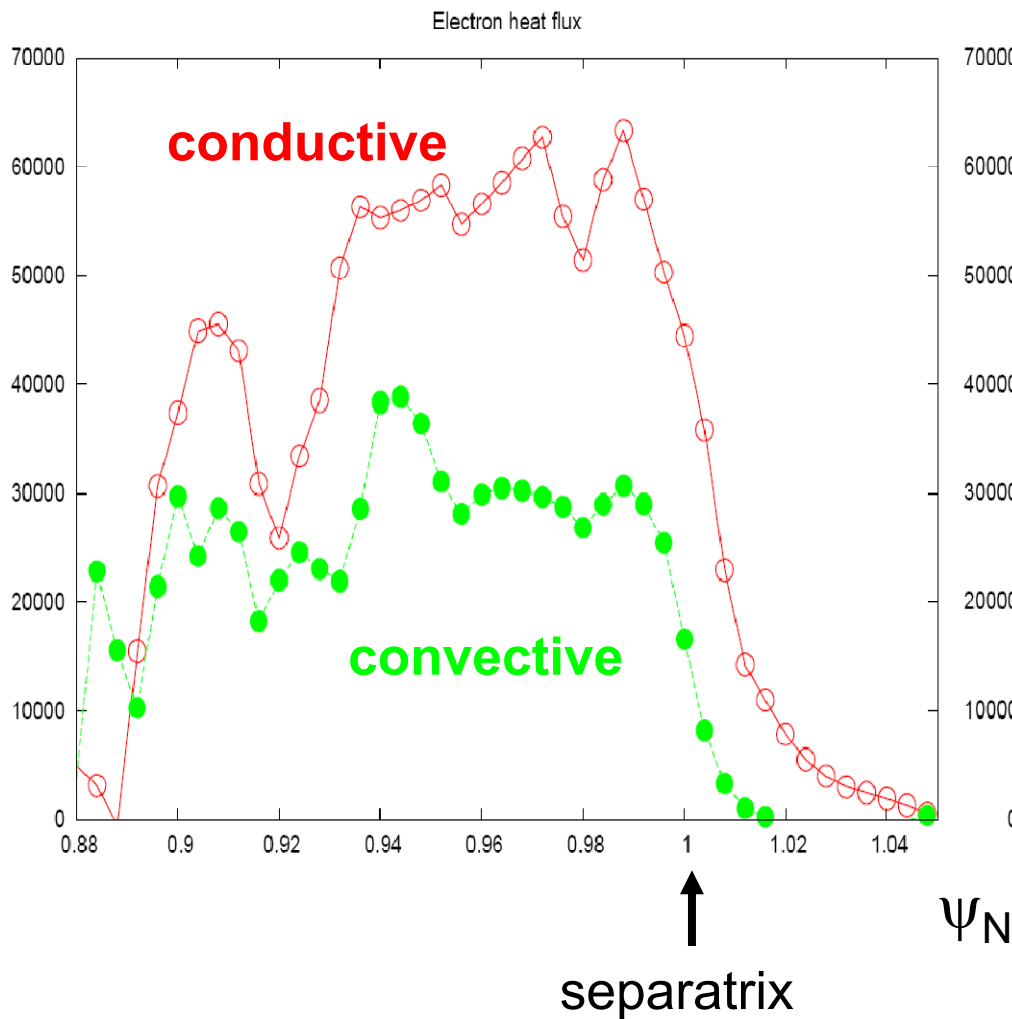


ψ_N

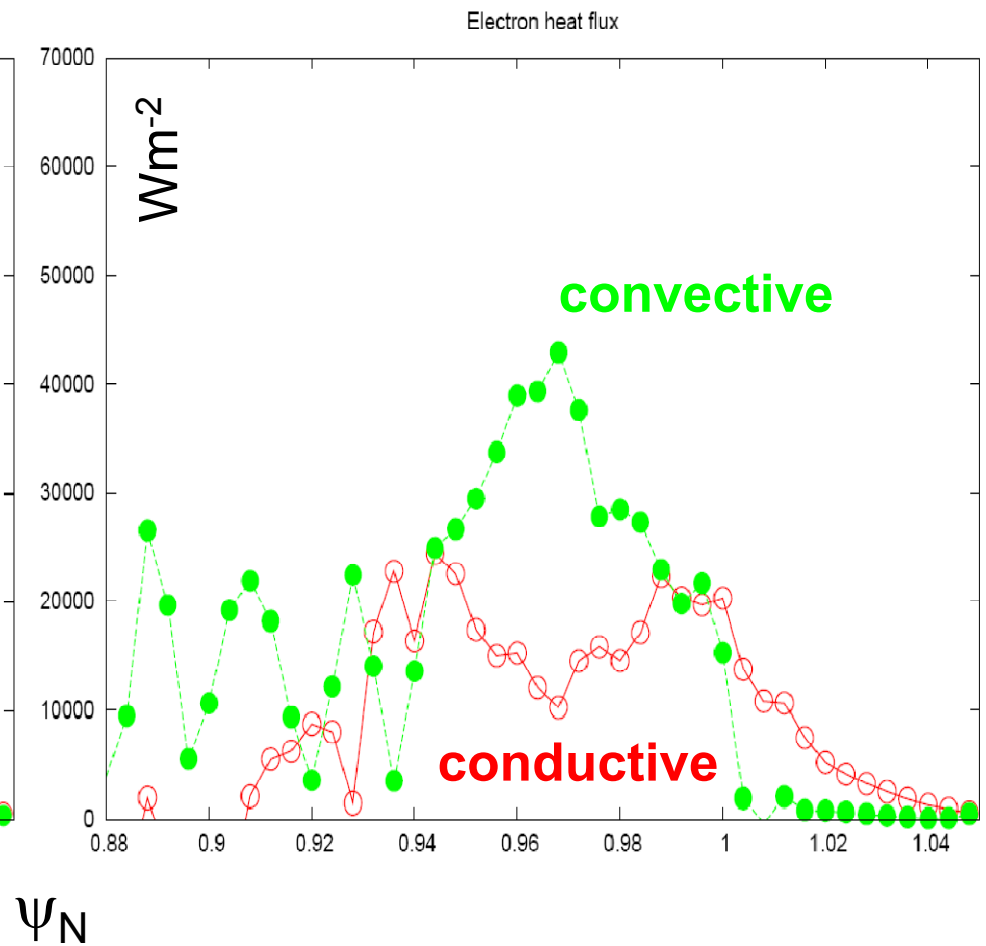
- T_e does not rise as much due to RMP enhanced heat transport.

Conductive radial electron heat loss ~ convective, inside the separatrix

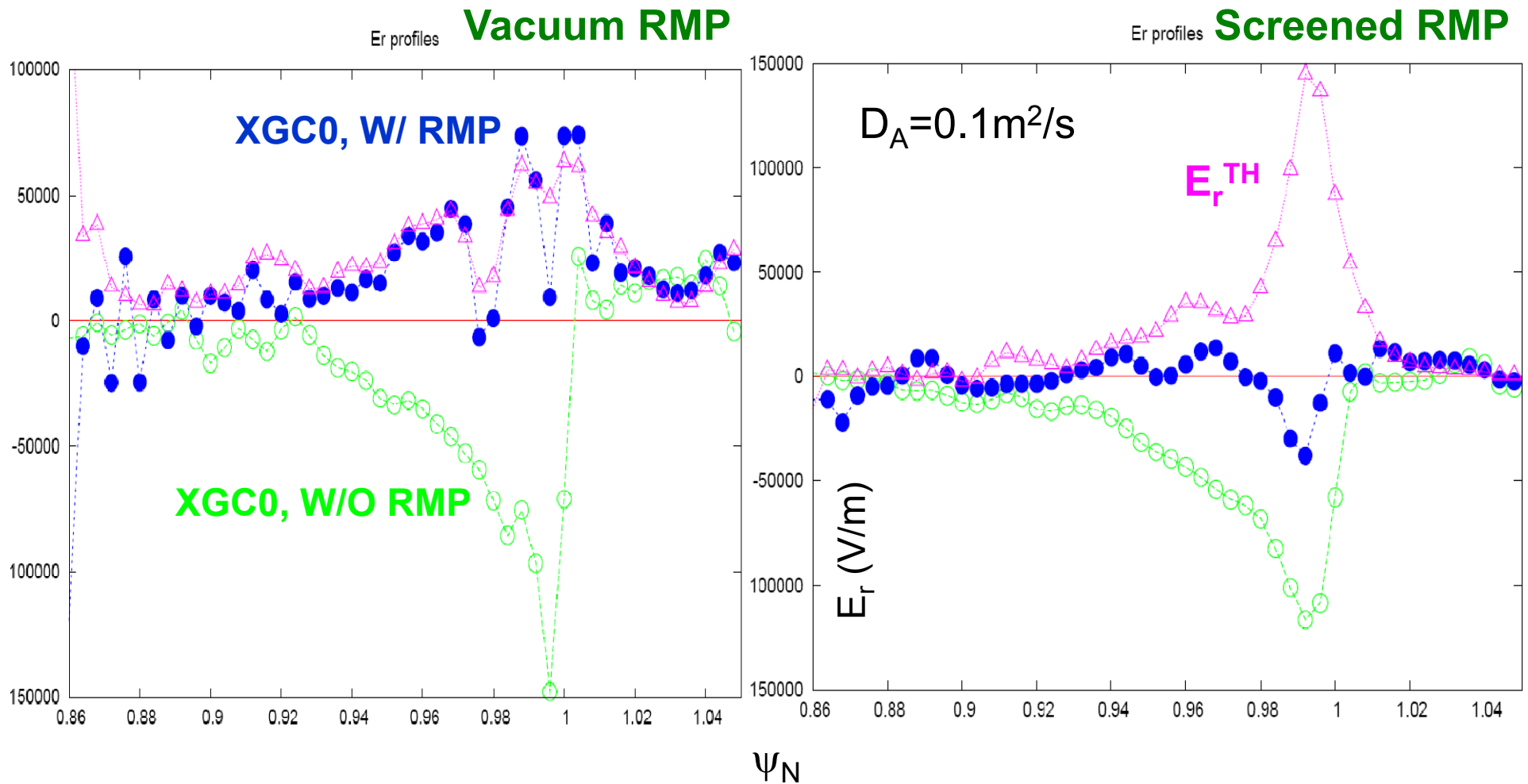
Vacuum RMP



Screened RMP

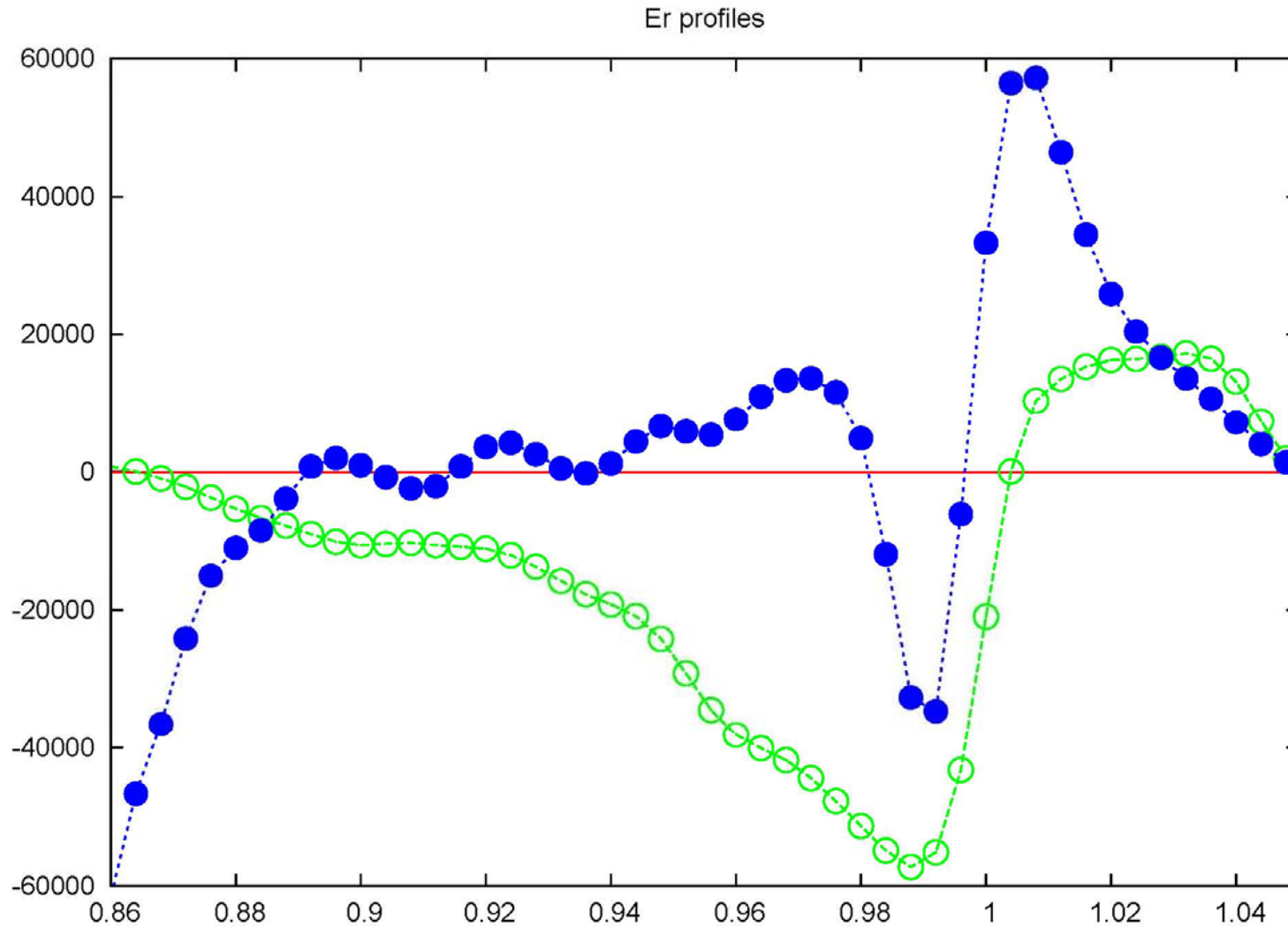


Narrow radial electric field well (from ion X-loss) can persist under screened RMPs



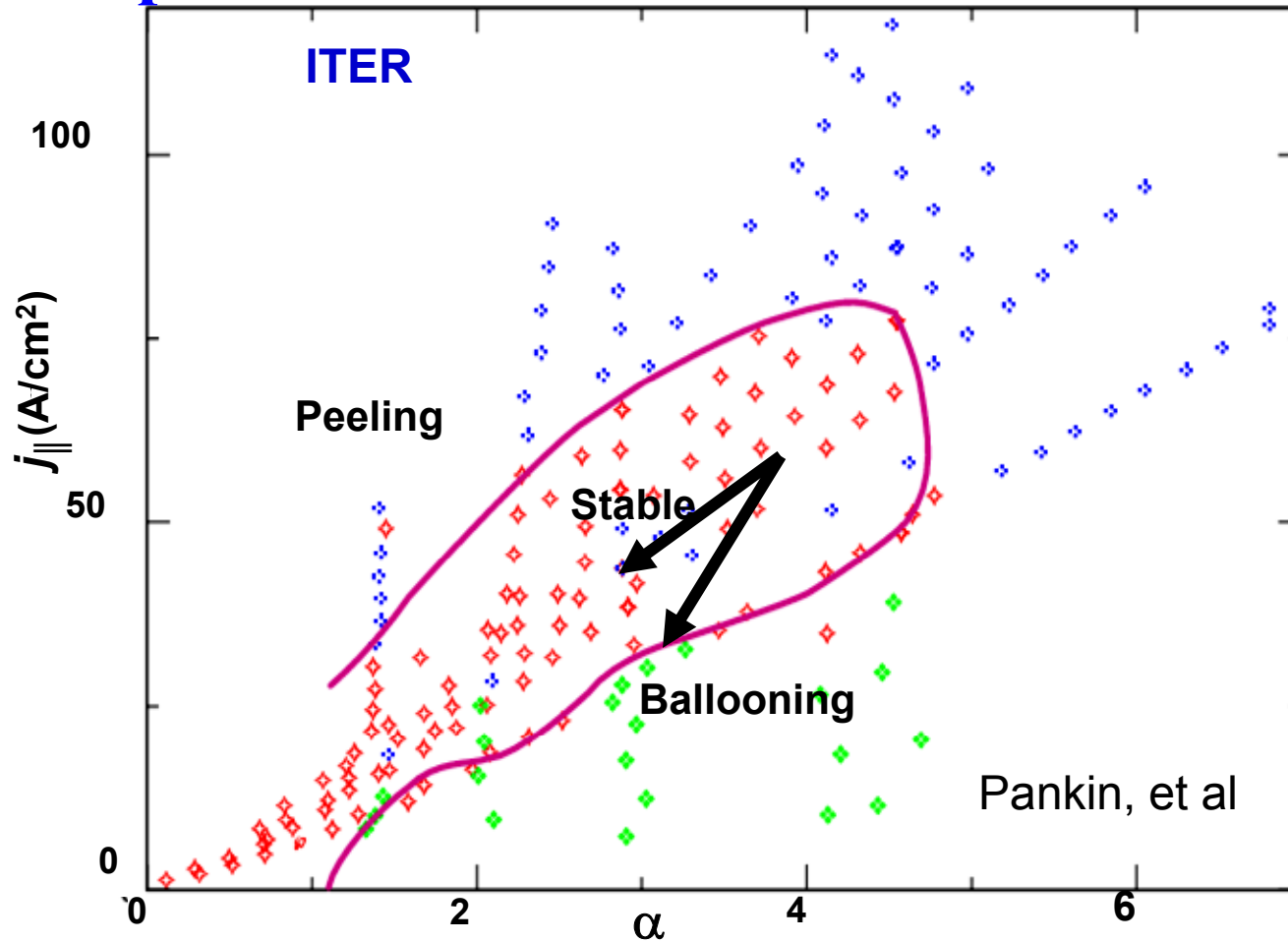
- Electron confining E_r from Einstein mobility relation is similar to XGC0 E_r in vacuum RMP but much different in screened RMP

When we increase D_A to $0.25\text{m}^2/\text{s}$, we get a much better agreement with experiment \rightarrow another prediction for D_A

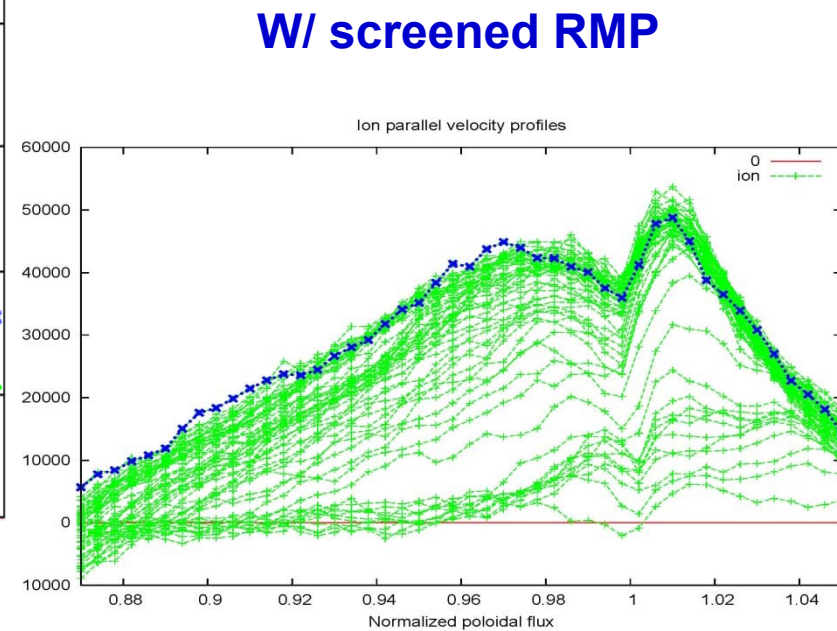
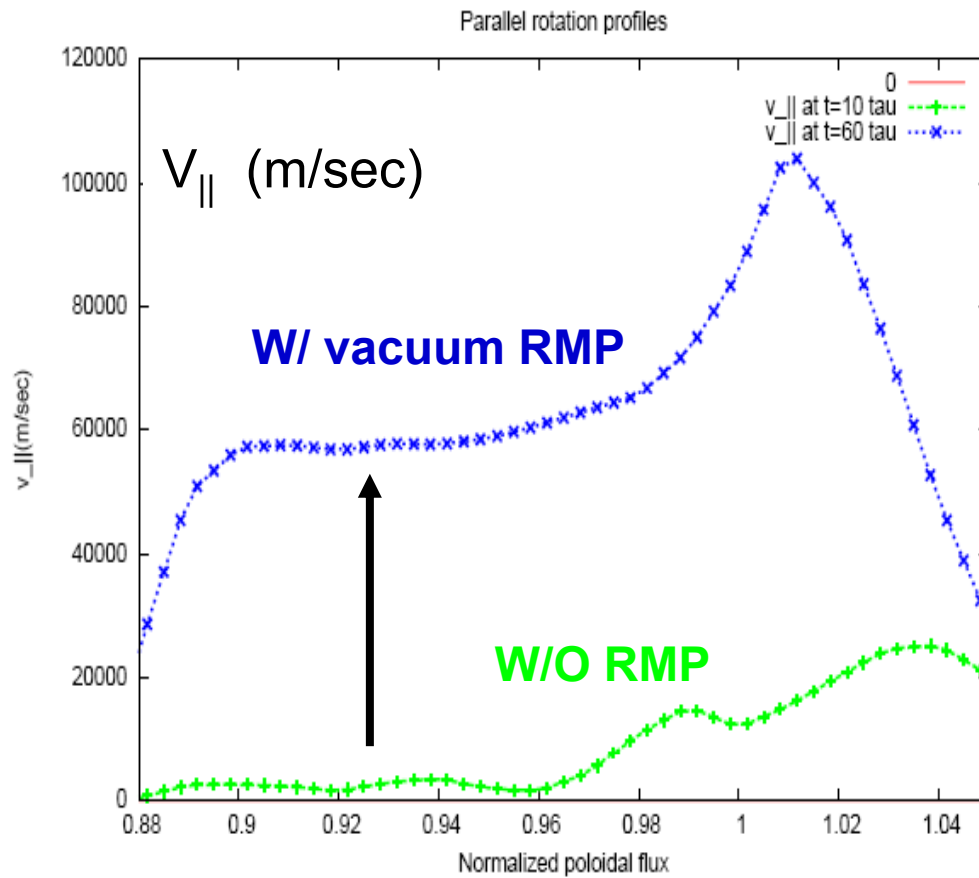


Different RMP effect on NSTX from DIII-D?

Self-consistent bootstrap current is needed for a better prediction of the RMP effect on ELM



Parallel rotation increases significantly by RMPs



Ψ_N

The observed phenomena on DIII-D have been reproduced.

Black part is from the vacuum RMP, predicted from XGC0 before experimental collaboration

- Density collapses (below $2 \times 10^{19} \text{ m}^{-3}$), but not the temperatures
- T_i increases significantly, T_e does not collapse.
- Toroidal rotation increases promptly by RMP

Red part is from screened RMP model, after experimental collaboration

- T_e steepens just inside the separatrix (screened RMP), with little change of ∇P there
- E_r well persists at narrow radial width just inside the separatrix (screened RMP): ion X-loss effect surfaces up at lower RMP

We think we have the basic understanding of the RMP transport.

We will improve the plasma screening effect from the XGC-M3D framework (NIMROD?).

At the same time, we can study the RMP concept improvement using the XGC-Elite framework.